

On-orbit Absolute Blackbody Emissivity Determination Using the Heated Halo Method

Jonathan Gero, Joe Taylor, Fred Best, Ray Garcia, Hank Revercomb, Bob Knuteson, Dave Tobin, Doug Adler, Nick Ciganovich and Steve Dutcher

University of Wisconsin—Madison

Space Science and Engineering Center

NEWRAD 2011

Maui, Hawaii

September 19, 2011



Overview

- Introduction
 - Traceability of infrared blackbody radiance
 - Paint degradation in space
 - On-orbit Absolute Radiance Standard
- Heated Halo Emissivity Monitor
 - Test configuration
 - Emissivity results and uncertainty
 - Comparison between S-HIS, ARI and NIST measurements
- Summary

Traceable Blackbody Radiance

- Planck function:
$$B_{\tilde{\nu}}(T) = \frac{2hc^2\nu^3}{\exp(h\nu c / k_B T) - 1}$$
- Blackbody radiance:
$$I_{\tilde{\nu}, Blackbody}(\epsilon_{\tilde{\nu}}, T) = \epsilon_{\tilde{\nu}} \cdot B(T)$$

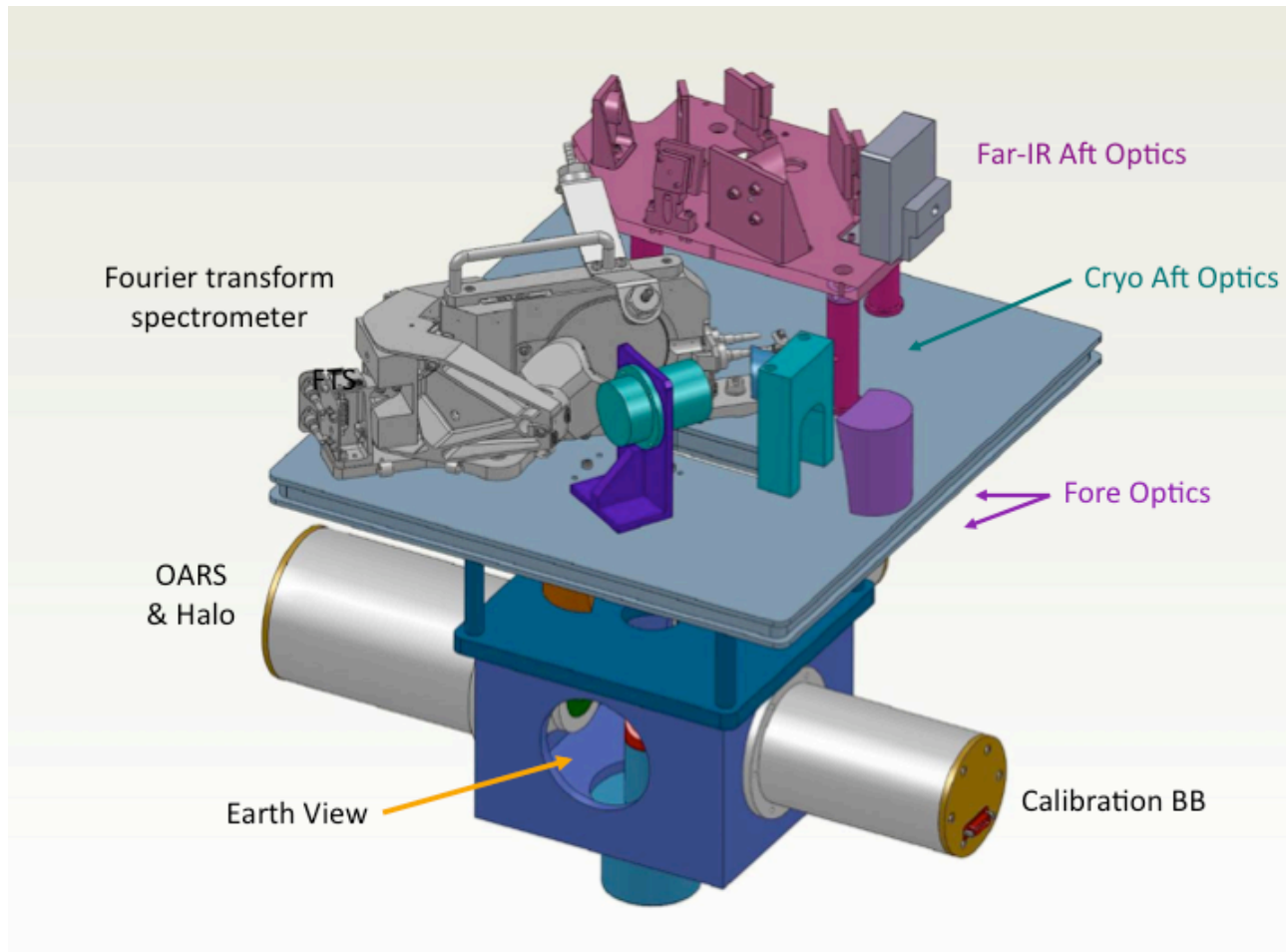
Traceable Blackbody Radiance

- Planck function:
$$B_{\tilde{\nu}}(T) = \frac{2hc^2\nu^3}{\exp(h\nu c / k_B T) - 1}$$
- Blackbody radiance:
$$I_{\tilde{\nu}, Blackbody}(\epsilon_{\tilde{\nu}}, T) = \epsilon_{\tilde{\nu}} \cdot B(T)$$
- Both temperature and emissivity of a blackbody must be known — on-orbit — throughout the lifetime of the instrument

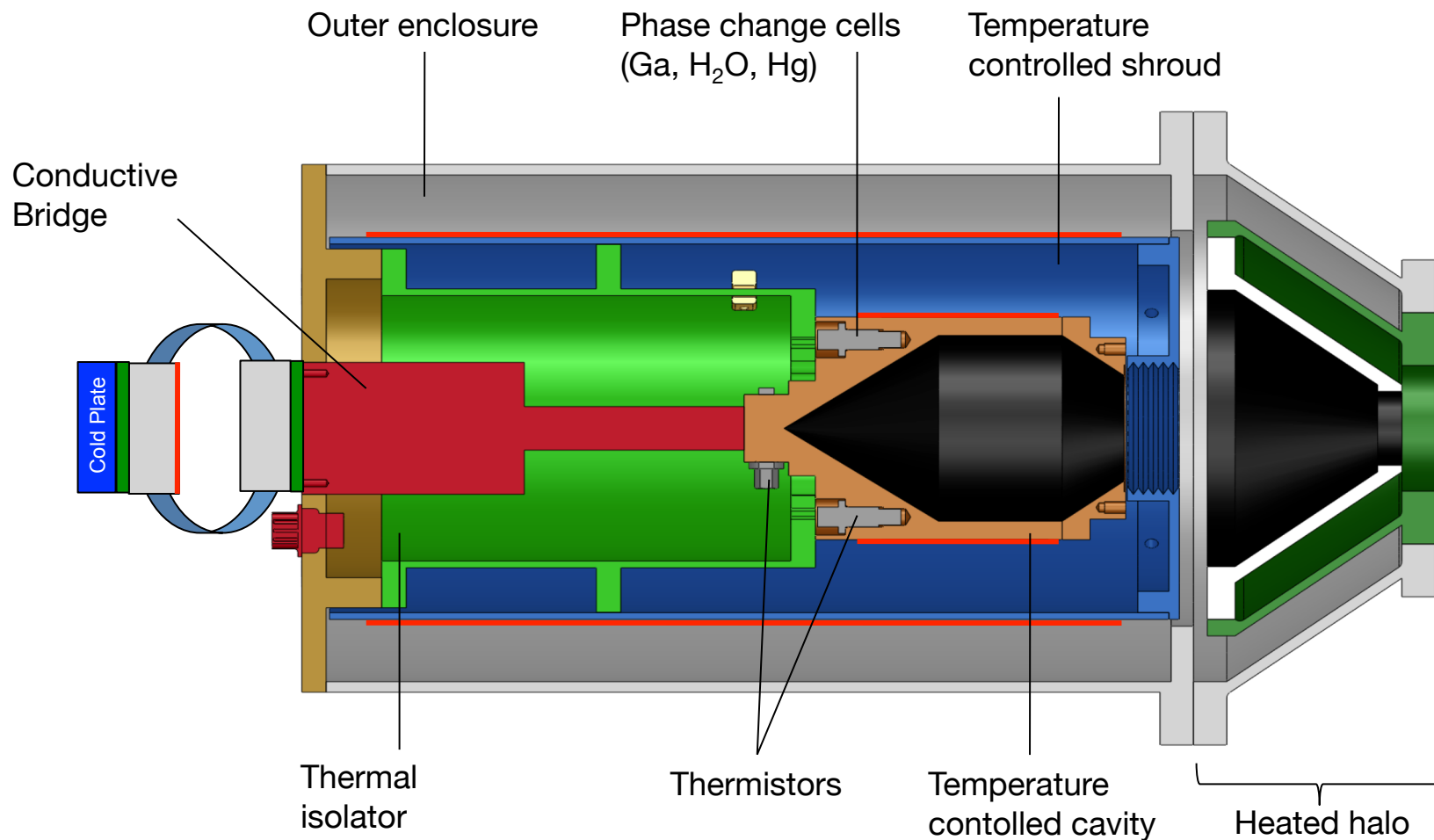
Traceable Measurements for Climate

- Goody *et al* 1998 *BAMS*
- Anderson *et al* 2004 *JQSRT*
- Revercomb *et al* 2005 *OSA*
- Dykema and Anderson 2006 *Metrologia*
- National Research Council 2007 Decadal Survey
- Best *et al* 2007 *CALCON*
- Ohring 2008 *ASIC3*
- Leroy *et al* 2008 *J. Climate*

Absolute Radiance Interferometer

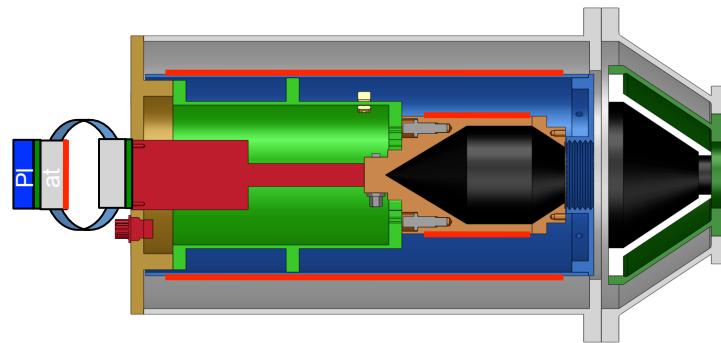


On-orbit Absolute Radiance Standard



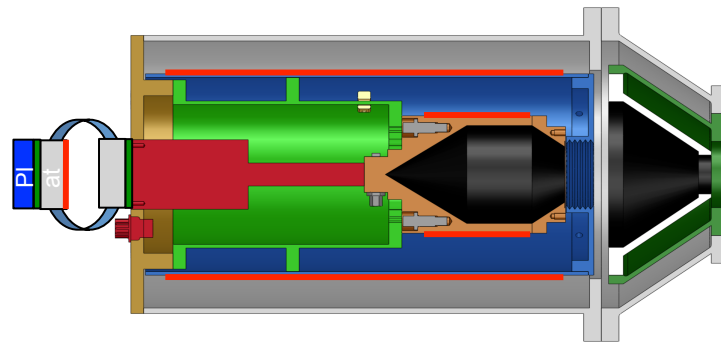
Heritage for the On-orbit Absolute Radiance Standard

- Krutikov *et al* 2006 *Metrologia*
- Gero *et al* 2008 *J. Atmos. Oceanic Technol.*
- Best *et al* 2008 *SPIE*
- Gero *et al* 2009 *J. Atmos. Oceanic Technol.*
- Gero *et al* 2011 *Metrologia* (submitted)



NEWRAD 2011: Absolute Radiance Interferometer

- [Joe Taylor](#), The University of Wisconsin Space Science and Engineering Center Absolute Radiance Interferometer
- [Fred Best](#), On-orbit absolute radiance standard for future IR remote sensing instruments
- [John Dykema](#), Infrared laser-based reflectance measurements for blackbody cavity emissivity determination



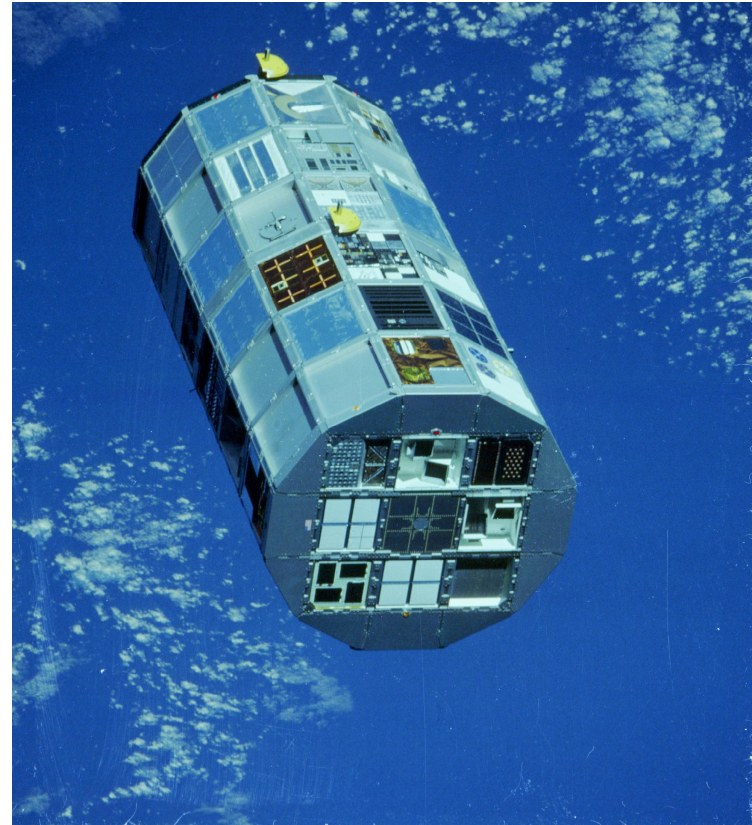
Paint Degradation in Space

Long Duration Exposure Facility

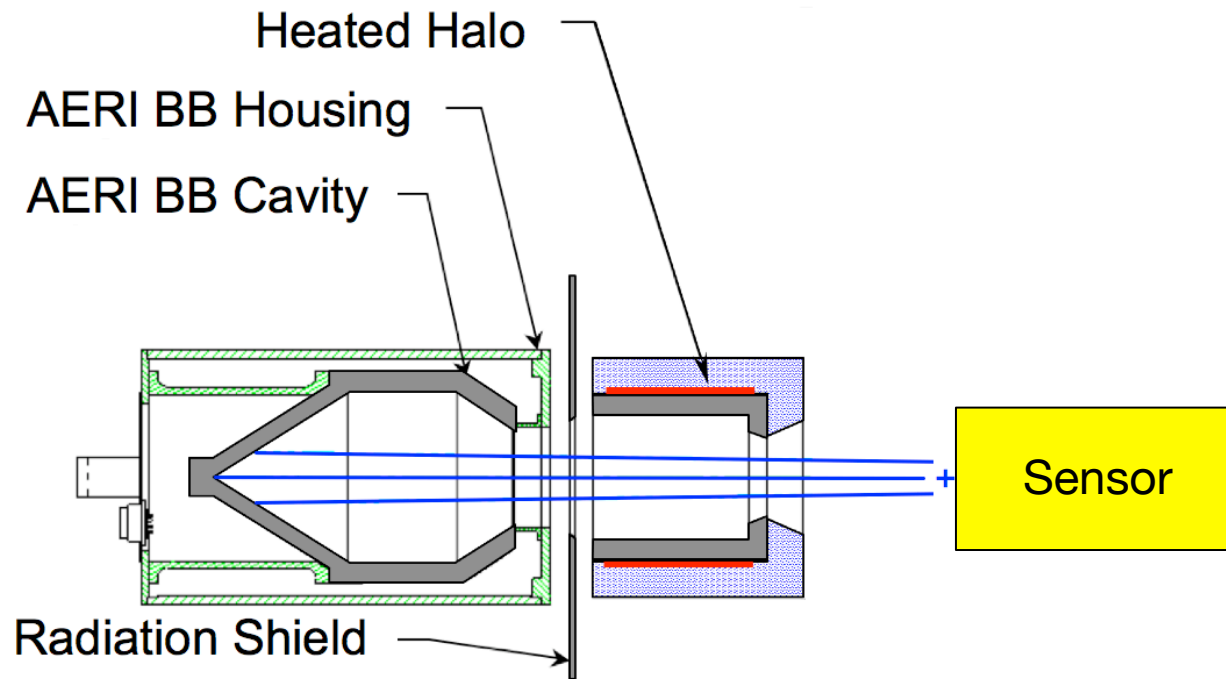
- Study effects of LEO exposure on various materials
- In LEO 1984-1990 (5.7 years)
- Samples of Z306 on Aluminum

Results

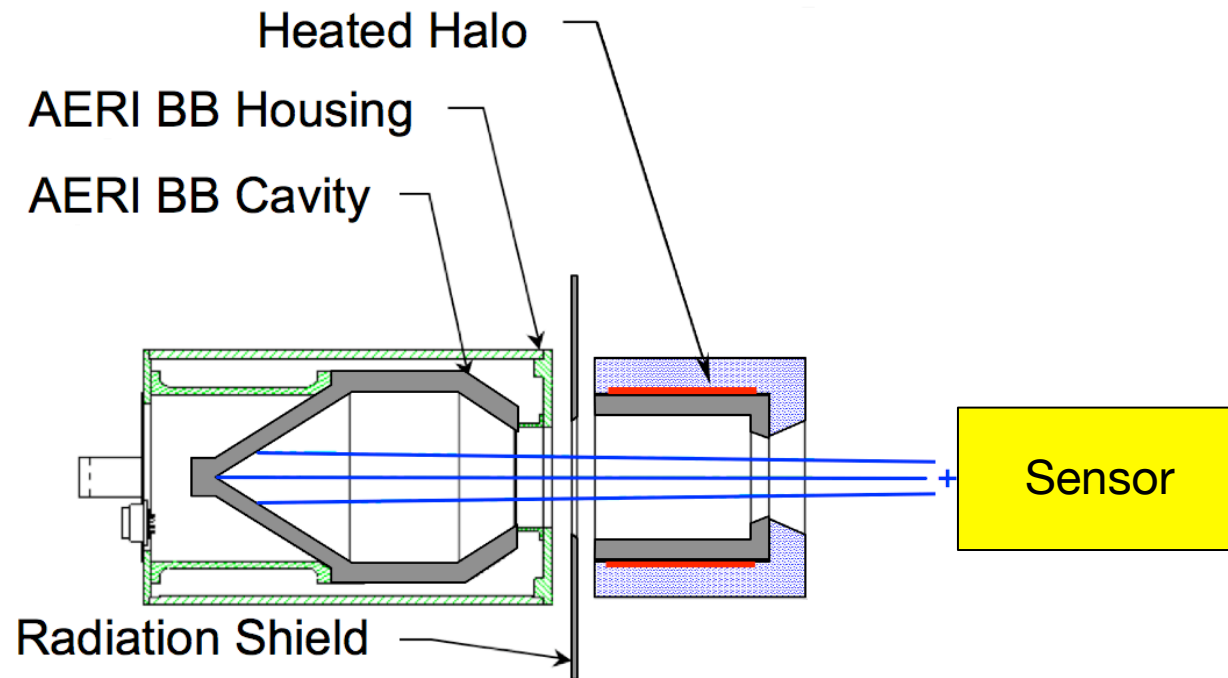
- Evidence of oxidation, erosion, removal of resins, appearance of silicate residues, cracking
- Quantitative changes in optical properties



Heated Halo Concept



Heated Halo Concept



$$R_{\text{obs}} = \underbrace{\varepsilon \cdot B(T_{\text{bb}})}_{\text{Direct radiance from BB}} + \underbrace{(1 - \varepsilon) \cdot [F \cdot B(T_{\text{halo}}) + (1 - F) \cdot B(T_{\text{room}})]}_{\text{Reflected radiance from BB}}$$

Emissivity Calculation

Observed radiance:

$$R_{\text{obs}} = \varepsilon \cdot B(T_{\text{bb}}) + (1 - \varepsilon) \cdot R_{\text{bg}},$$

Emissivity Calculation

Observed radiance:

$$R_{\text{obs}} = \varepsilon \cdot B(T_{\text{bb}}) + (1 - \varepsilon) \cdot R_{\text{bg}},$$

$$R_{\text{bg}} = [F \cdot B(T_{\text{halo}}) + (1 - F) \cdot B(T_{\text{room}})]$$

Emissivity Calculation

Observed radiance:

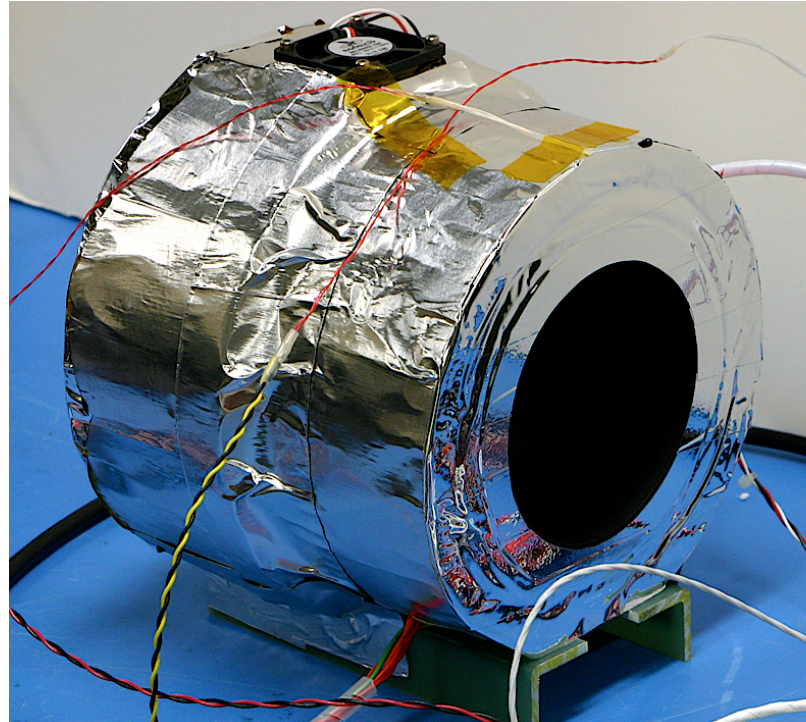
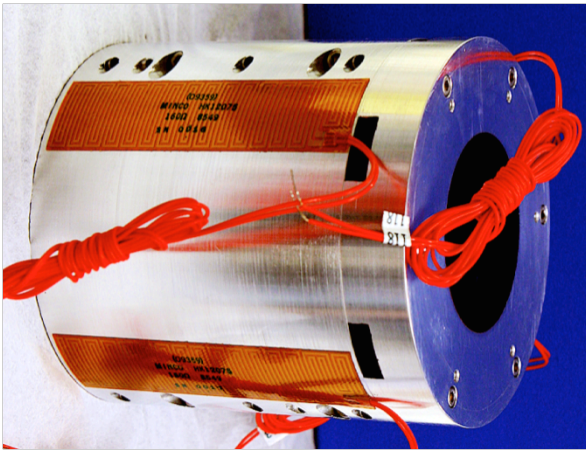
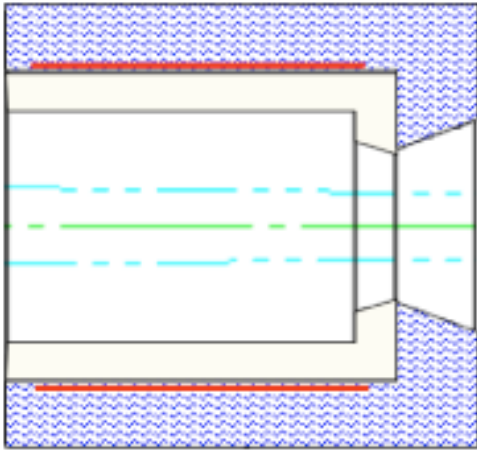
$$R_{\text{obs}} = \varepsilon \cdot B(T_{\text{bb}}) + (1 - \varepsilon) \cdot R_{\text{bg}},$$

$$R_{\text{bg}} = [F \cdot B(T_{\text{halo}}) + (1 - F) \cdot B(T_{\text{room}})]$$

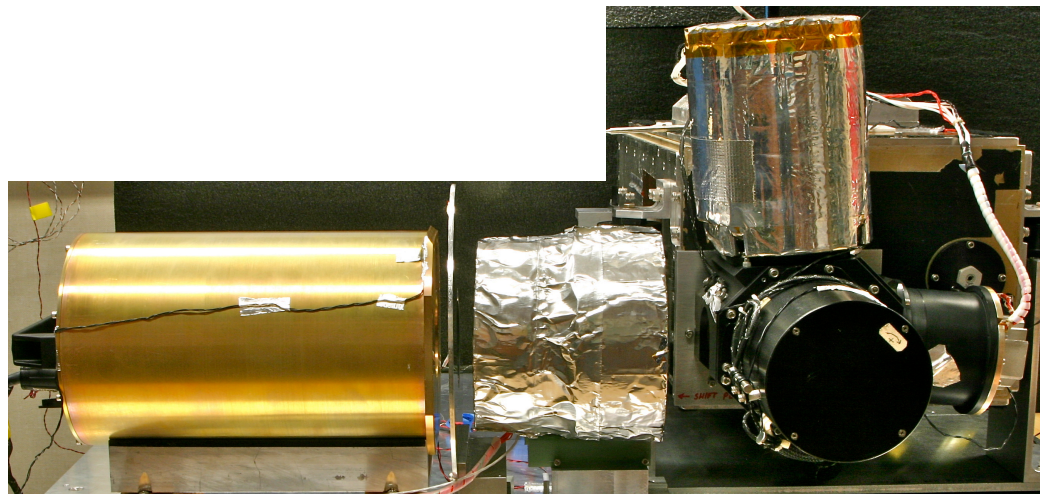
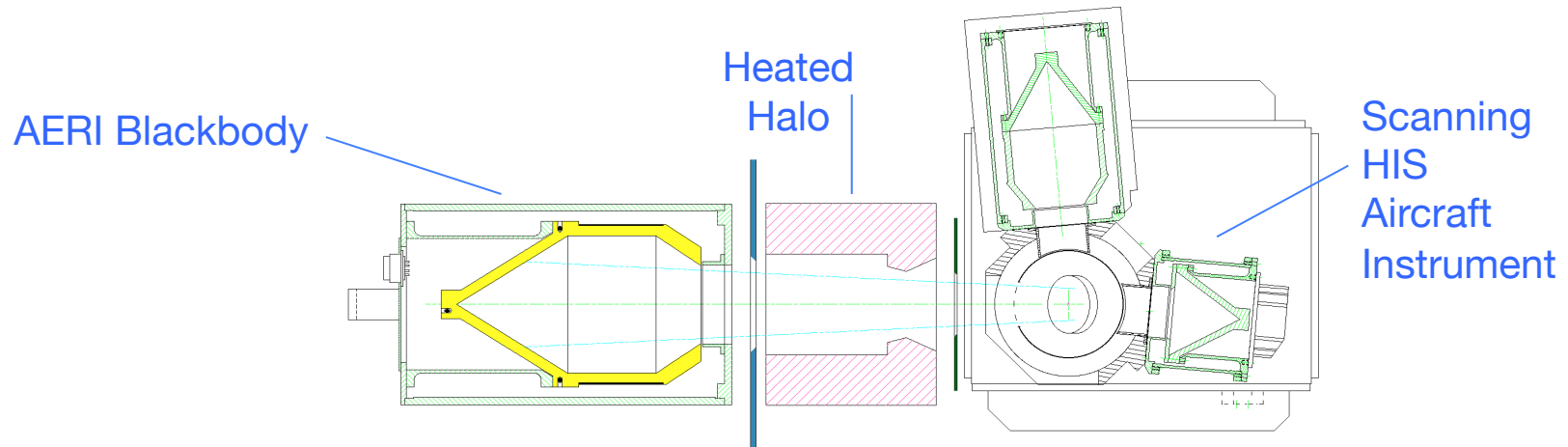
Emissivity/reflectivity measurement:

$$\langle 1 - \varepsilon(t) \rangle_t = \left\langle \frac{R_{\text{obs}}(t) - B[T_{\text{bb}}(t)]}{R_{\text{bg}}(t) - B[T_{\text{bb}}(t)]} \right\rangle_t$$

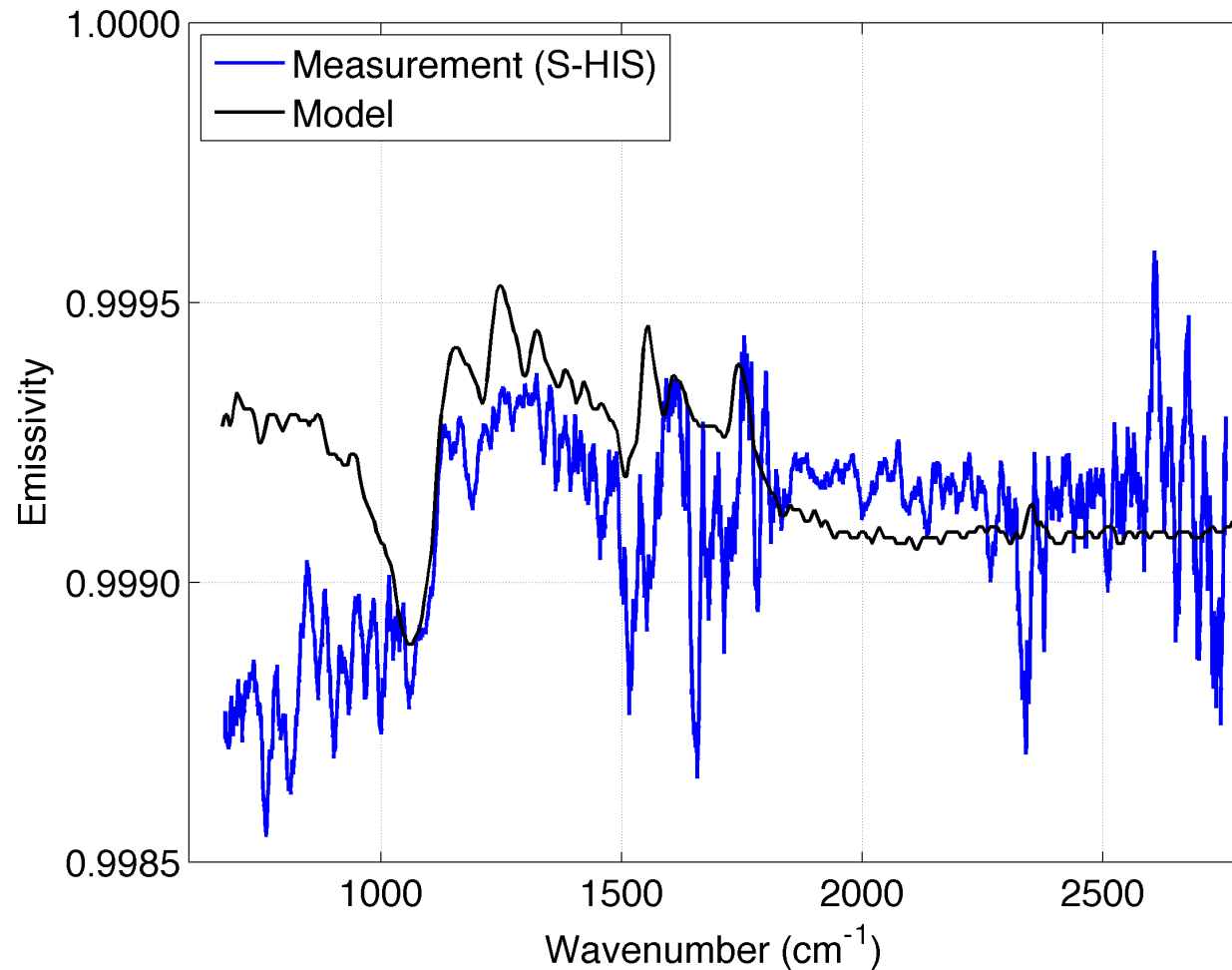
Heated Halo Gen. 1



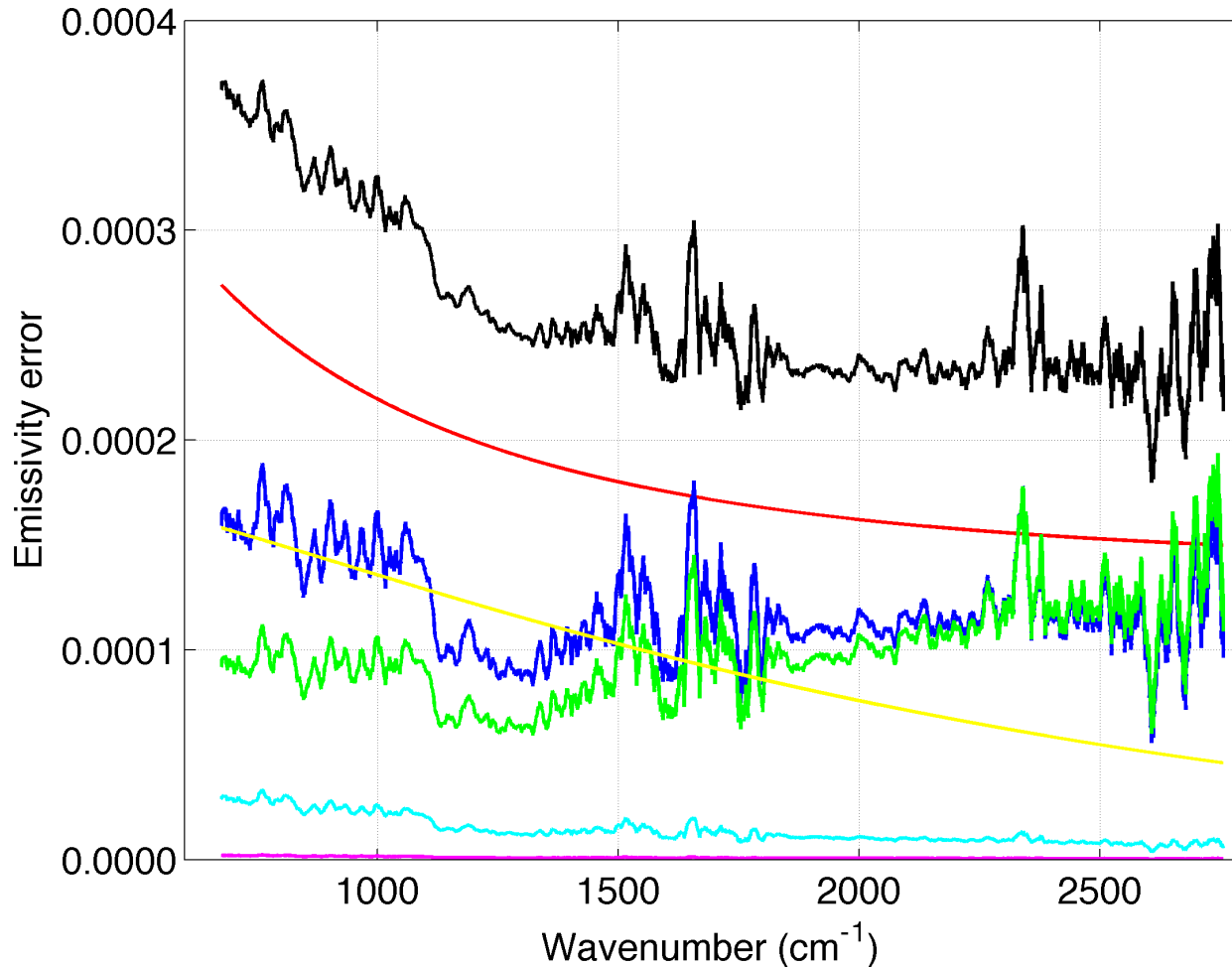
Heated Halo Gen. 1 Test Configuration (S-HIS)



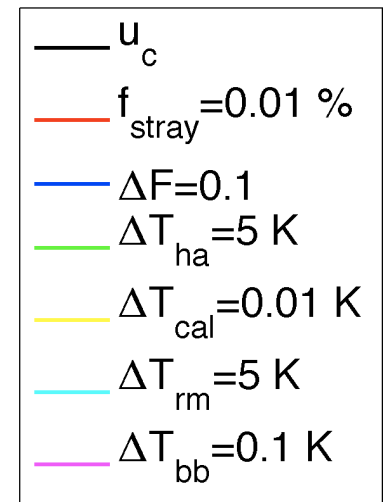
AERI Blackbody Emissivity (Halo 1, S-HIS)



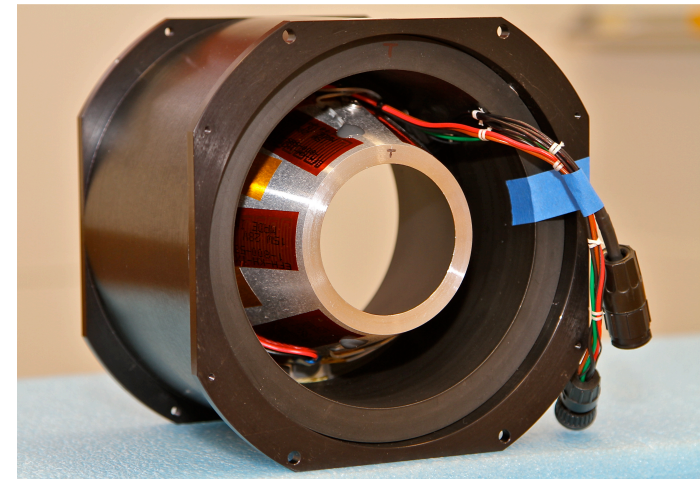
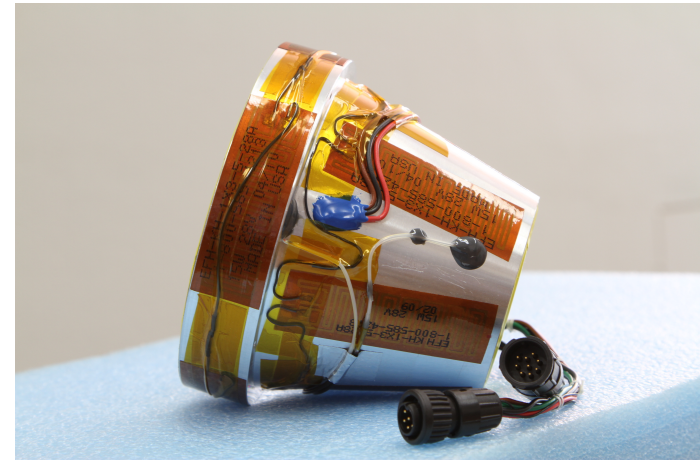
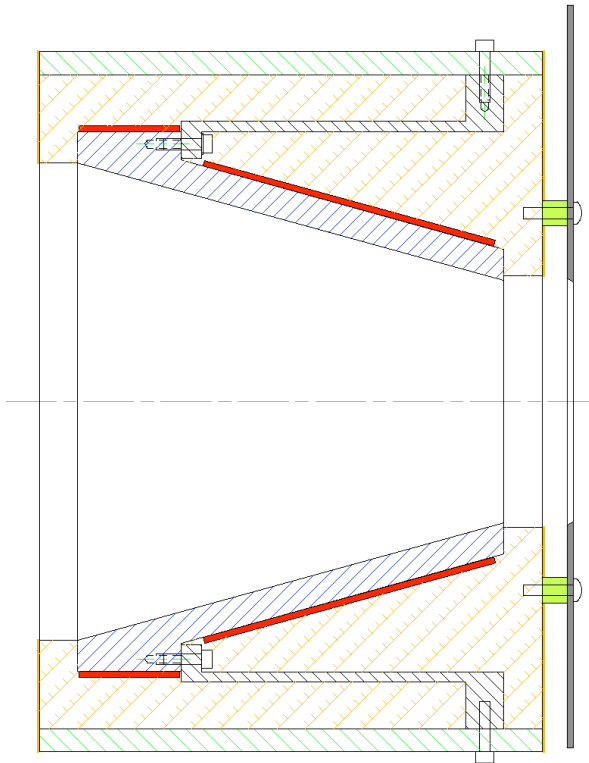
Emissivity Uncertainty (Halo 1, S-HIS)



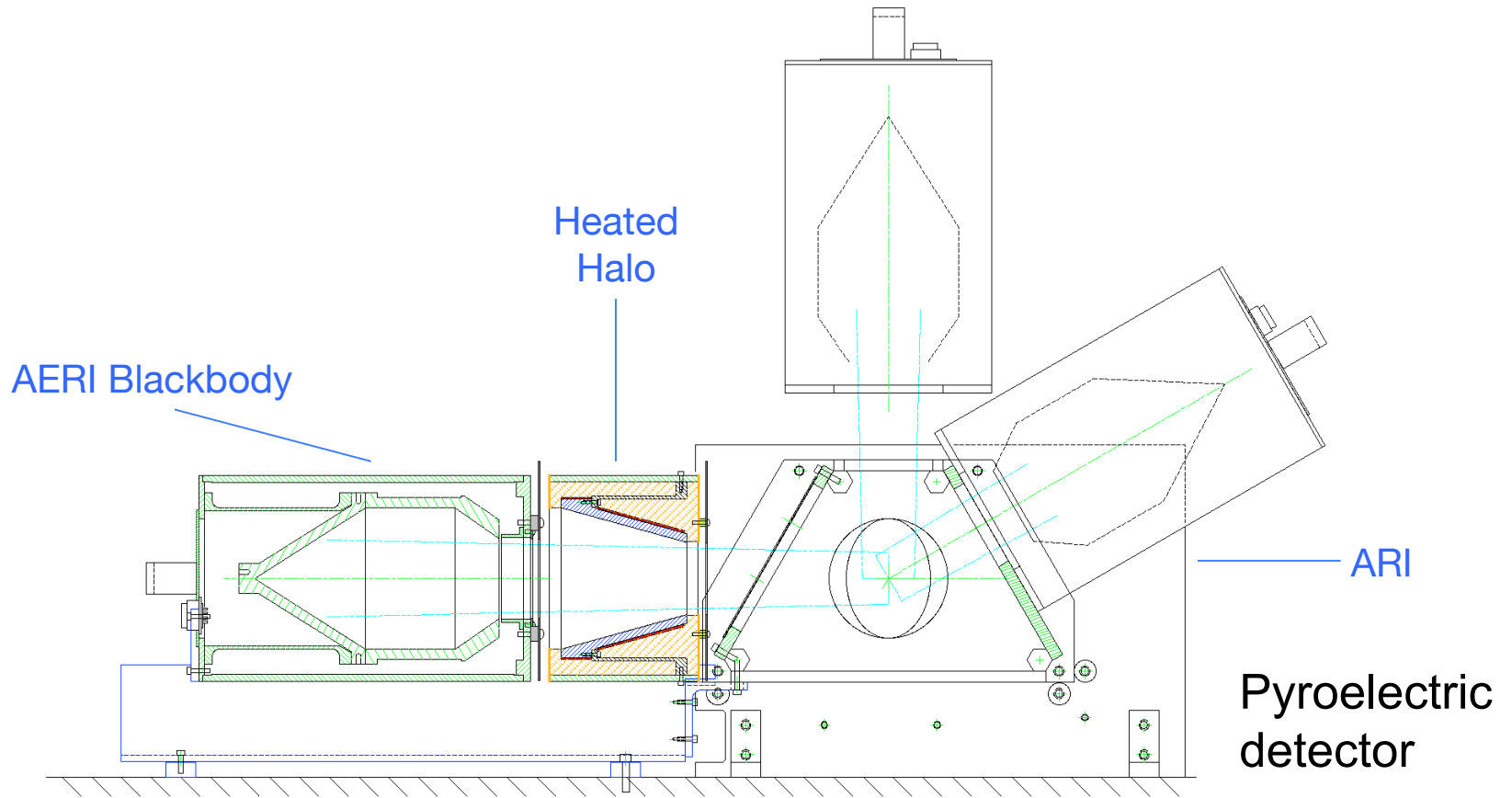
Type B
measurement
uncertainty
($k = 3$)



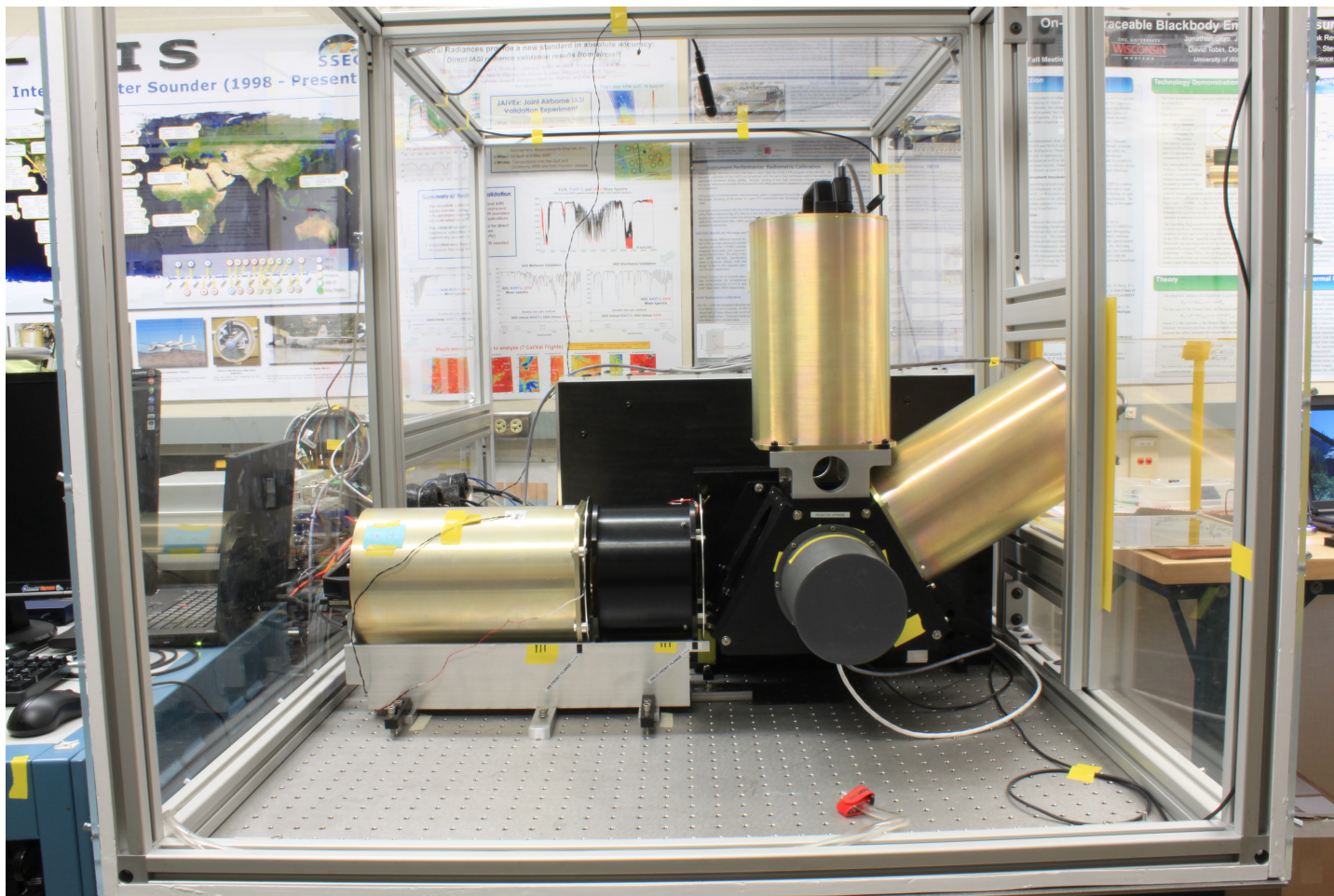
Heated Halo Gen. 2



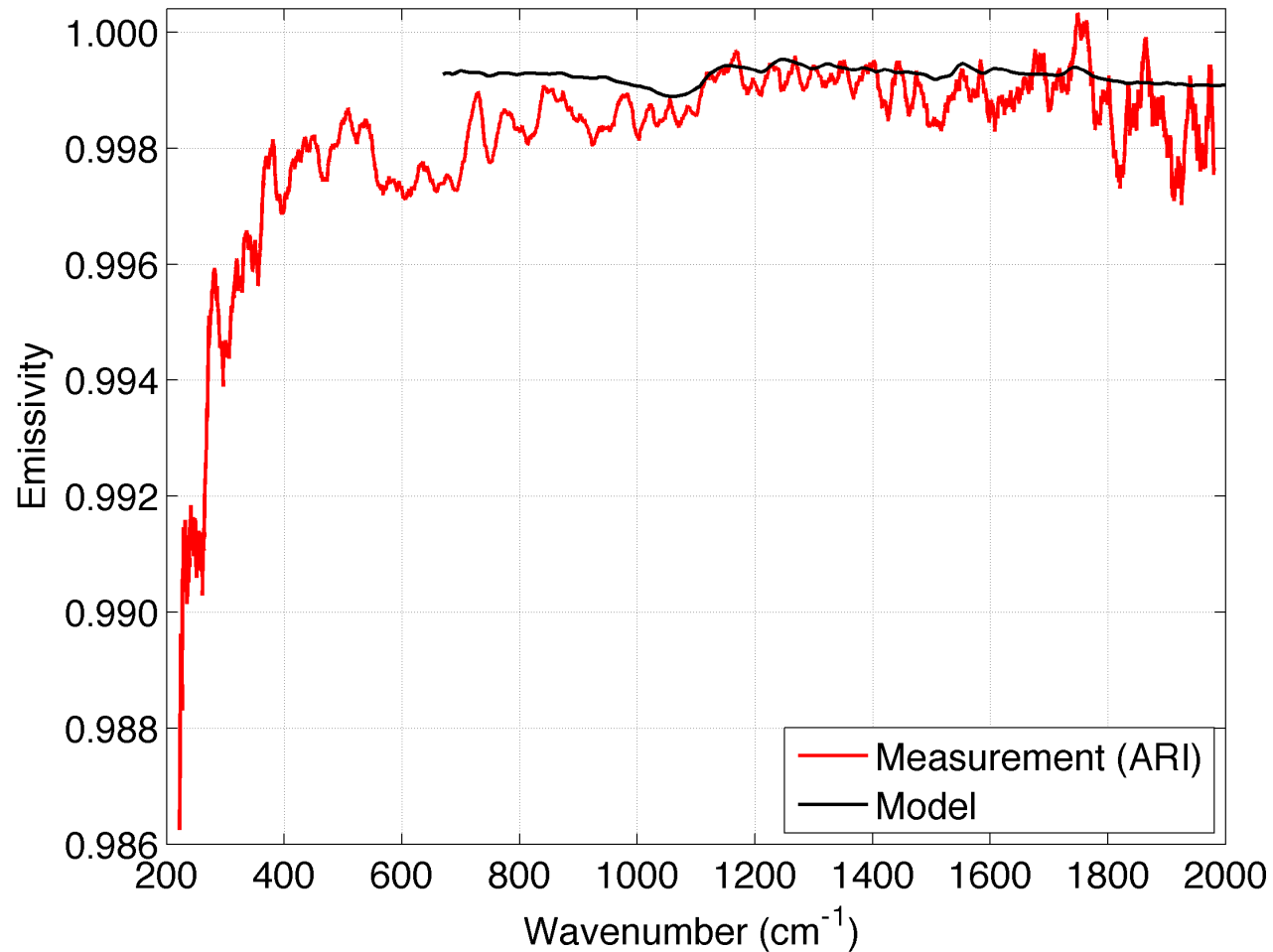
Heated Halo Gen. 2 Test Configuration (ARI)



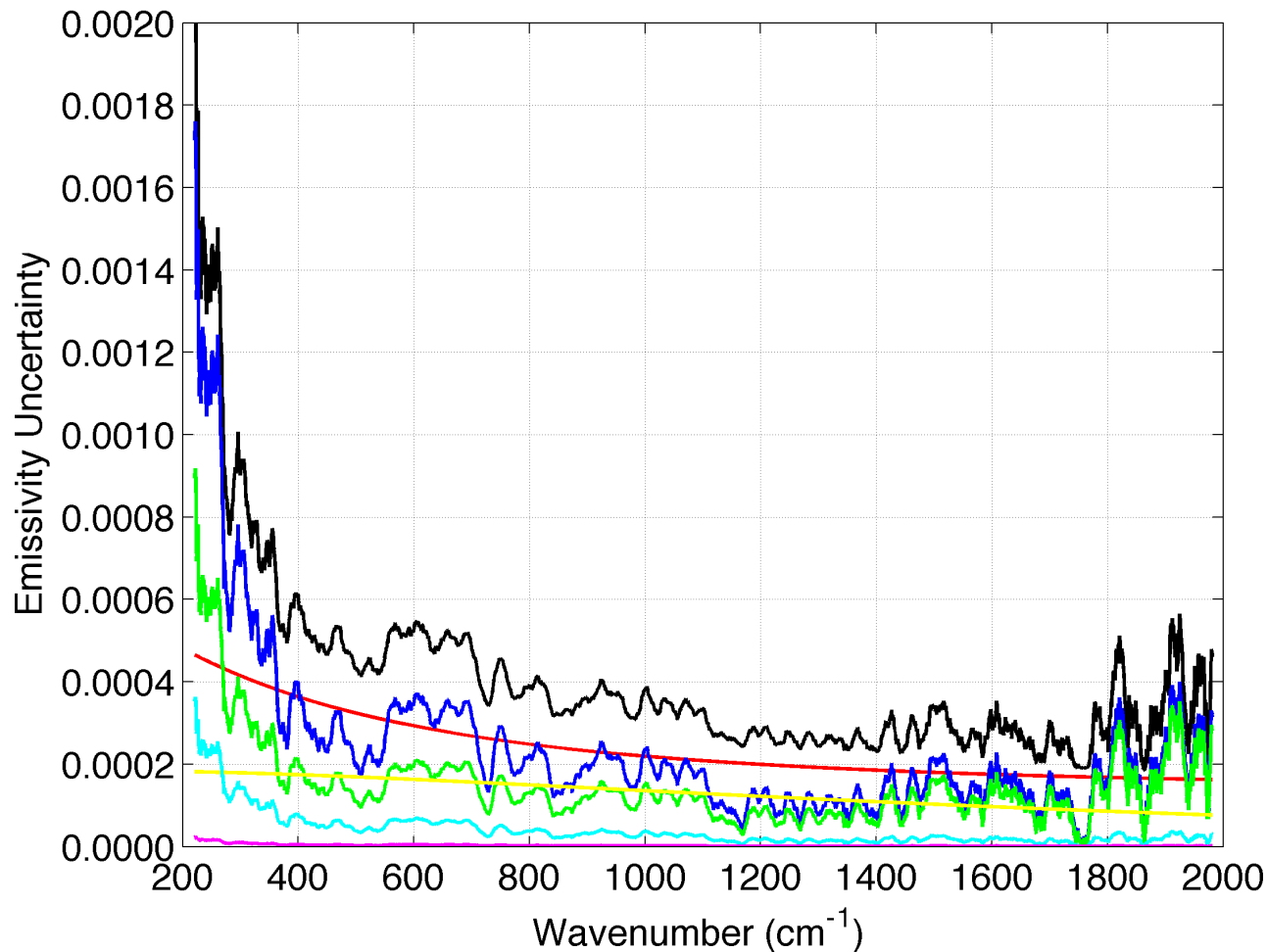
Heated Halo Gen. 2 Test Configuration (ARI)



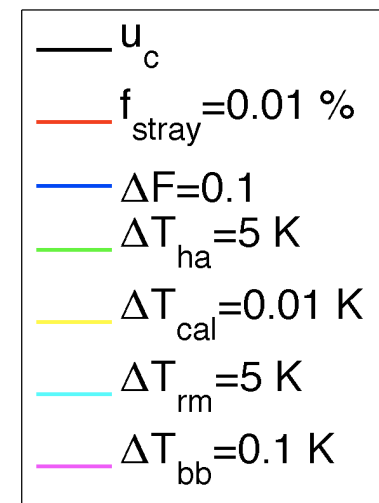
AERI Blackbody Emissivity (Halo 2, ARI)



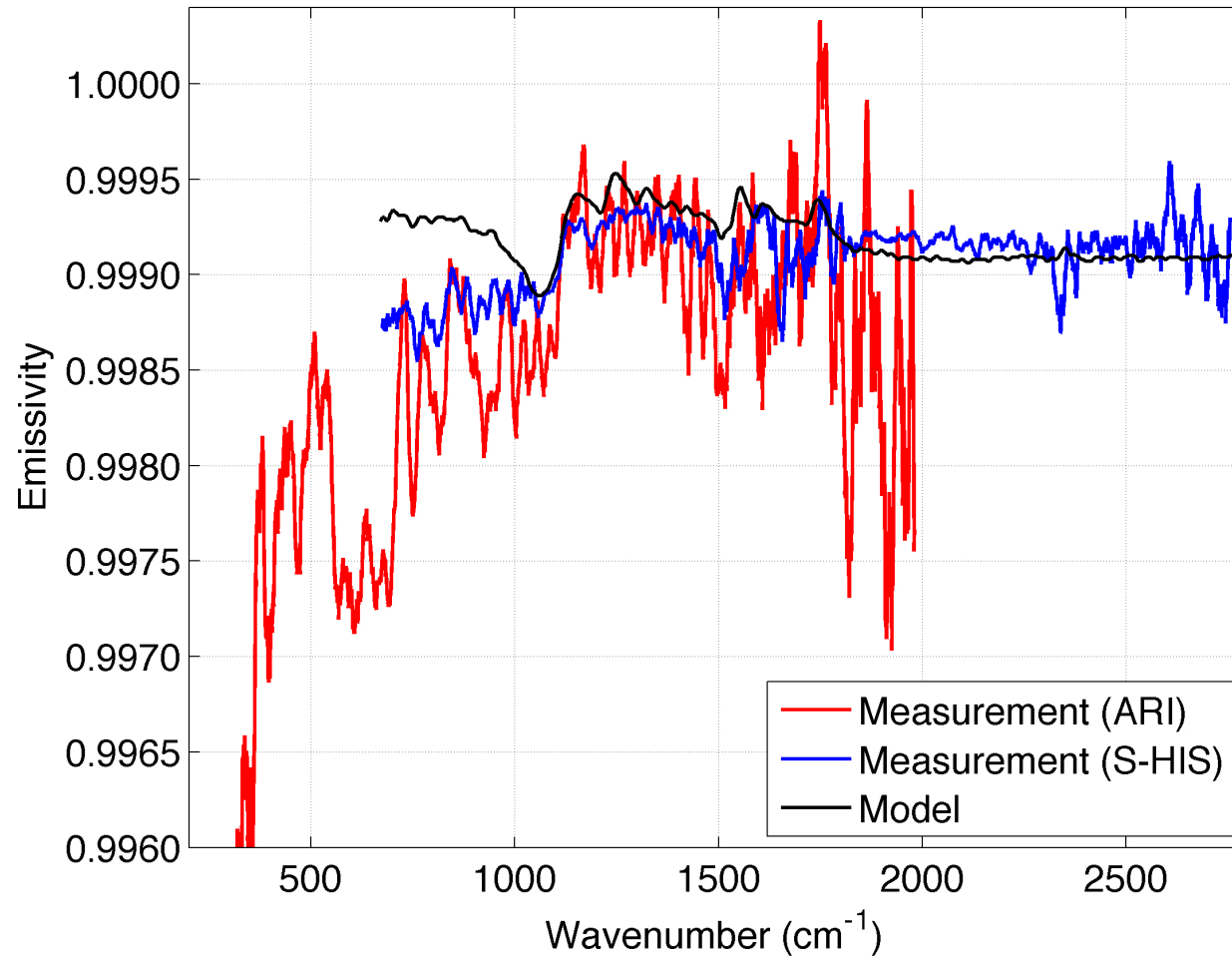
Emissivity Uncertainty (Halo 2, ARI)



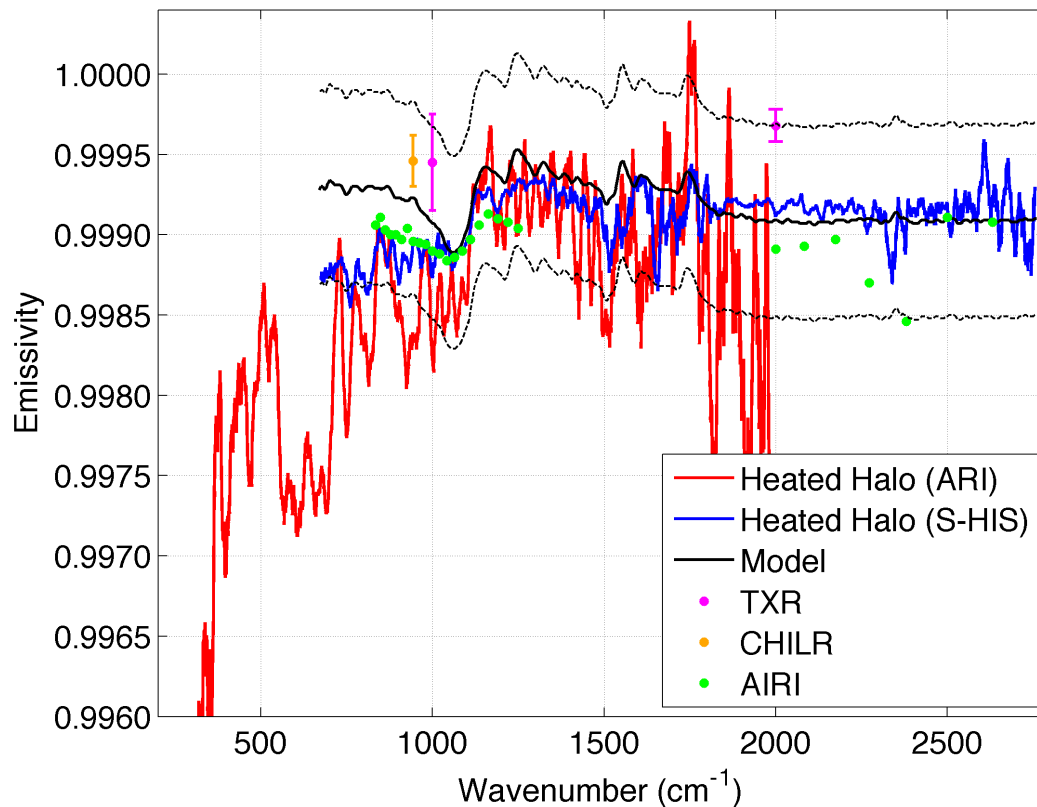
Type B
measurement
uncertainty
($k = 3$)



AERI Blackbody Emissivity (Halo 1, Halo 2)

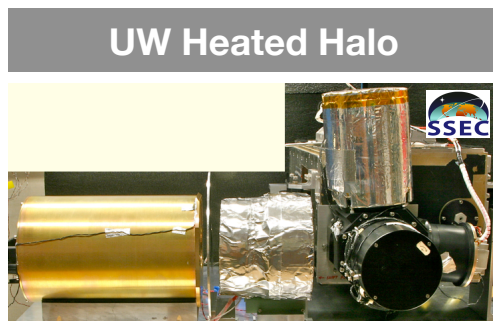


AERI Blackbody Emissivity Comparison



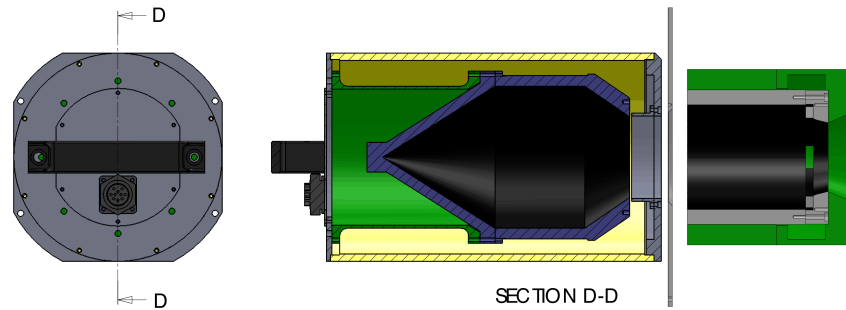
Comparison with NIST measurements

Continued work corroborates earlier results and helps reduce uncertainty



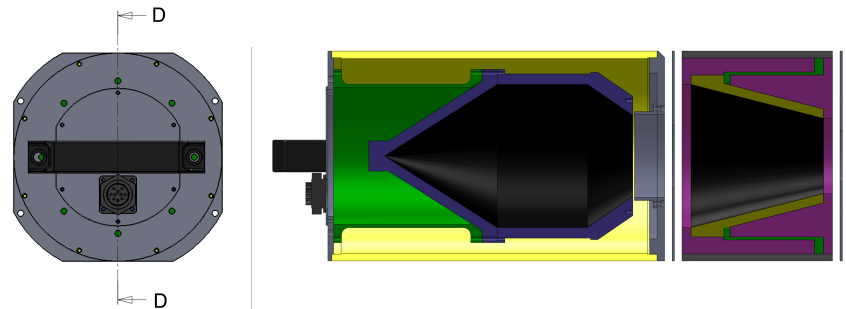
Next Generation Heated Halo

Gen. 1



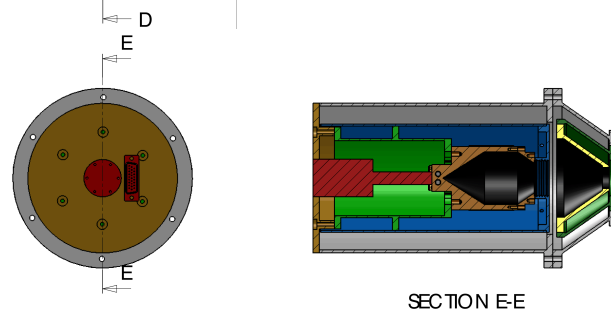
TRL 4

Gen. 2



TRL 5

Gen. 3

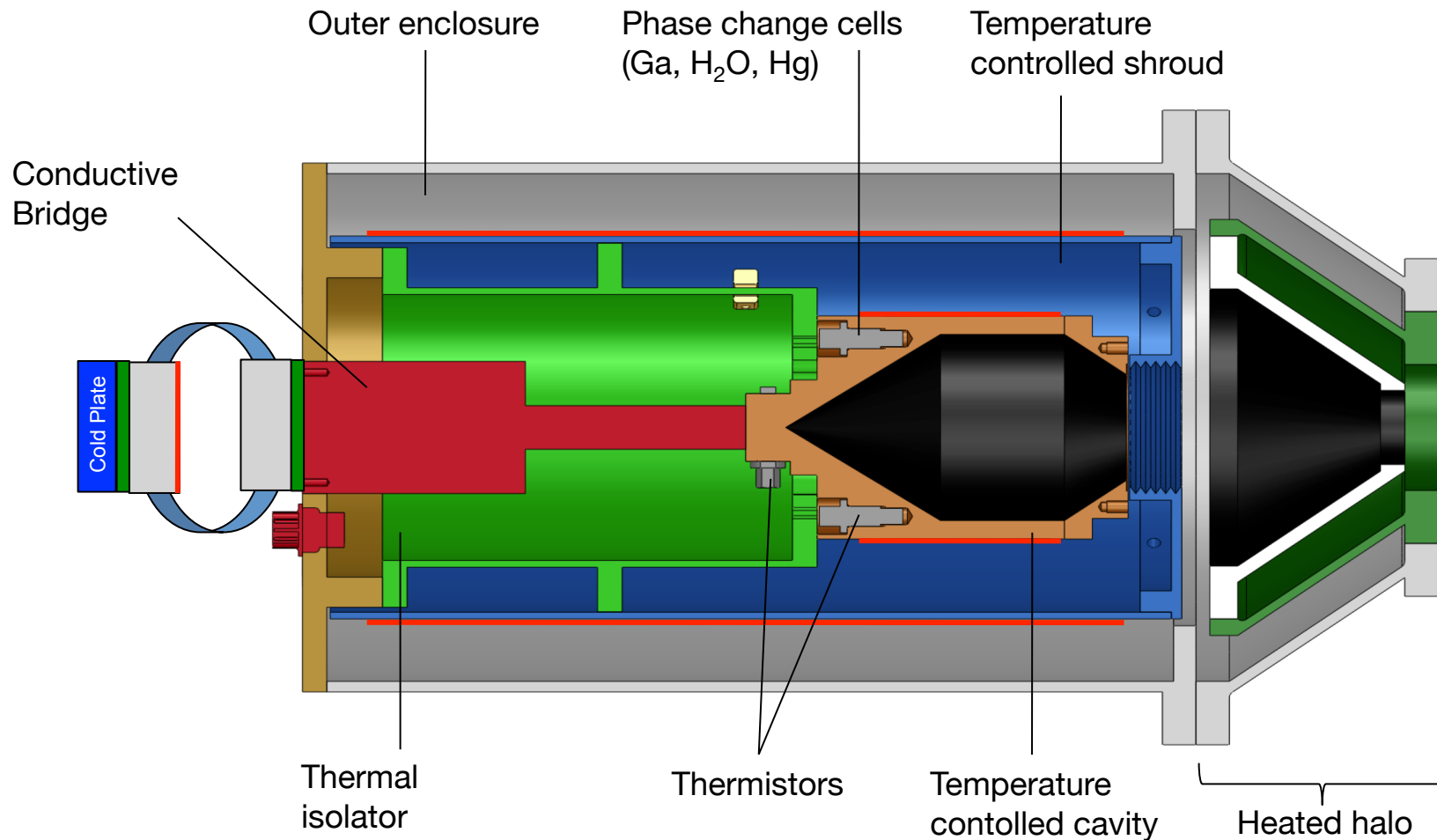


TRL 6

Summary

- Spectral emissivity measurement has been demonstrated with the Heated Halo configured with both the S-HIS and the ARI, using an AERI blackbody as the target
- 0.0006 measurement uncertainty achievable across most of the thermal infrared
- Primary “lesson learned” is the importance of controlling stray light contributions
- Agreement between observations using two different instruments validates the process for emissivity measurement with the Heated Halo

On-orbit Absolute Radiance Standard



AERI Blackbody Emissivity Comparison

