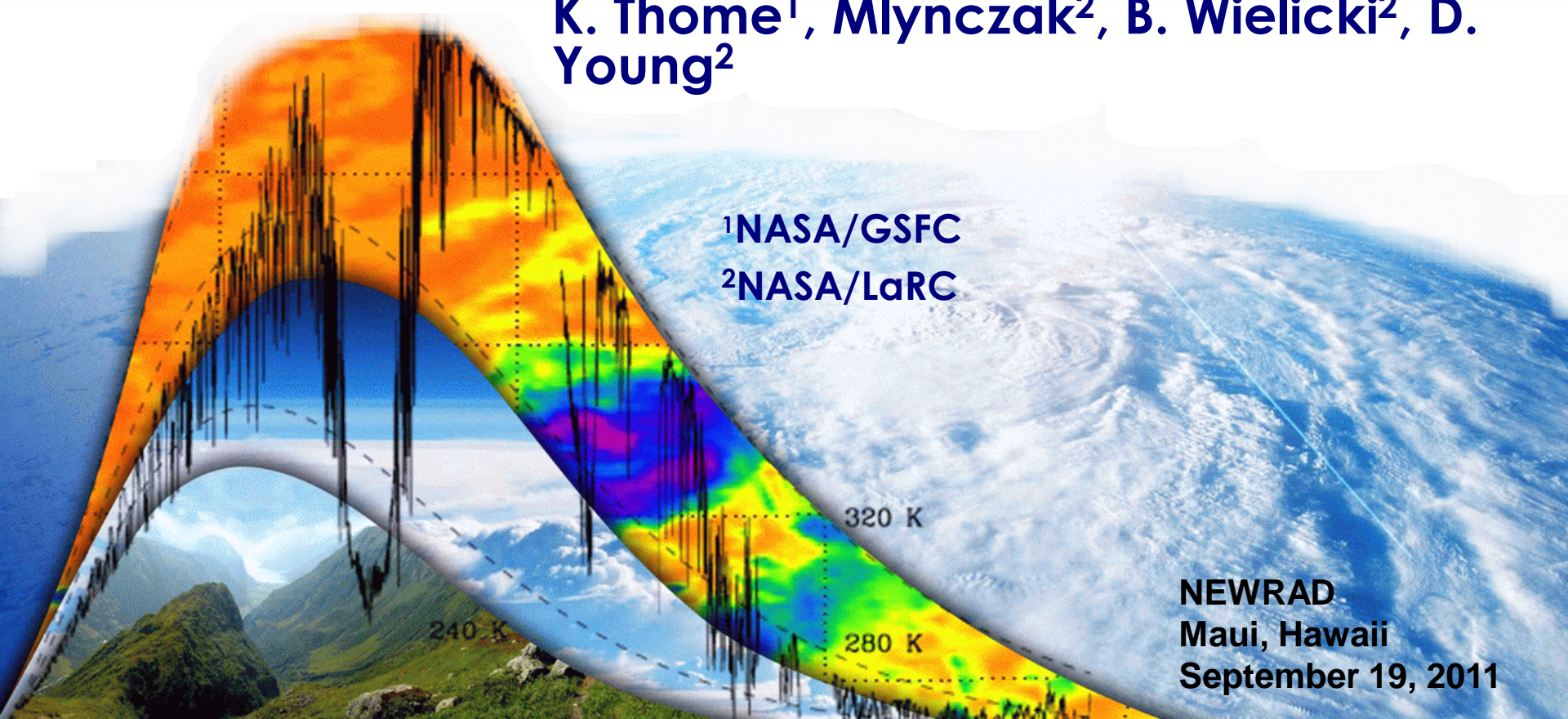


Calibration Accuracy for Climate-Quality Data Sets from Space Observations

K. Thome¹, Mlynczak², B. Wielicki², D. Young²

¹NASA/GSFC
²NASA/LaRC

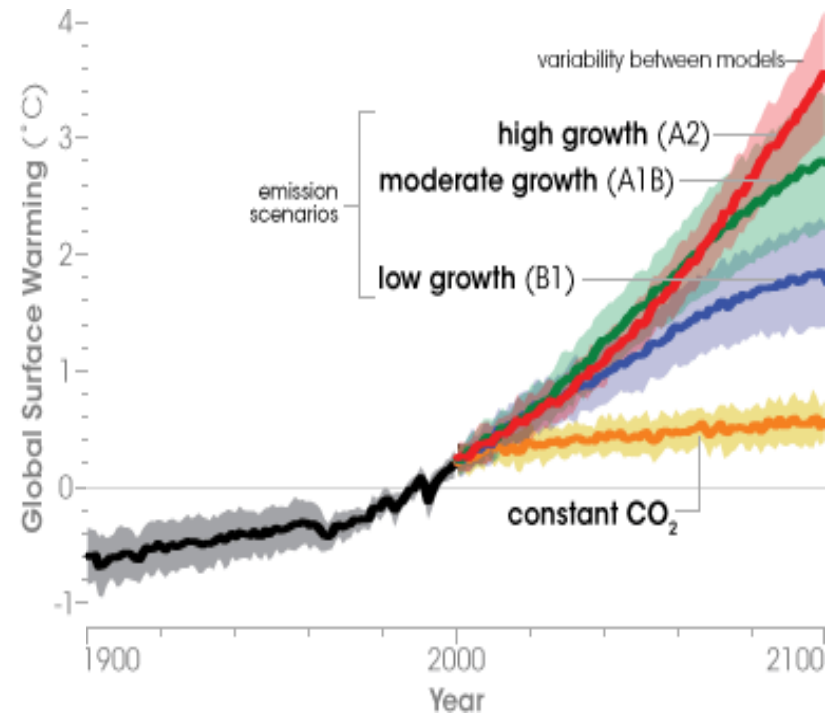
NEWRAD
Maui, Hawaii
September 19, 2011



Talk overview

Climate quality data records require high accuracy and SI traceability

- One goal is to understand climate change projections
- Summarize methods to determine the observing requirements
 - Reflected solar and IR
 - SI-traceable uncertainty
 - Temperature to 0.07 K ($k=2$)
 - Reflectance to 0.3% ($k=2$)
- Instrument approaches and calibration methods



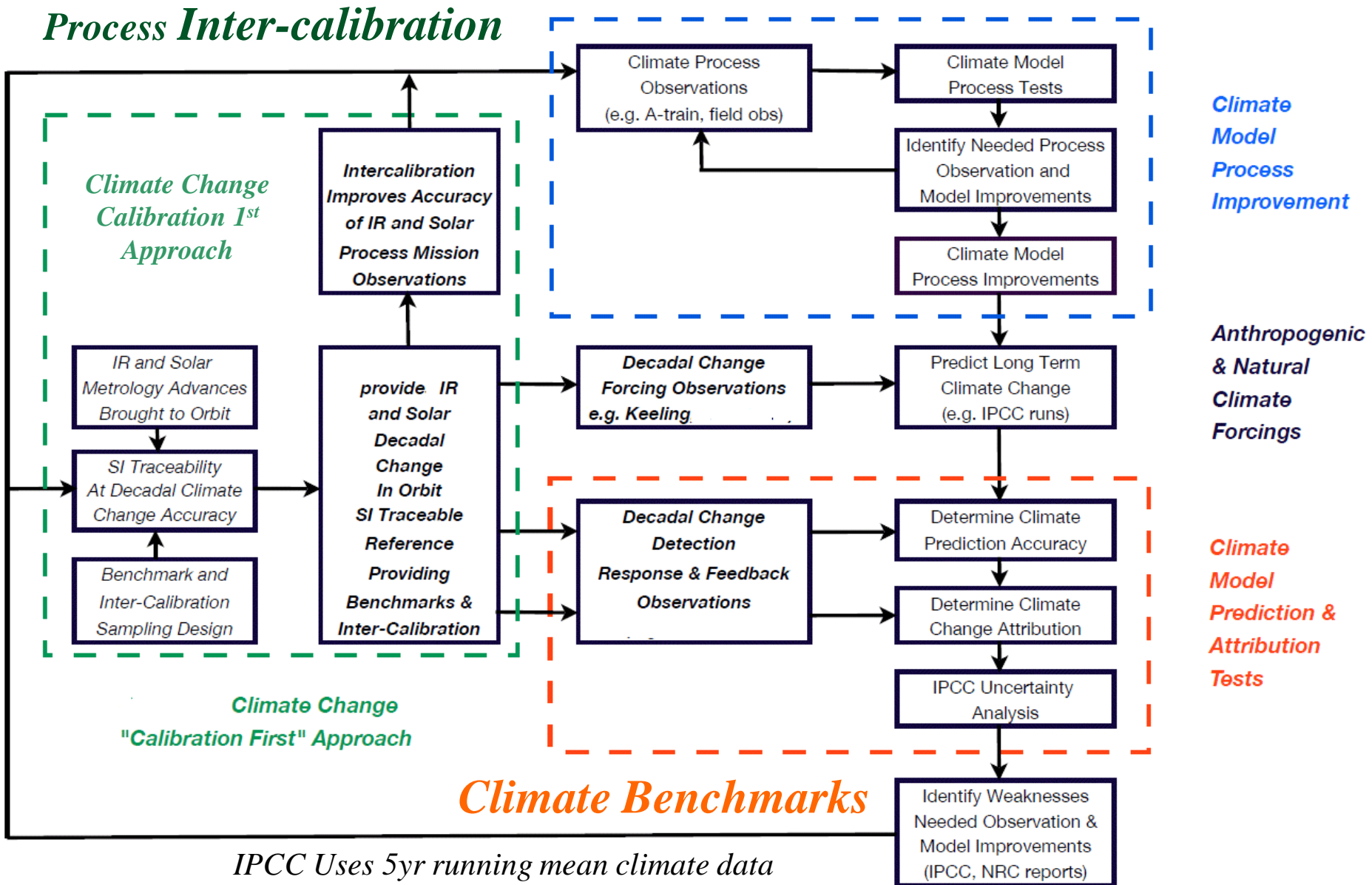
Climate model test and decadal change

Most powerful test of climate model prediction accuracy is decadal change observations

- Accuracy required at large time and space scales
 - Zonal annual, not instantaneous field of view
 - Very different than typical process missions
- Questions are
 - How long a data record is needed?
 - What variables are key?
 - What accuracy relative to perfect observing system is needed?
- How are the above determined?



Climate Science and Observations

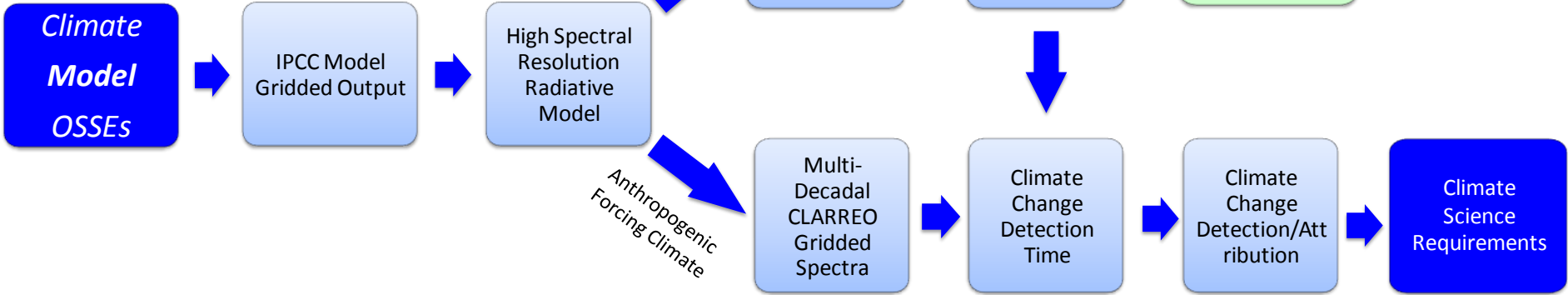


Climate science requirements

Observations and climate model predictions combined to develop requirements

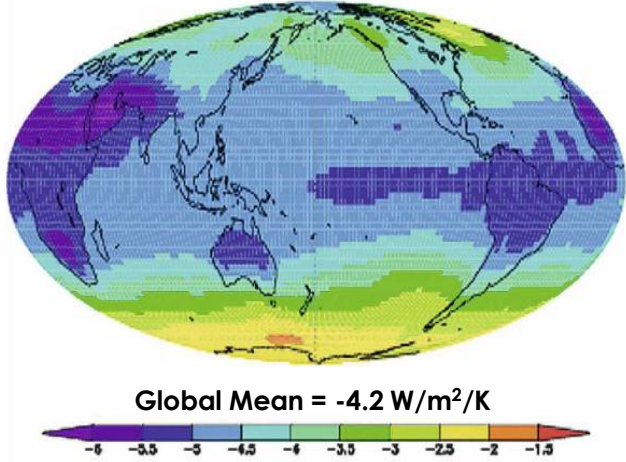


Observing System Simulation Experiments (OSSEs)

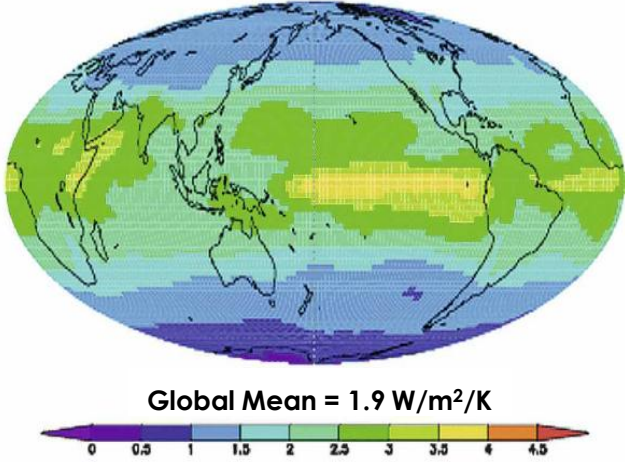


Feedbacks Drive Sampling Requirements

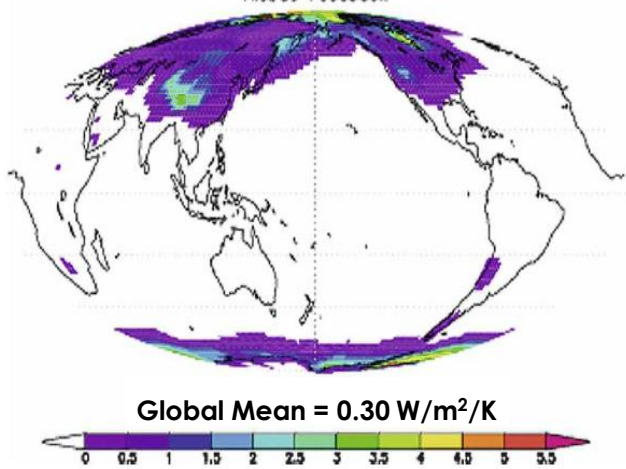
Temperature Feedback



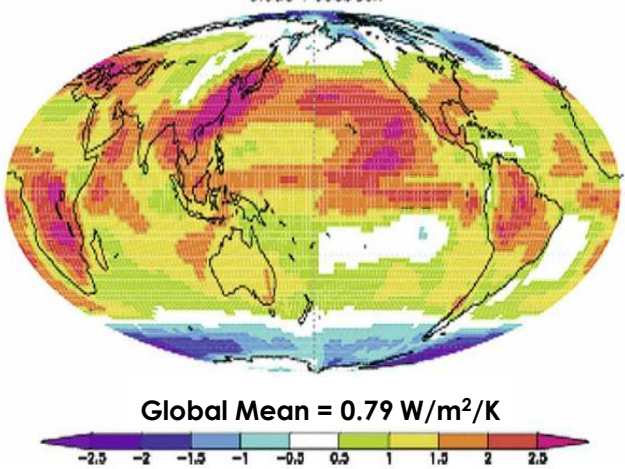
Water Vapor Feedback



Albedo Feedback



Cloud Feedback



Zonal annual means of land and ocean for lapse rate, water vapor, and surface albedo feedbacks

1000 km regional scale required for cloud feedbacks

Seasonal cycle required for reflected solar: cloud feedback, snow/ice albedo feedback

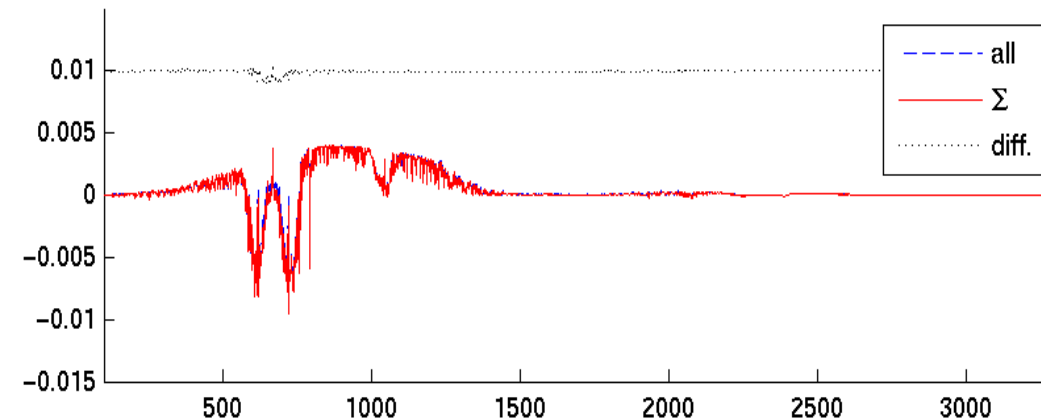
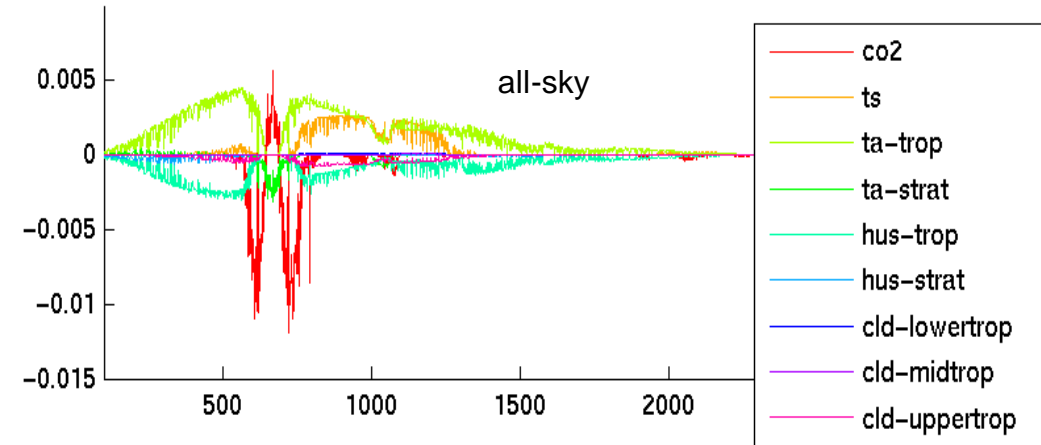
FIG. 8. Multimodel ensemble-mean maps of the temperature, water vapor, albedo, and cloud feedback computed using climate Multi-model ensemble-mean maps of the temperature, water vapor, albedo, and cloud feedback, computed using climate response patterns from the IPCC AR4 models and the GFDL radiative kernels (Soden et al., 2008)



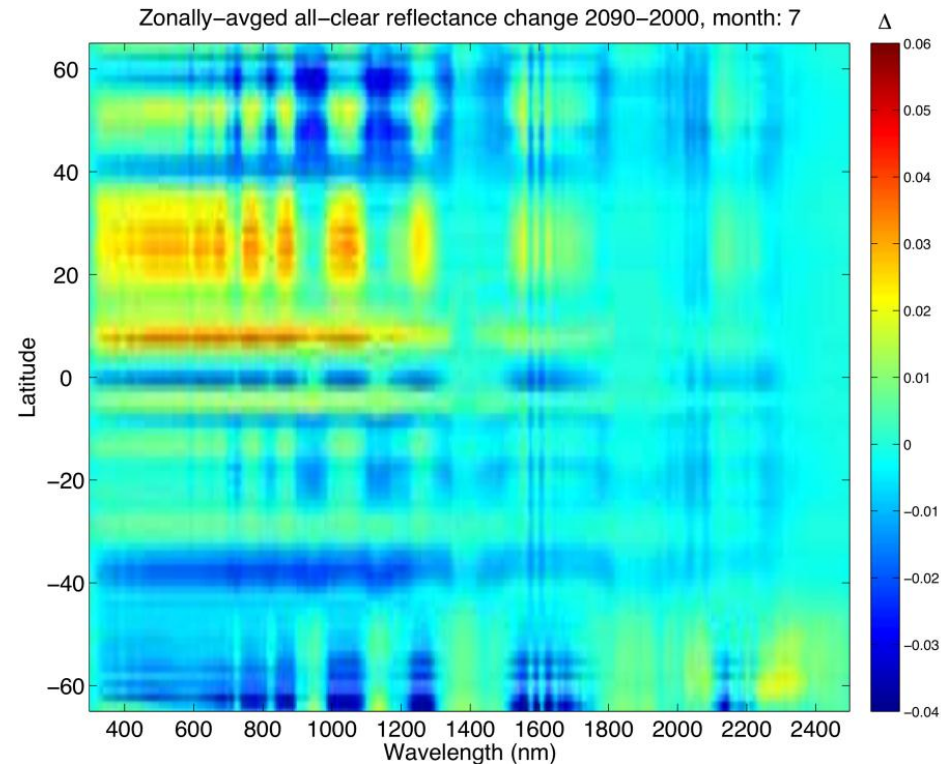
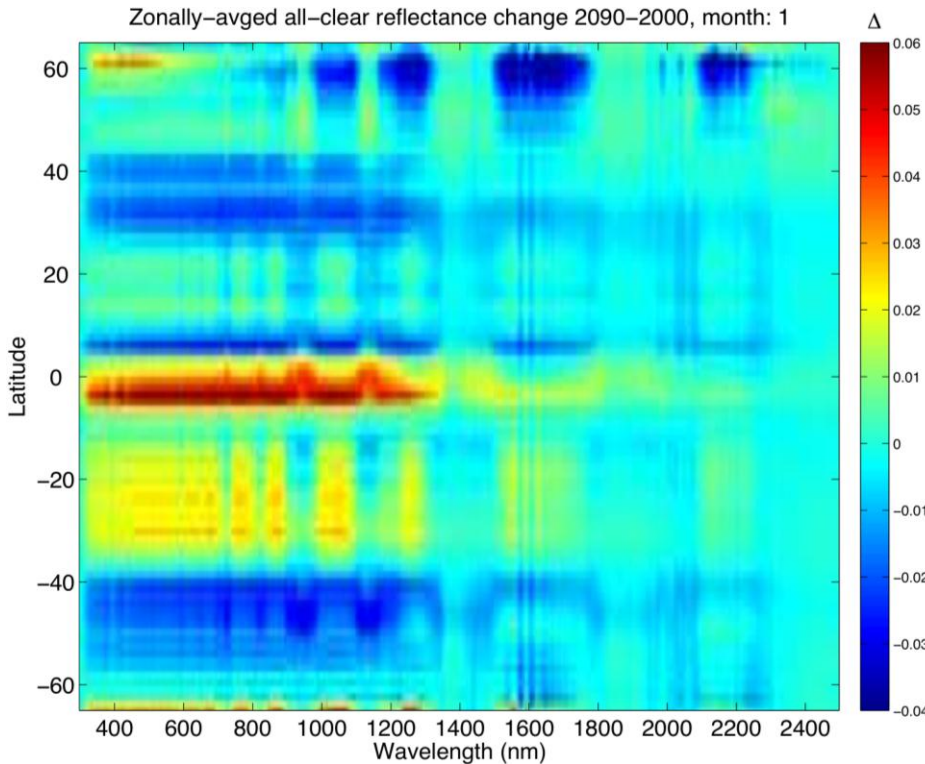
IR OSSEs

Studies by GFDL/ Harvard demonstrate the linearity of all-sky decadal change IR signals

- Doubled CO₂ change of all-sky global spectral radiance from $T(z)$, $q(z)$, CO₂, clouds, for the CFMIP CCCMA coupled climate model
- Examined linear sum of components versus total
- Eliminates the requirement for global clear-sky observations (Huang and Leroy, 2009)



RS OSSEs



Collins and Feldman, AGU Fall Meeting, 2010

Studies demonstrate linearity & information content of decadal change solar-reflected radiance signals

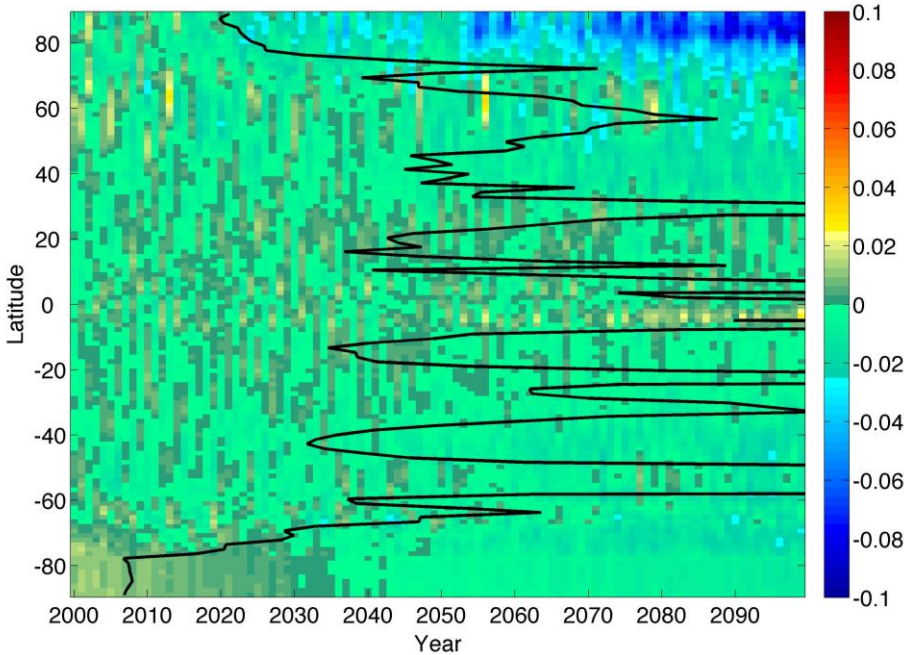
- U-Cal Berkeley, LASP, and LaRC



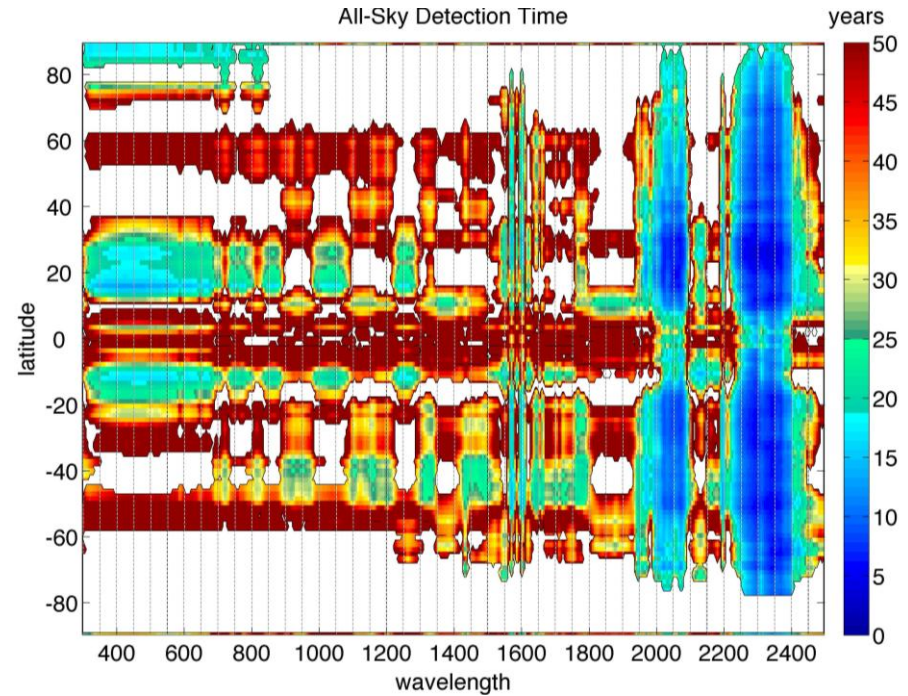
RS OSSEs

Δ All-sky spectral albedo evaluation shows time to detect (95% confidence)

Zonally-averaged annual mean all-sky albedo anomaly: experiment-control Δ



All-Sky Detection Time



Spectral measurements *shorten* time to detection over tropics and subtropics.

Collins et al, AGU Fall Meeting 2010



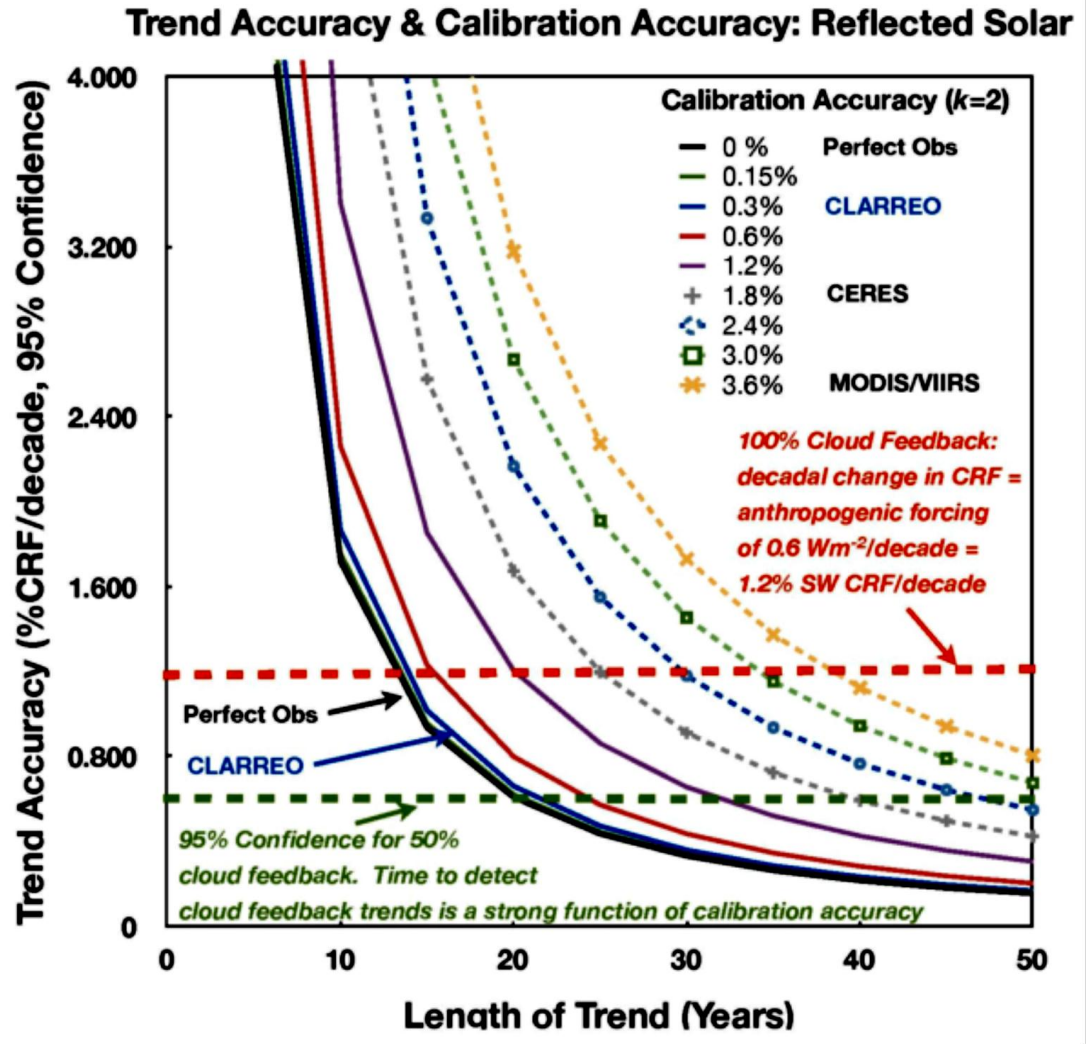
Accuracy Requirements

- Instrument absolute accuracy requirements are derived with the goal of achieving measurements
 - Within 20% of a perfect climate observing system
 - Time to detect trends within 15% of perfect observing system
- 0.1 K ($k=3$) for the IR spectra absolute accuracy required driven by natural variability of IR spectra.
- 0.3% ($k=2$) for the RS spectra (nadir reflectance) is required driven by natural variability of cloud radiative forcing, cloud fraction, cloud optical depth, particle size
- 0.03% ($k=2$) refractivity, consistent with accuracy of 0.1K ($k=3$) for temperature profile for 5km to 20km altitudes.
- SI traceable observations to survive short gaps in record



Accuracy requirements

Accuracy requirements can be inferred relative to natural variability & time to detect



- Traceable uncertainties from decadal change observation to SI standards, at decadal change accuracy levels
- Accuracy requirements driven by a time to detect above natural variability



Measurement approach

Benchmark climate data records provide the decadal change observations

- Benchmarks can be formed in two complementary ways
 - Well-calibrated, highly accurate sensor (e.g. TRUTHS, CLARREO) for direct data collection
 - Well-calibrated sensor calibrates operational sensors
- Dual approach provides test analogous to using independent measurements and analysis in metrology
- Accuracy of benchmark observations is required only at large time and space scales



Verification Systems Provide SI Traceability

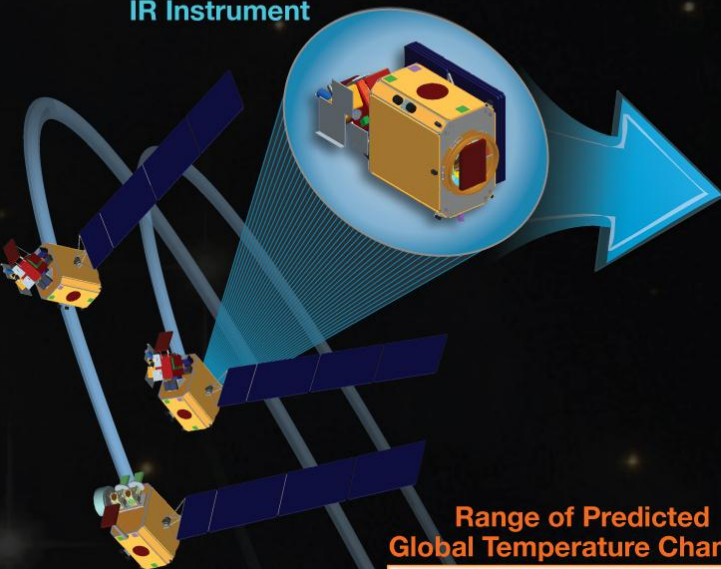
SI Traceability: Unbroken chain of comparison with stated uncertainties

- Partially met through traditional methods
 - Blackbody and deep space for radiance calibration
 - Lunar views
 - Temperature sensors to detect and correct for detector package changes
 - Preflight calibration
- On-orbit verification system
 - Quantify sources of bias
 - Polarization effects
 - Instrument line shape
 - Stray light
 - Verify accuracy through known sources

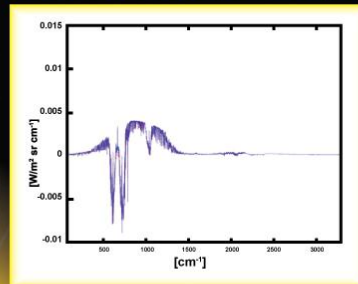


CLARREO: Calibrating Planet Earth

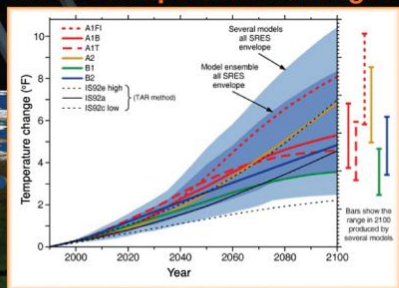
IR Instrument



CLARREO Anticipated Infrared Decadal Change



Range of Predicted Global Temperature Changes

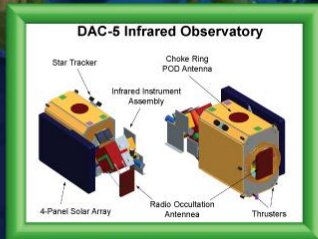


Informing Policy

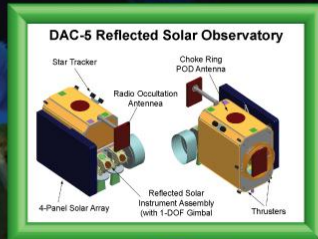


CLimate
Absolute
RadiancE &
REfractivity
Observatory

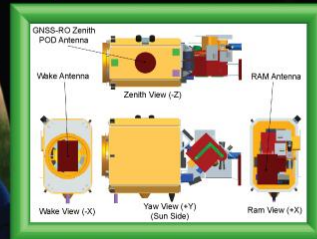
IR
Infrared



RS
Reflected Solar



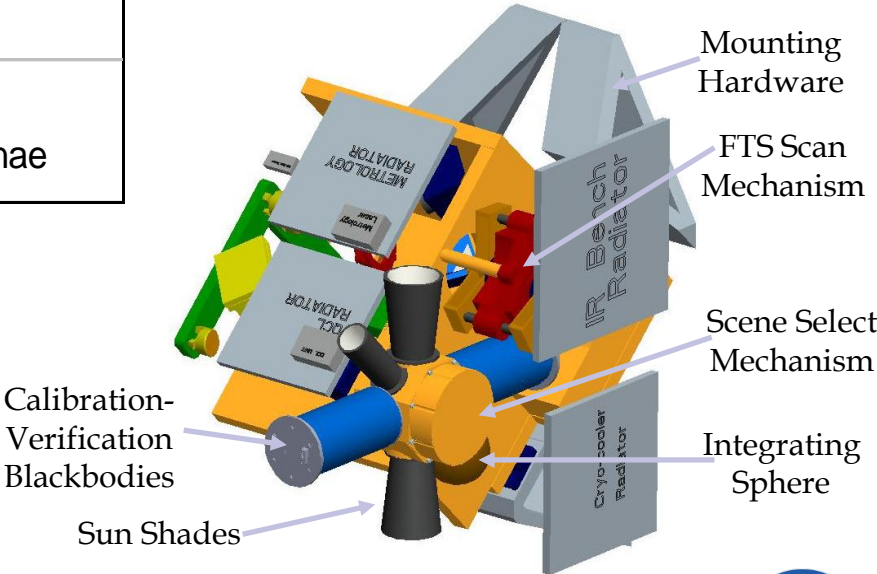
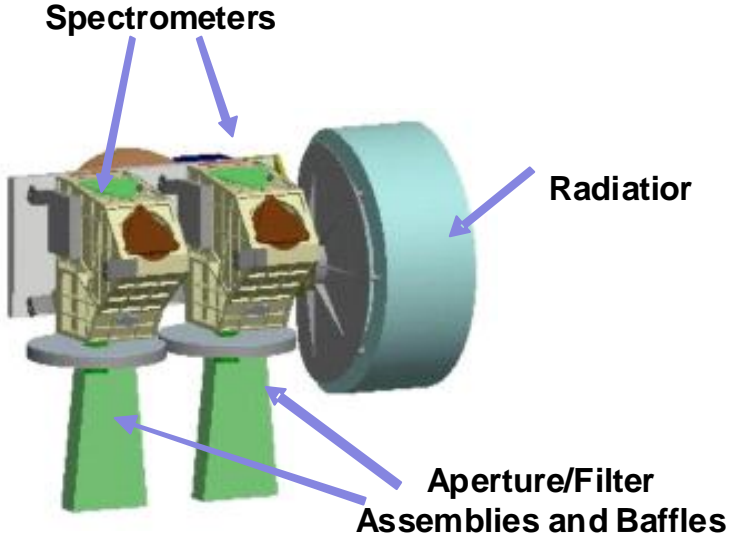
GNSS-RO
Global Navigation Satellite
System Radio Occultation



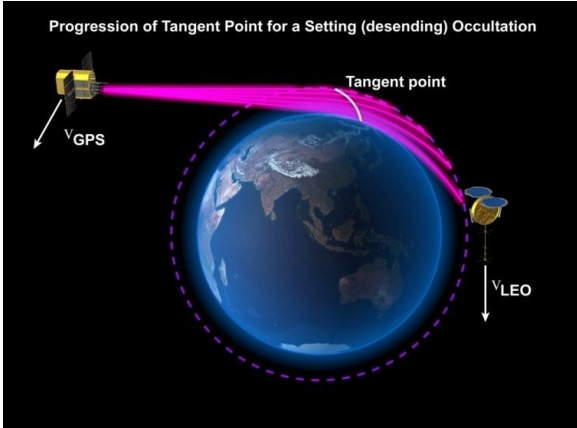
CLARREO Mission Overview

Payload Suite

Instrument	Features	
Infrared Spectrometer	Type	Fourier transform spectrometer
	Spectral Range	5 to 50 micron
	Configuration	Single combined instrument
Reflected Solar Spectrometer	Type	Grating spectrometer
	Spectral Range	320-2300 nanometer
	Configuration	Two box design
GNSS Radio Occultation System	Signal Range	GPS and Galileo
	Configuration	Receiver Two occultation antennae



IR Instrument Concept Design (BB Radiator and Enclosure Partially Removed)



CLARREO Project update

It has been an up and down ten months for CLARREO

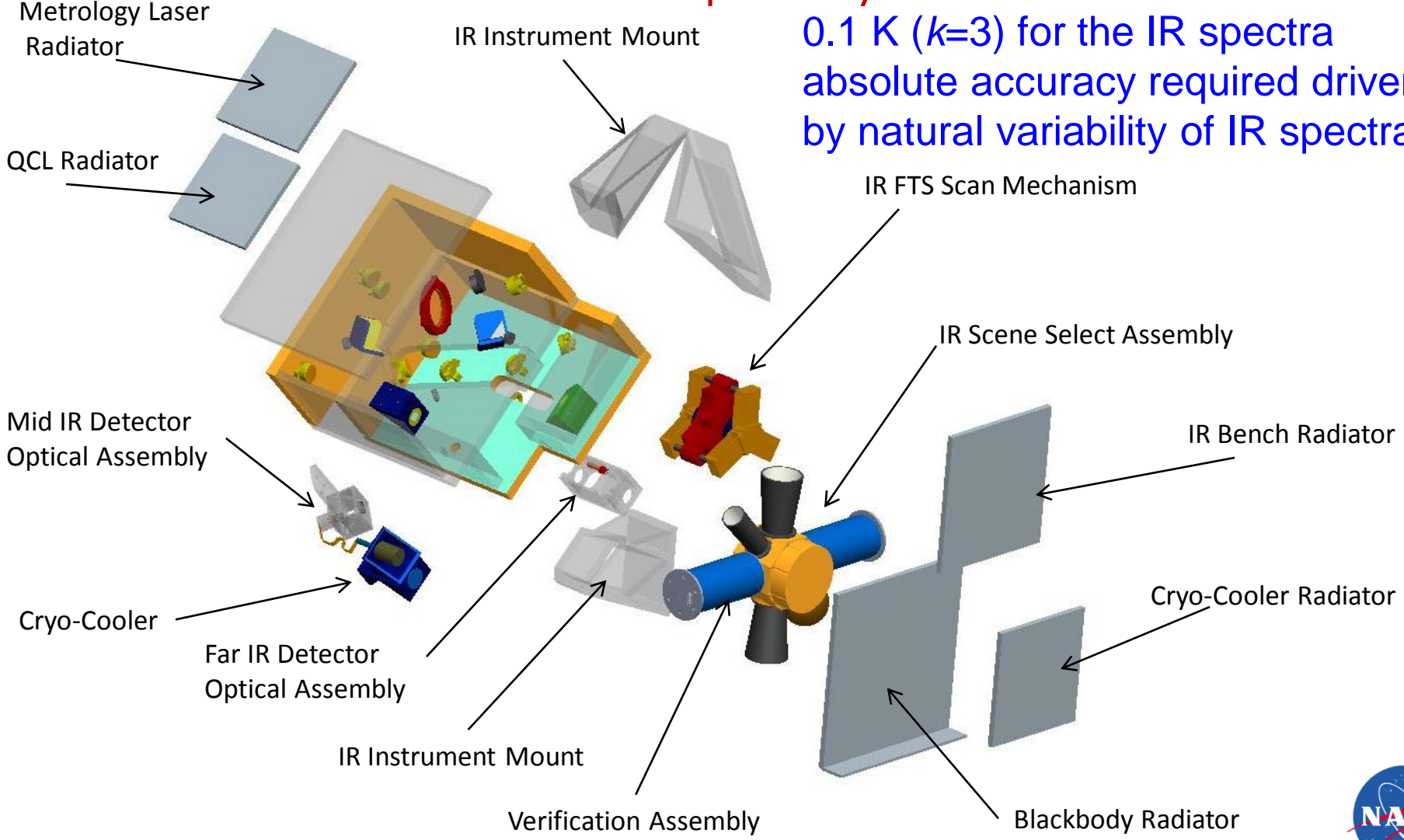
- Demonstrated readiness to begin Phase A at a successful Mission Concept Review in November 2010
- Science Definition Team selected January 2011
- Unfortunately, FY2012 NASA budget removed \$1.2 billion from proposed NASA Climate Initiative in FY2012 - FY2015
 - NASA was directed not to proceed to Phase A for CLARREO in FY2011
 - CLARREO was placed in “Extended Pre-Phase A” from FY2012-FY2016
- Funding will continue to allow work to concentrate on
 - Advancing CLARREO science, including continuation of Science Definition Team
 - Exploring cost-effective alternatives for achieving some portion of the CLARREO science objectives
 - Reducing risk for achieving on-orbit SI-traceability
 - Continuing engagement with the user community (GSICs, CEOS) and potential international partners



IR Instrument Suite

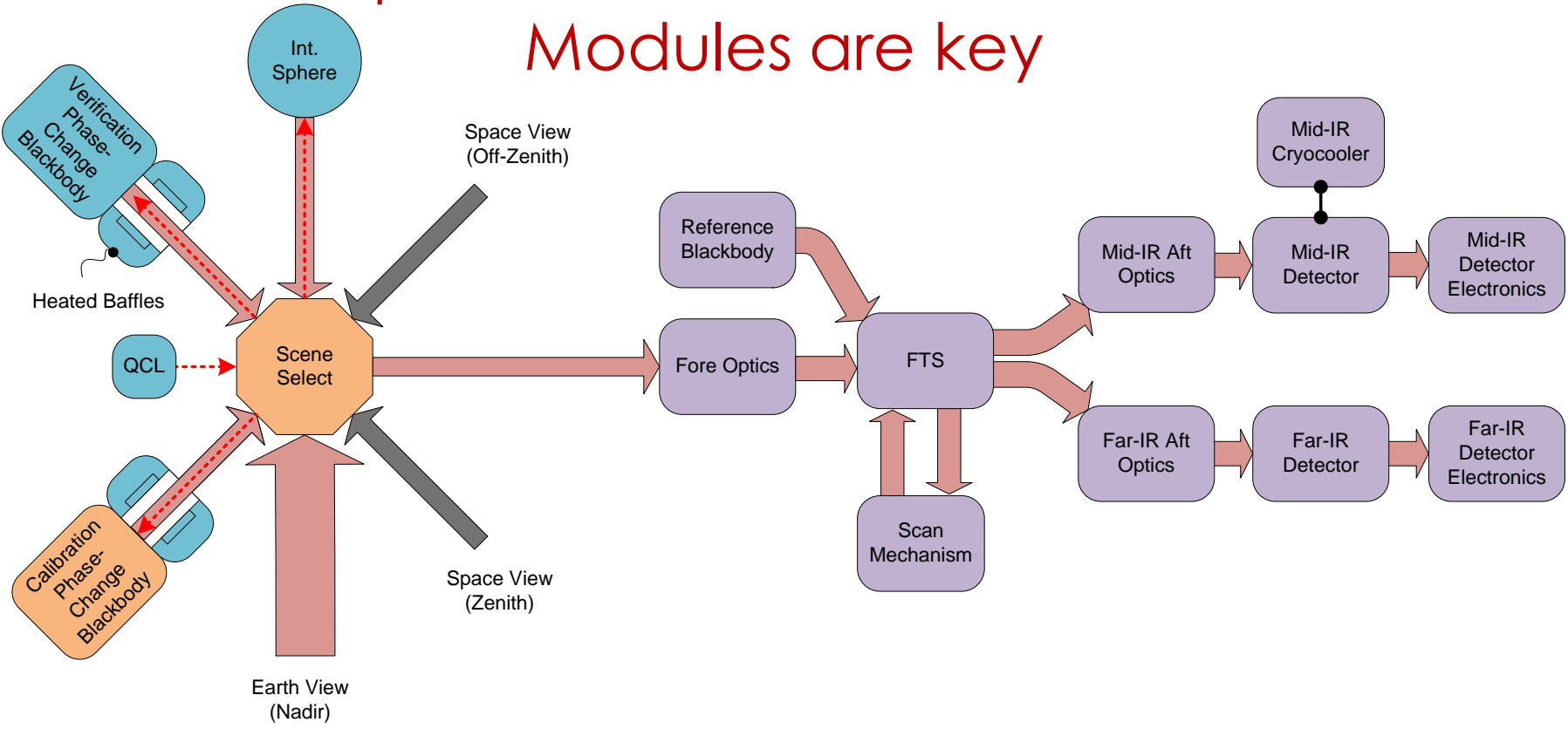
Instrument suite is of moderate size and complexity

0.1 K ($k=3$) for the IR spectra absolute accuracy required driven by natural variability of IR spectra.



IR Concept

Electro-Optical and Calibration-Verification Modules are key



IR Error budget

Estimated $k=3$ uncertainties at 1000 cm^{-1} for scene temperature of 250K , with calibration BB at 270K

Total Combined Uncertainty
54 mK

IR Level 1 Requirement
 100 mK , 3σ ($k=3$)

Combined Type B Uncertainty
54 mK

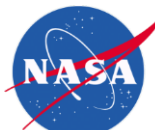
Annual Type A Uncertainty
< 1 mK

Calibration Blackbody Radiance
31 mK

Space View Radiance
< 1 mK

Gain Nonlinearity
29 mK

FTS Uncertainty Terms
33 mK



On-Orbit Calibration of Verification Blackbody

- Referenced to International Temperature Scale (ITS) through pure materials (water, gallium, mercury)
- Thermistors in blackbody continuously calibrated providing fundamental tie to SI through Kelvin

$$L_{BB}(\nu, T_{BB}) = \epsilon_{BB}(\nu) * \frac{2h\nu^3}{c^2} * \frac{1}{(e^{(h\nu/kT_{BB})} - 1)}$$

Planck Equation

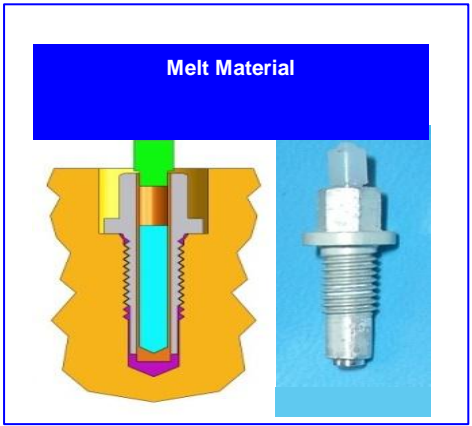
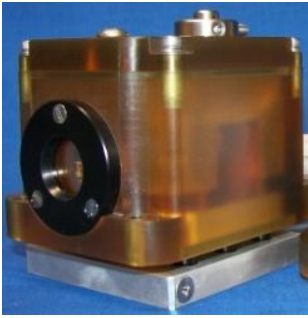
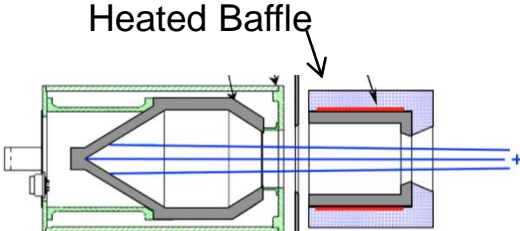
Emissivity
Temperature

Cavity Emissivity Measurement

3 Phase Change Cells Provide SI Traceable Fixed Points (-40°C, 0°C, 30°C)

Quantum Cascade Laser (QCL)

Phase Change Cells



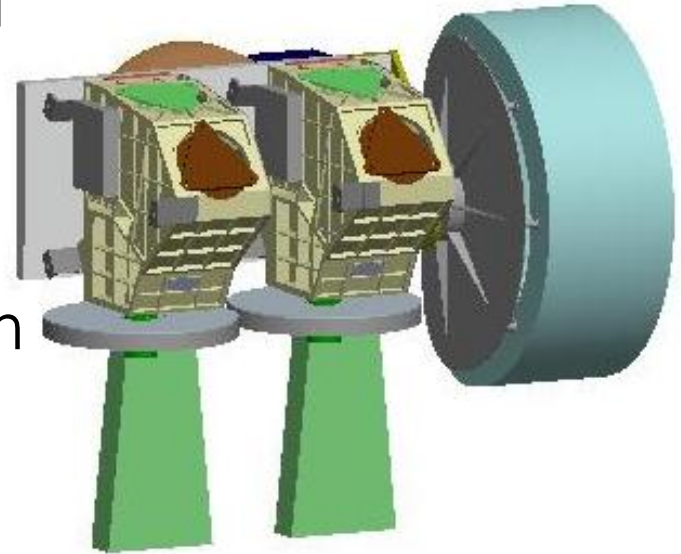
SI (Kelvin)-Based IR Radiance Scale Realization



RS instrument

Reflectance must be traceable to SI standards
at an absolute uncertainty $<0.3\%$

- Reflectance obtained from ratio of radiance while viewing earth's surface to measurements of irradiance while viewing the sun
- Spectral range from 320 to 2300 nm
- 500-m GIFOV with 100-km swath
- Goal of sensor design is to reduce complexity for accurate calibration
- Spectrometer approach with two separate entrance apertures
- Commonality of design of two boxes aids in calibration



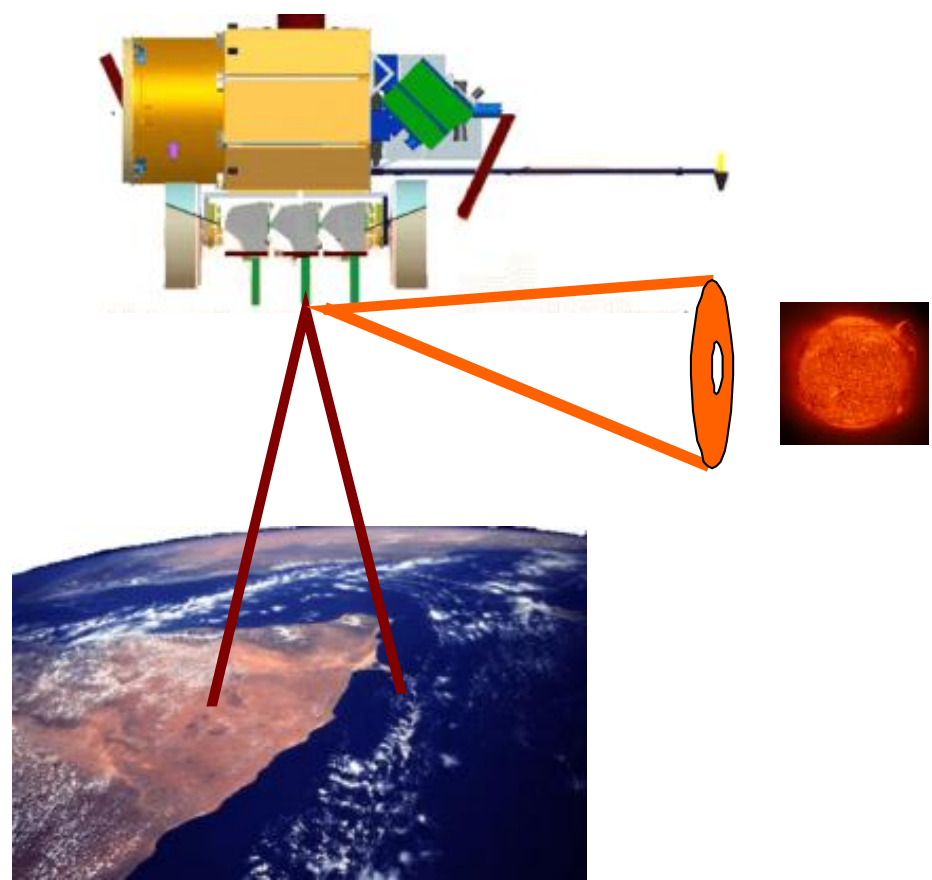
Reflectance approach

- Retrieve reflectance via ratio of earth-view data to solar-view data
 - Single detector scans entire solar disk
 - Response of i^{th} detector is

$$R_{i,\lambda}^{\text{sensor}} = \frac{\sum_{x_{\text{solar}}} \sum_{y_{\text{solar}}} S_{i,\lambda}^{\text{solar}}(x'_{\text{solar}}, y'_{\text{solar}})}{(T_{\text{attenuator}} A_{\text{attenuator}}) E_{\text{solar}}}$$

- Bidirectional reflectance distribution function (BRDF) is

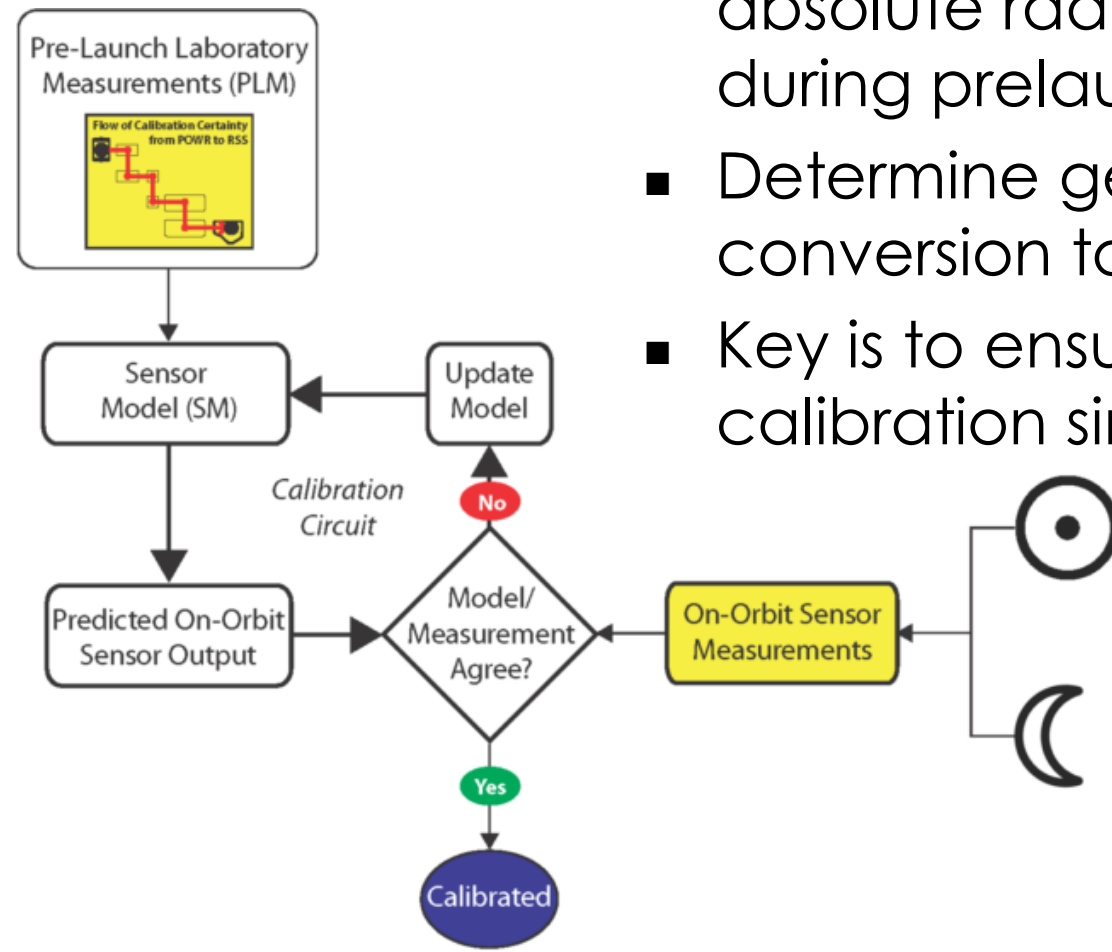
$$BRDF_{i,\lambda}^{\text{earth}} = \frac{L_{i,\lambda}^{\text{earth}}}{E_{\text{sun}} \cos \theta_{\text{solar}}} = \frac{S_{i,\lambda}^{\text{earth}}}{R_{i,\lambda}^{\text{sensor}} A_{\text{sensor}} \Omega_{\text{sensor}}} \frac{(T_{\text{attenuator}} A_{\text{attenuator}}) R_{i,\lambda}^{\text{sensor}}}{\cos \theta_{\text{solar}} \sum_{x_{\text{solar}}} \sum_{y_{\text{solar}}} S_{i,\lambda}^{\text{solar}}(x'_{\text{solar}}, y'_{\text{solar}})}$$



Calibration approach

Successful transfer to orbit through accurate prediction of sensor behavior

- Characterize sensor to SI-traceable, absolute radiometric quantities during prelaunch calibration
- Determine geometric factors for conversion to reflectance
- Key is to ensure prelaunch calibration simulates on-orbit sources



Key error terms

Developed a preliminary error budget based on a nominal design for the RS sensor

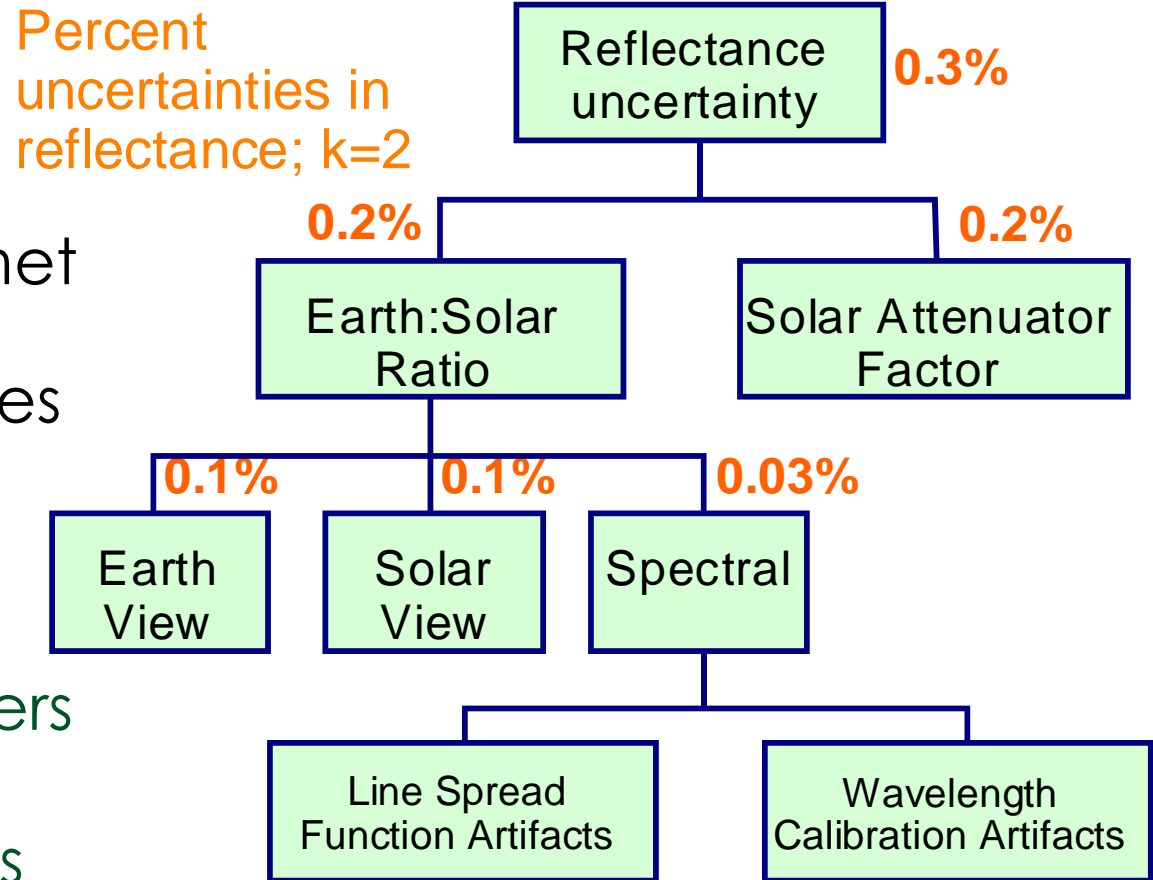
- Key uncertainties are
 - Geometry differences between the solar and earth views
 - Knowledge of attenuator behavior when viewing the sun
 - Sensor behavior
 - ◆ Detector linearity
 - ◆ Noise behavior
 - Polarization

$$BRDF_{i,\lambda}^{earth} = \frac{S_{i,\lambda}^{earth}}{R_{i,\lambda}^{sensor} A_{sensor} \Omega_{sensor}} \frac{(T_{attenuator} A_{attenuator}) \langle R_{\lambda}^{sensor} \rangle}{\cos \theta_{solar} \sum_k \sum_l S_{k,l}^{solar} r_{k,\lambda}^{flat\ field}} \frac{a_{sensor}^{straylight} \omega_{sensor}^{straylight} a_{attenuator}^{straylight}}{r_{i,\lambda}^{flat\ field} r_{i,\lambda}^{nonlinearity} r_{i,\lambda}^{polarization}}$$



SIRCUS-based Error Budget

- Radiometric calibration requirements of RS instrument can be met with currently-available approaches
- Requires inclusion of NIST-based methods
 - Detector-based transfer radiometers
 - Narrow-band SIRCUS approaches
 - Hyperspectral image projector-based scene projections



Summary

Climate-quality data records are achievable with current technology

- Climate OSSEs are invaluable in the definition of mission requirements
- Give confidence that the measurement approach would meet the goals of the Decadal Survey
- Climate model OSSEs should be used to provide more rigorous climate observing system requirements where possible including perturbed physics ensembles
- Lessons learned from CLARREO mission definition are applicable to other missions
- A goal of CLARREO pre-Phase A efforts is to ensure that these lessons are not lost



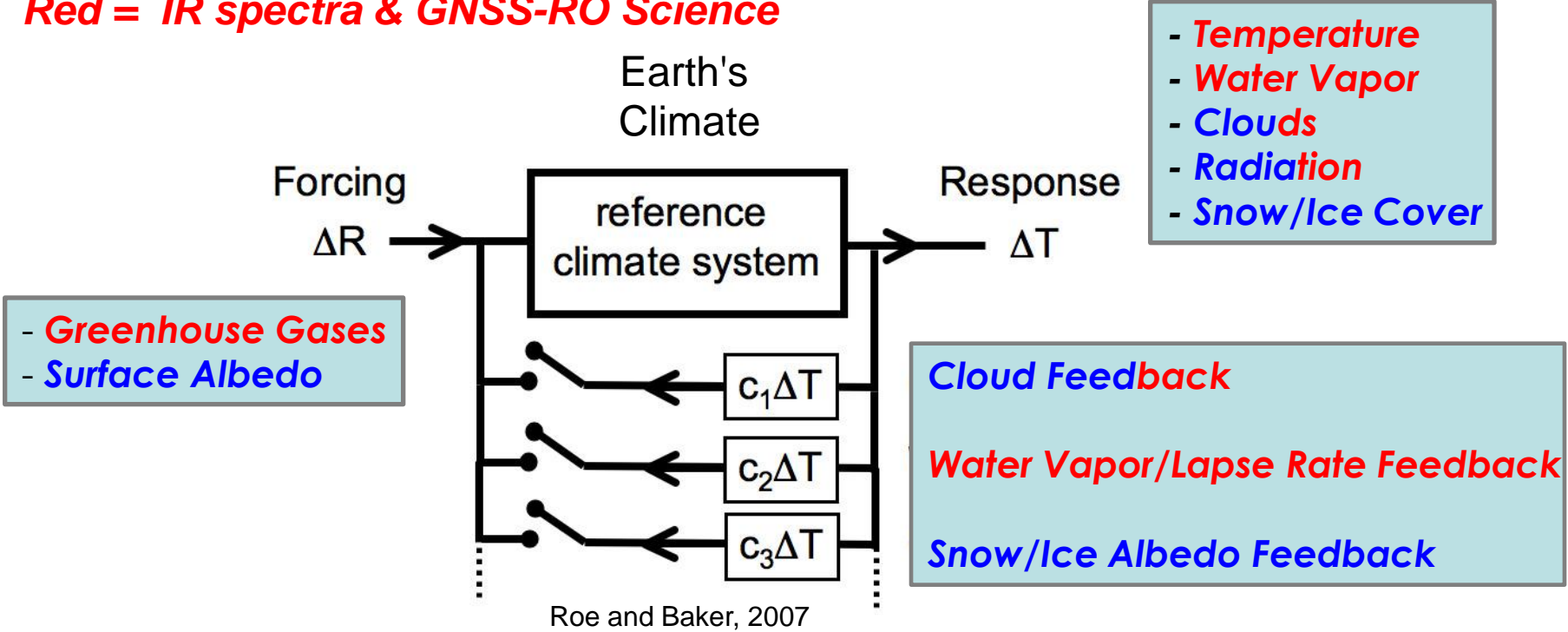
Backup Slides



Forcing, Response, Feedback

100% of science value is in the data accuracy

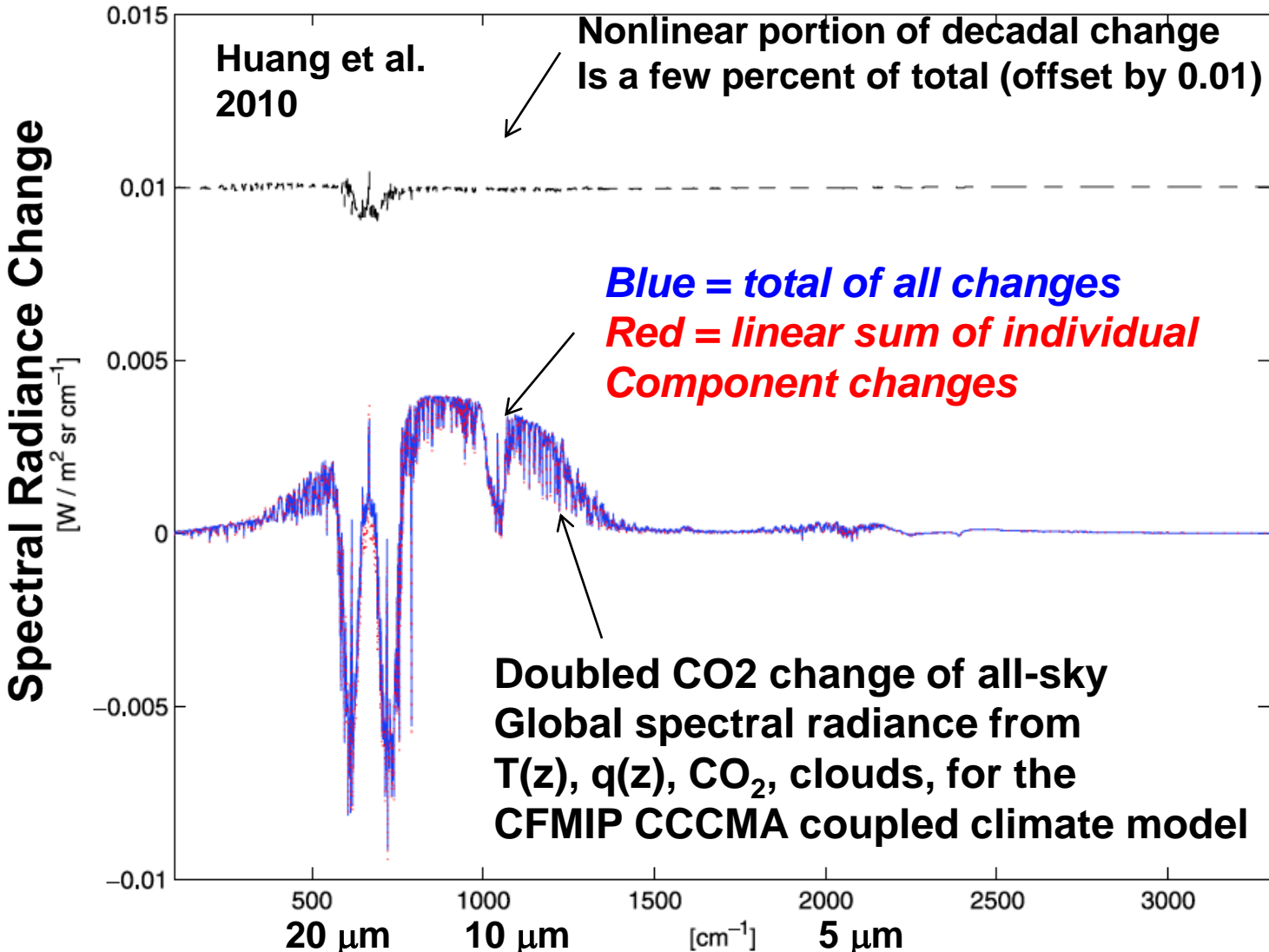
Blue = Solar Reflected Spectra Science
Red = IR spectra & GNSS-RO Science



50% of Science Value is in Reflected Solar Spectra
50% of Science Value is in Infrared Spectra & GNSS-RO



Spectral Decadal Change is Linear



Instantaneous changes are nonlinear: decadal change is highly linear

Similar linearity found for decadal change of reflected solar spectra Jin et al., 2010

