

# ***Future Long-term Measurements of Solar Spectral Irradiance by the TSIS Spectral Irradiance Monitor: Improvements in Measurement Accuracy and Stability***



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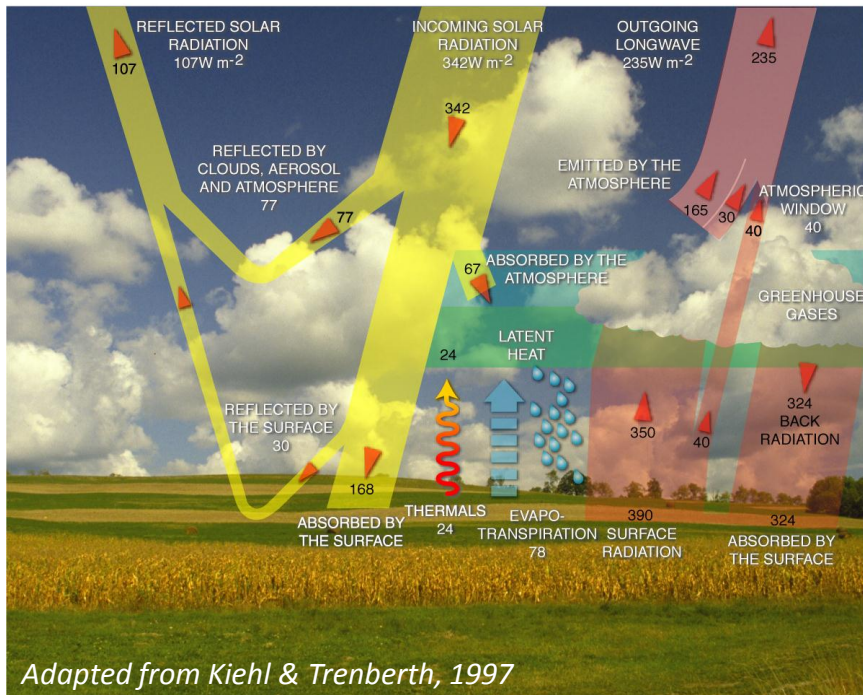


# Spectral Irradiance Contributions Key to Climate Issues

## • **What is the solar forcing at decadal and longer timescales?**

- **Solar Irradiance Climate Data Record (CDR):** time series of measurements of sufficient length, consistency, and continuity to determine climate variability and change.

### Global Energy Budget Contributions



## Solar Radiative Forcing Questions

### • **How does the climate system respond?**

✓ **Requires measurement of wavelength-dependent irradiance variability.**

**Knowledge of TOA spectral distribution of solar radiation is crucial in interpreting the highly spectrally dependent radiative processes in atmosphere and at the surface.**

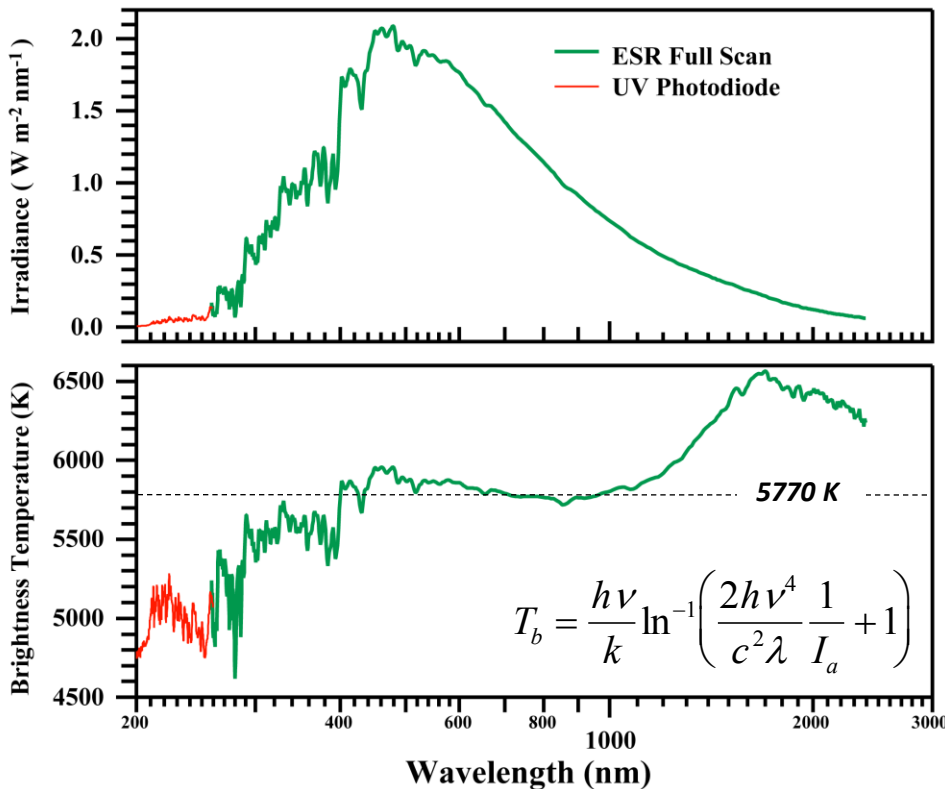


# SIM Measures Solar Spectral Irradiance (96% TSI)

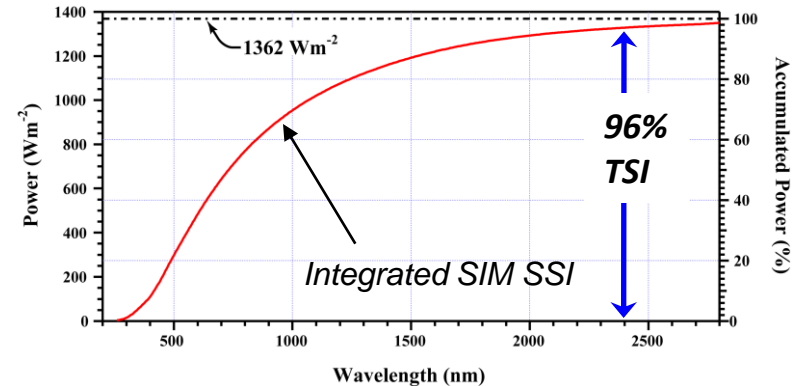
## Solar Spectral Irradiance ( $W m^{-2} nm^{-1}$ )

Solar Spectral Irradiance is defined as the radiant power per unit area per unit wavelength interval incident on a plane surface at the top of the atmosphere that is normal to the direction from the Sun.

## SORCE SIM (200 – 2400 nm)



## SIM Integrated Power vs. Wavelength

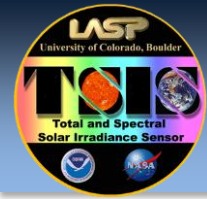


Total Solar Irradiance (TSI)

$$TSI_{TIM} = \int_{\lambda=0}^{\lambda=\infty} E_{\lambda} d\lambda \approx 1362 \text{ Watts/m}^2$$

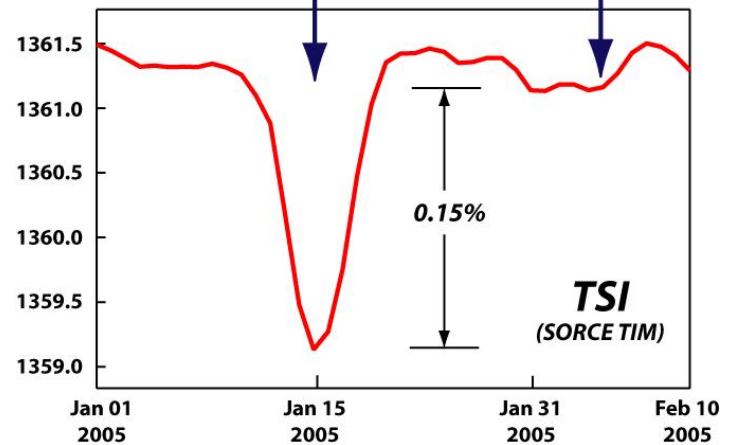
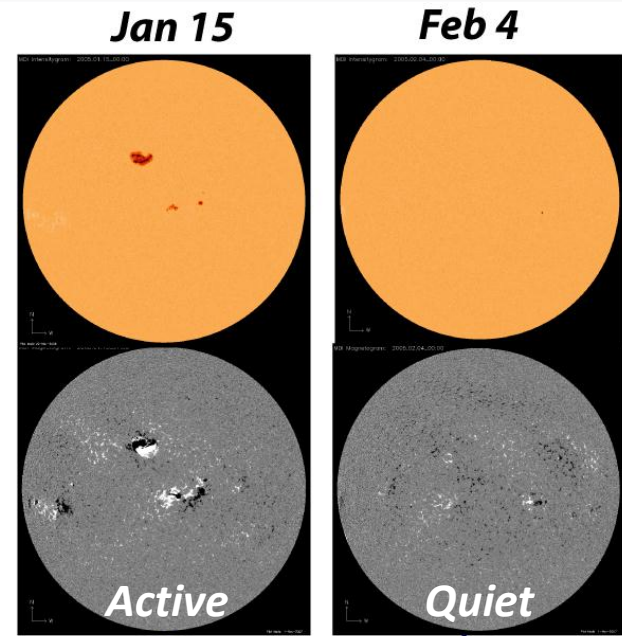
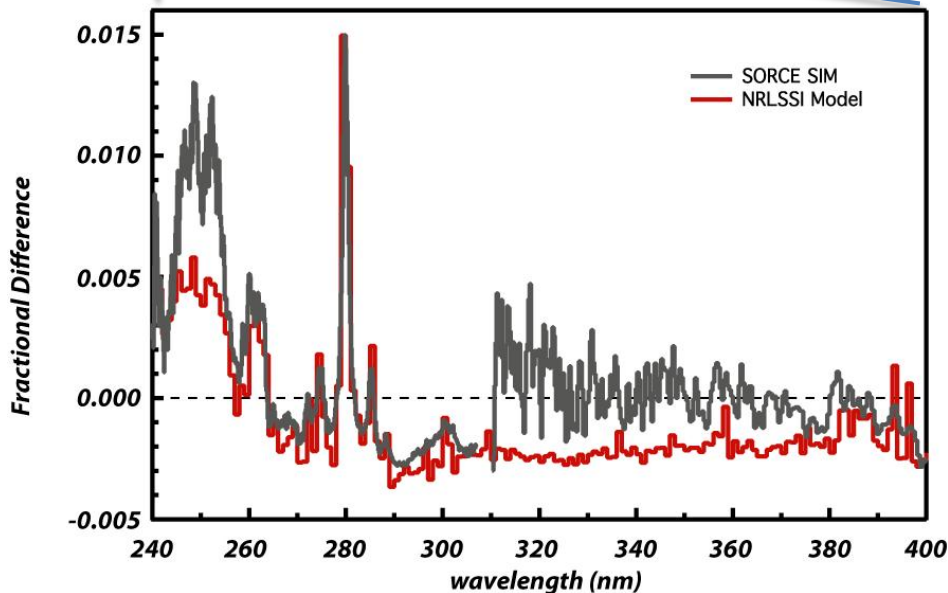
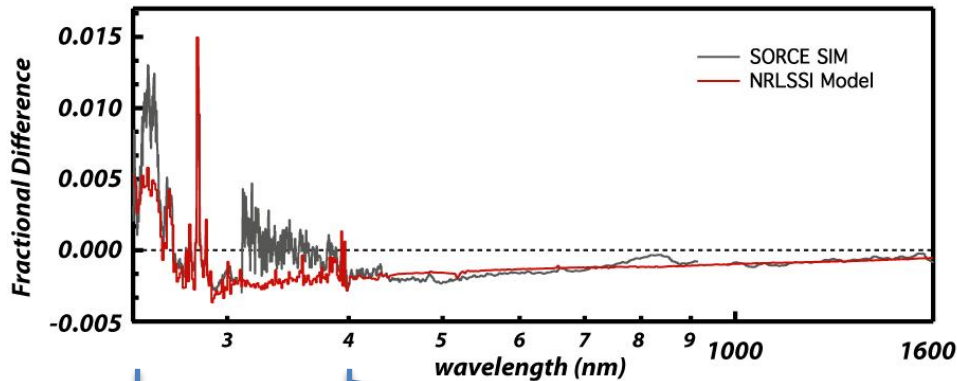
Spectral Solar Irradiance (SSI)

$$TSI_{SIM} = \int_{\lambda=200}^{\lambda=2400} E_{\lambda} d\lambda \approx 96\% \text{ of TSI}$$



# Short-term (Rotational) SSI Variability

## Spectral Irradiance Contrast

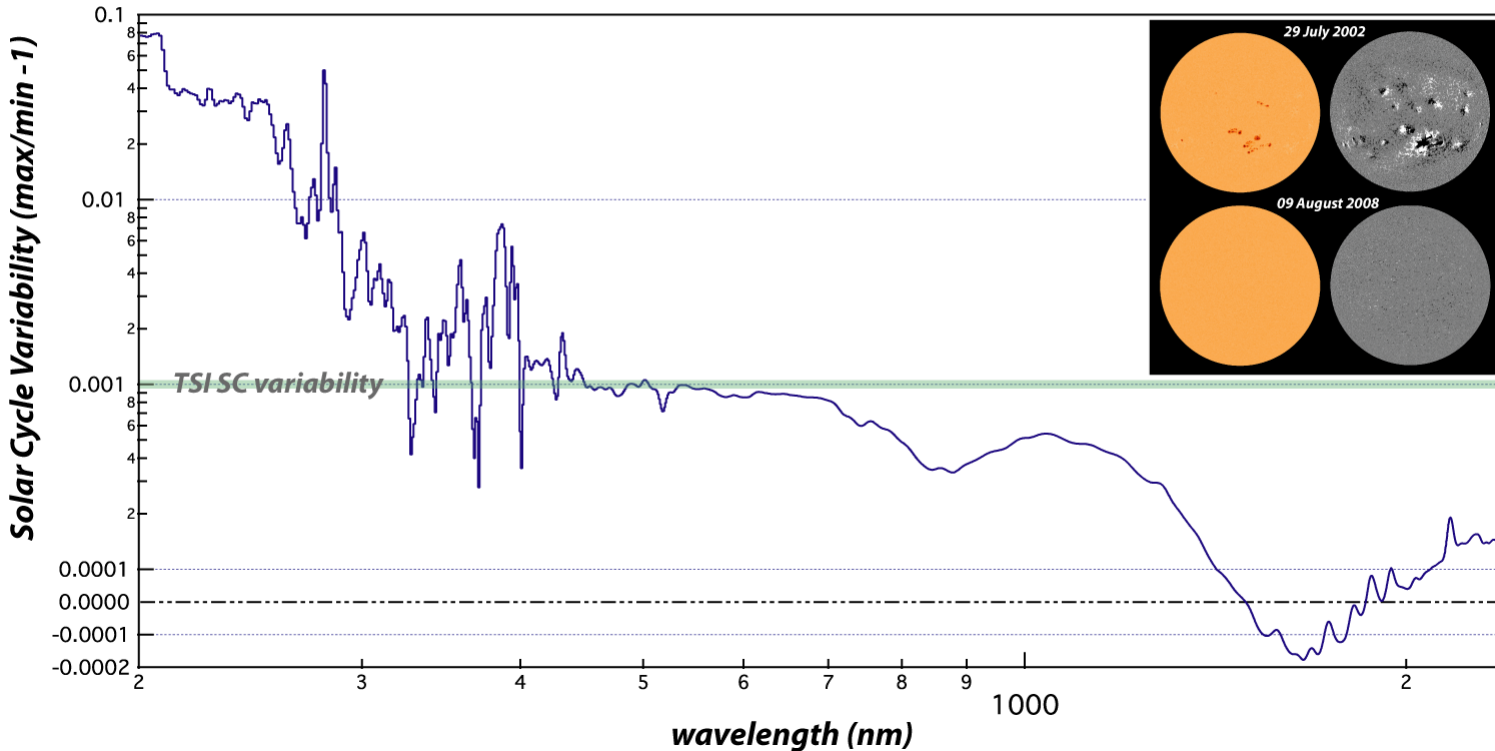






# Predicted Solar Cycle SSI Variability

↔ 174-242 nm: O<sub>2</sub> dissociation - O<sub>3</sub> production (1.8 W/m<sup>2</sup>)  
 ↔ 242-300 nm: O<sub>3</sub> absorption (12 W/m<sup>2</sup>)  
 ← 300 - >2000 nm: Climate Forcing (~1355 W/m<sup>2</sup>) →



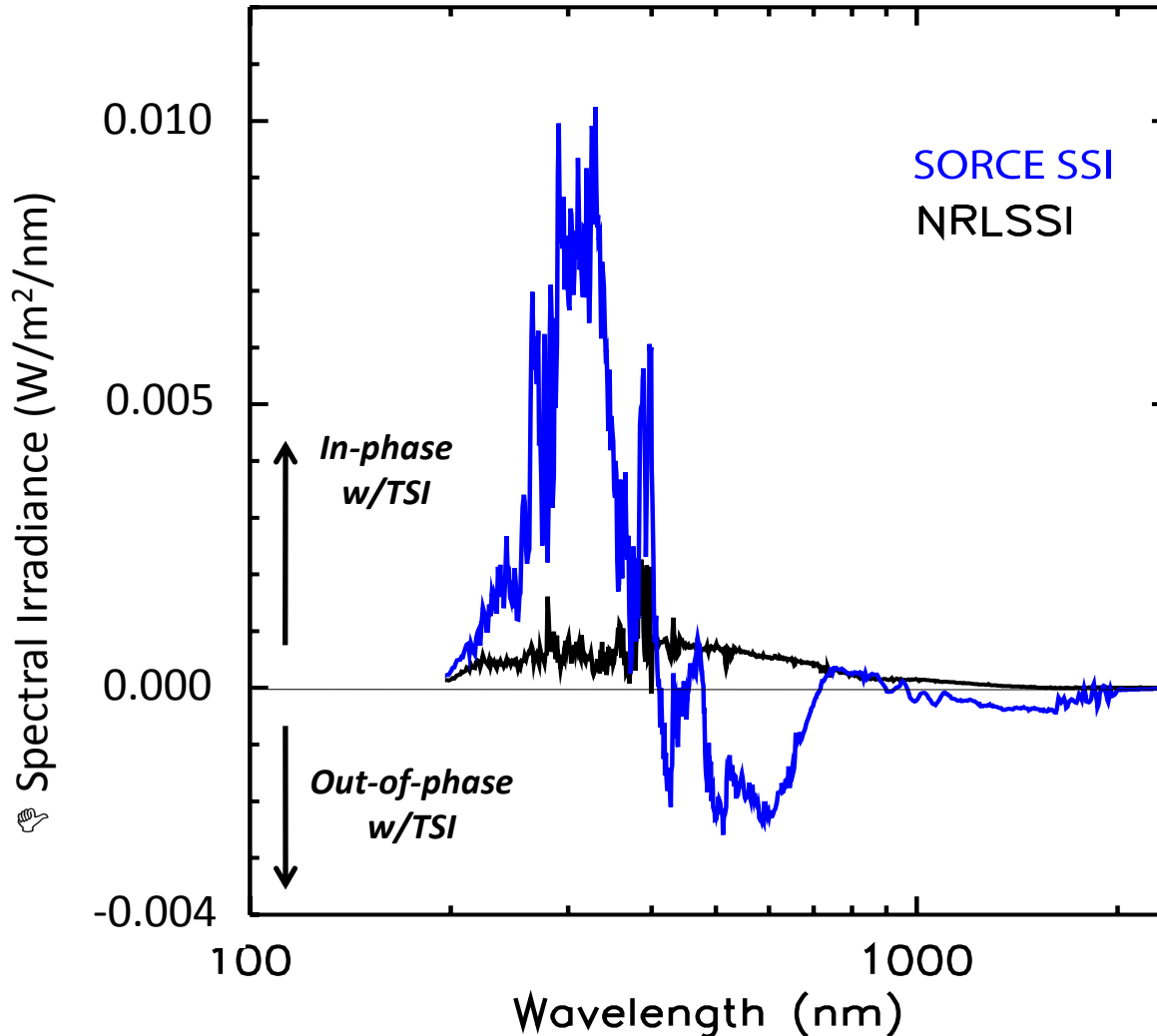
**NRLSSI Modeled spectral variability based on observations of UV (120-250 nm) and model of rotational modulation of plage and sunspot contrast.**

**Prior to *SORCE SIM* (2003) no continuous measurement of variability in the 400-2400 nm region**



# Observed “Long-term” SSI Variability

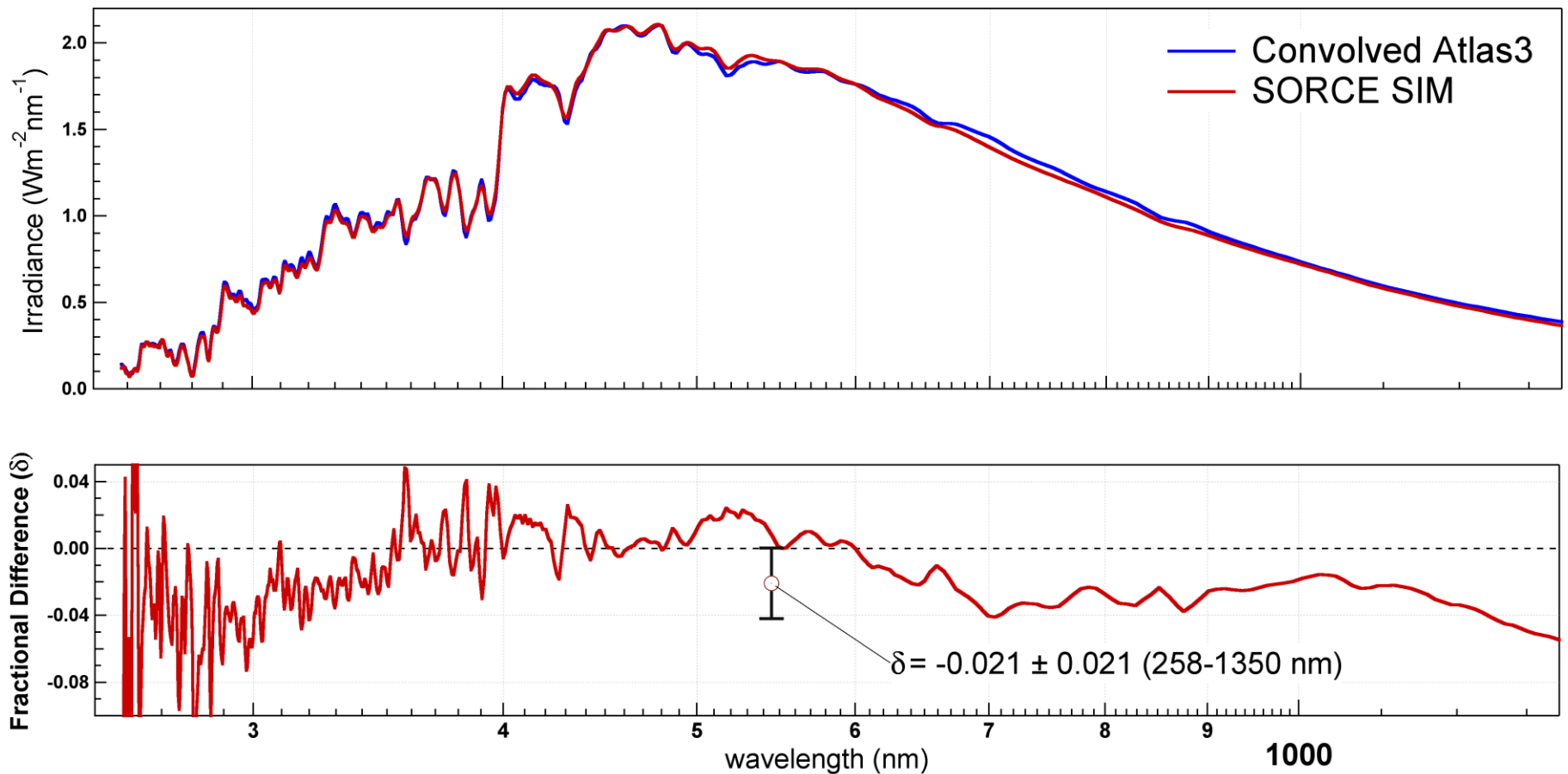
*Spectral Irradiance difference between Active in mid 2004 and Quiet in late 2007*

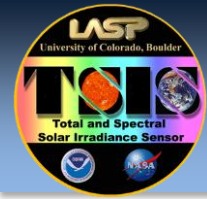




# SSI Absolute Accuracy Issues

## SORCE SIM and Atlas 3 Comparison (258-1350 nm)





# Attributes of a Climate Monitoring System

***“Design of climate observing and monitoring systems must ensure the establishment of global, long-term climate records that are of high accuracy, tested for systematic errors on-orbit, and tied to irrefutable international standards maintained in the U.S. by the National Institute of Standards and Technology (NIST).”***

***- Achieving Satellite Instrument Calibration for Climate Change (ASIC<sup>3</sup>)  
2007 Final Report***





# JPSS SSI Measurement Requirements

Attribute	Requirement	Justification
<b>Measurement Range ( <math>Wm^{-2}nm^{-1}</math> )</b> Spectral (0.2-2.4 $\mu m$ )	$10^{-4}$ - $10^1$	<b>Solar Spectrum</b> <i>Full scale of spectral irradiance magnitude</i>
<b>Long-term rel. stability (per year)</b> $0.2 \leq \lambda \leq 0.4 \mu m$ $0.4 < \lambda \leq 2.4 \mu m$	0.05% 0.01%	<b>Interpret Solar Cycle variability</b> <i>UV variability 0.1% - 10%</i> <i>Visible- Near IR variability <math>\leq 0.05\%</math></i>
<b>Measurement precision</b> Spectral (0.2-2.4 $\mu m$ )	0.01%	<b>Measure short term variability</b> <i>Sufficient SNR for Vis-NIR spectral variability</i>
<b>Measurement Accuracy</b> Spectral (0.2-2.4 $\mu m$ )	0.25%	<b>Climate modeling input</b> <i>Earth radaiton budget; Processes &amp; Mechanisms</i>
<b>Reporting Frequency (per day)</b> Spectral (0.2-2.4 $\mu m$ )	2	<b>Solar temporal variability</b> <i>sample short-term spectral varaitions with TSI</i>
<b>Spectral Resolution (nm)</b> $\lambda \leq 0.28 \mu m$ $0.28 \mu m < \lambda \leq 0.40 \mu m$ $\lambda > 0.40 \mu m$	1 5 45	<b>Solar wavelength variability</b> <i>Strongest wavelength dependence of UV variability</i> <i>Broader wavelength dependence of Vis-NIR var.</i>



# TSIS SIM Derives Heritage from SORCE SIM

**TSIS SIM designed for long-term spectral irradiance measurements (climate research)**

*Incorporate lessons learned from SORCE SIM (& other LASP programs) into TSIS SIM to meet measurement requirements for long-term JPSS SSI record*

## **Specific required capabilities over SORCE SIM**

✓ **Reduce uncertainties in prism degradation correction to meet long-term stability requirement**

- **Ultra-clean optical environment to mitigate contamination**
- **Addition of 3<sup>rd</sup> channel to reduce calibration uncertainties**

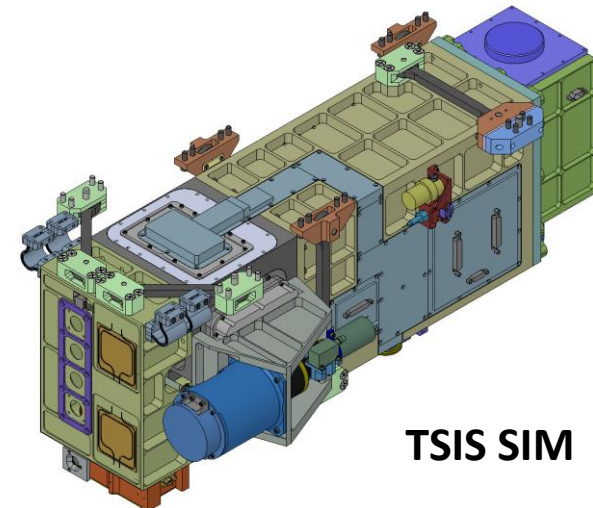
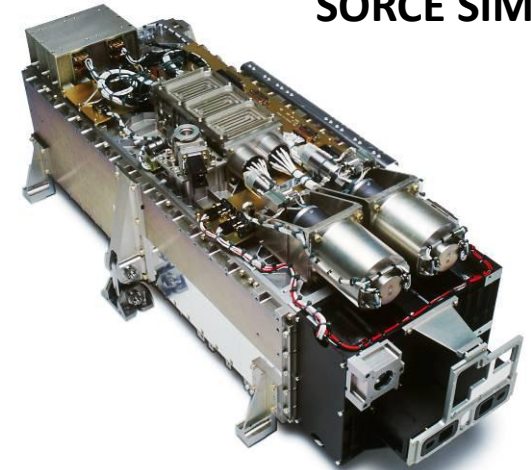
✓ **Improve noise characteristics of ESR and photodiode detectors to meet measurement precision requirement**

- **Improved ESR thermal & electrical design**
- **Larger dynamic range integrating ADC's (21 bits)**

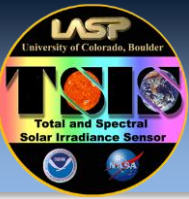
✓ **Improve absolute accuracy pre-launch calibration**

- **NIST SI-traceable Unit and Instrument level pre-launch spectral calibrations (SIMRF-SIRCUS)**

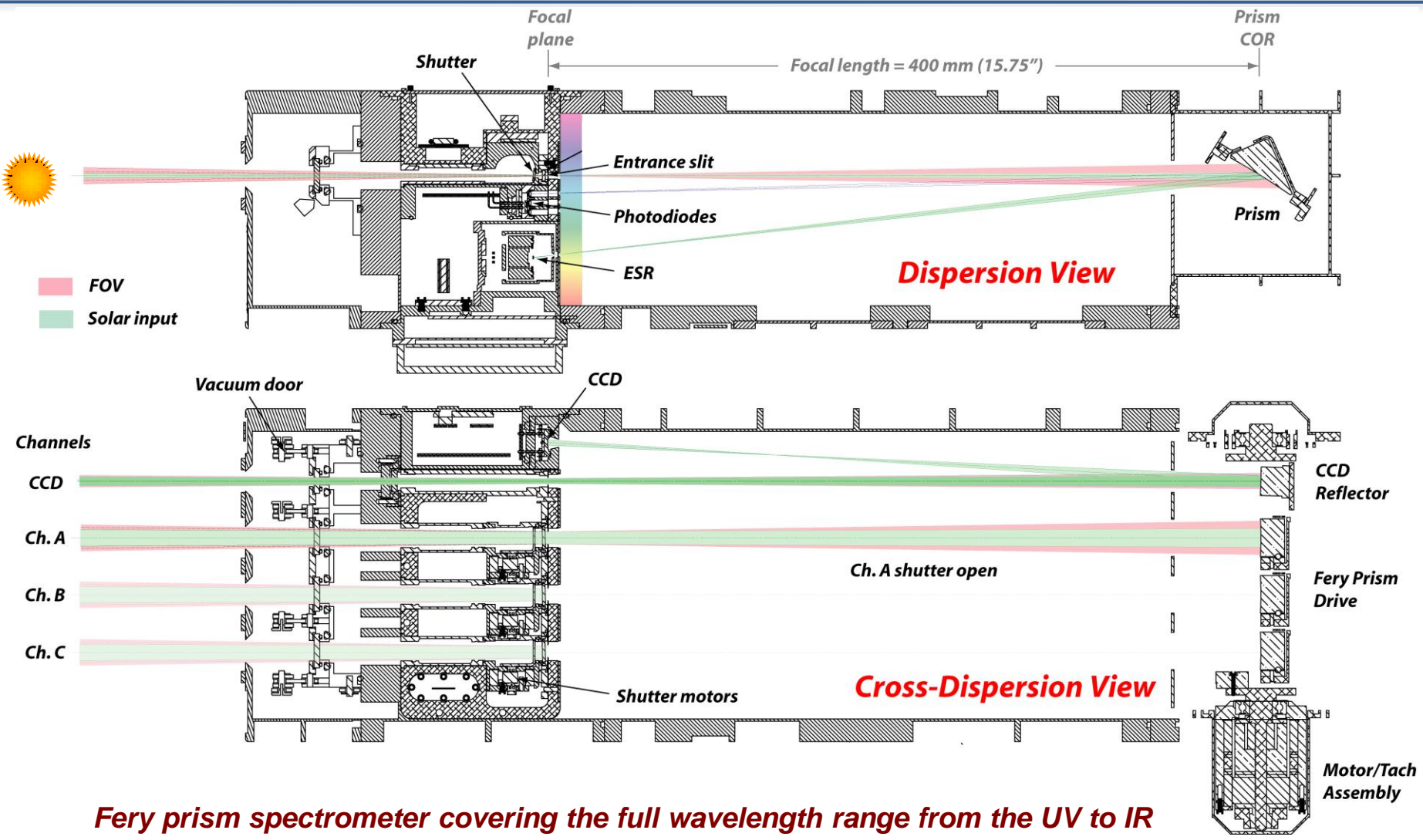
**SORCE SIM**



**TSIS SIM**



# TSIS SIM Design Overview

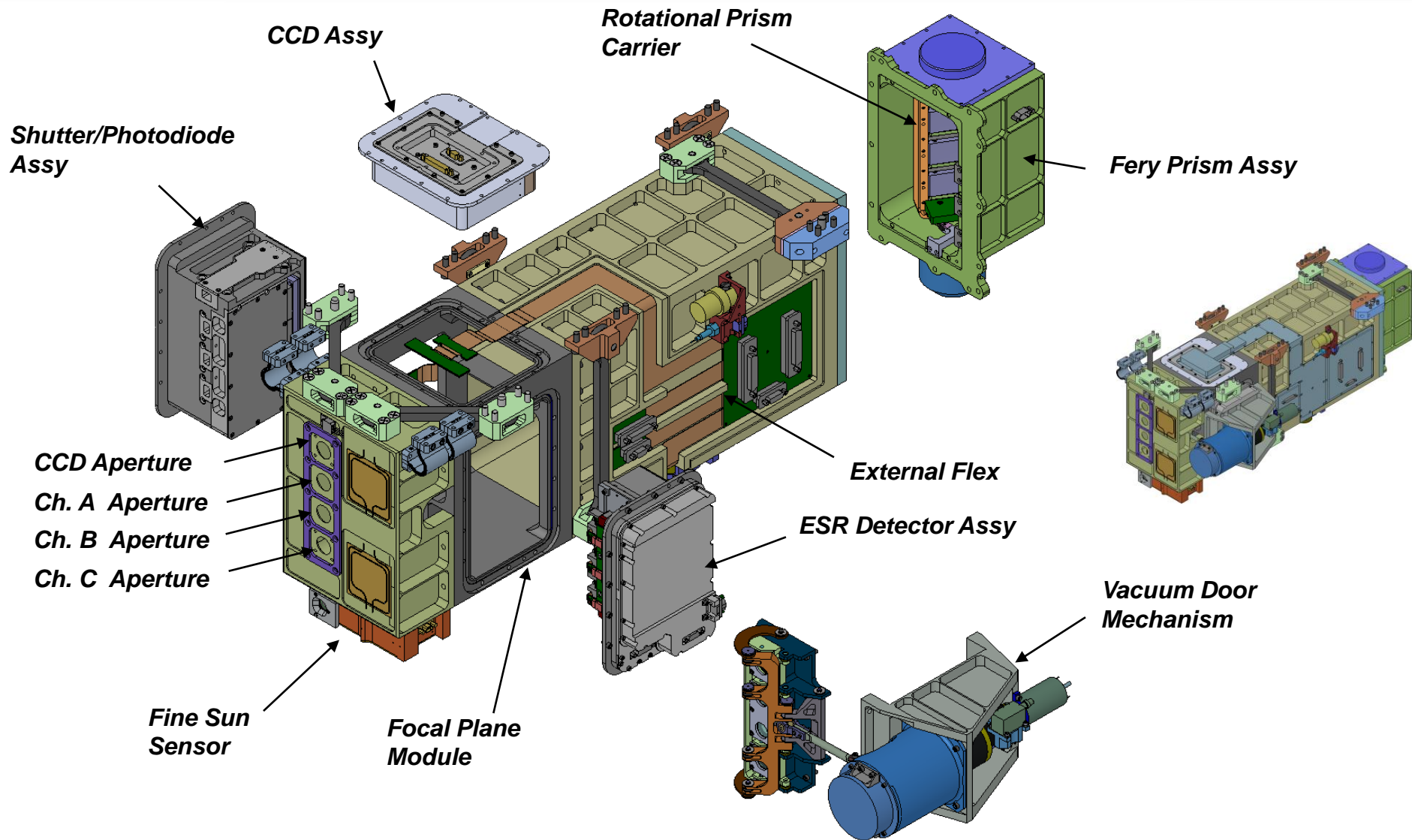


**Fery prism spectrometer covering the full wavelength range from the UV to IR using only one optical element for spectral dispersion and image quality**



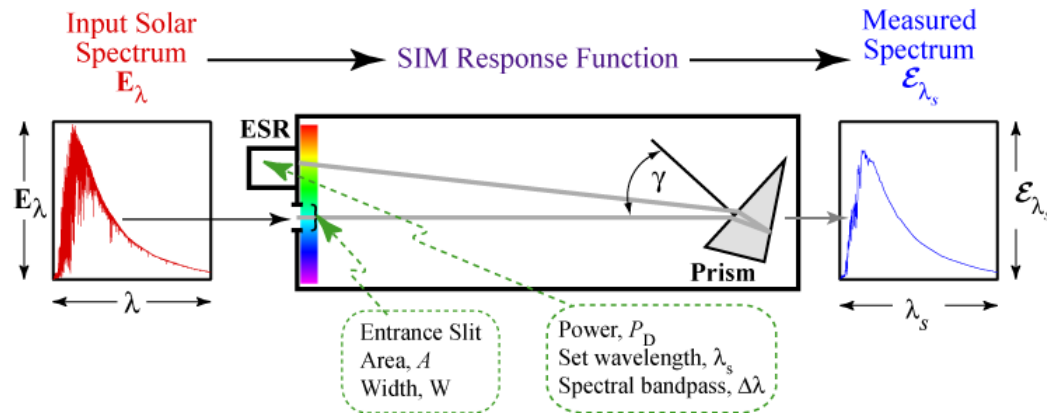


# Overview of TSIS SIM





# Measurement Equation Overview



$$\mathcal{E}_{\lambda}(\lambda_s) = \frac{\mathcal{P}_{\text{ESR}}(\lambda_s)}{A_{\text{slit}} \cdot \int \alpha_{\lambda} \cdot T_{\lambda} \cdot \phi_{\lambda} \cdot S(\lambda, \lambda_s) d\lambda}$$

← ESR detected power

↑

Aperture area

↑

ESR Absorptance

↑

prism transmission

↑

Slit diffraction

↑

Spectral Transfer function





# TSIS SIM Calibration Flow

## Component-Level Calibrations:

### Calibrations Common with TIM:

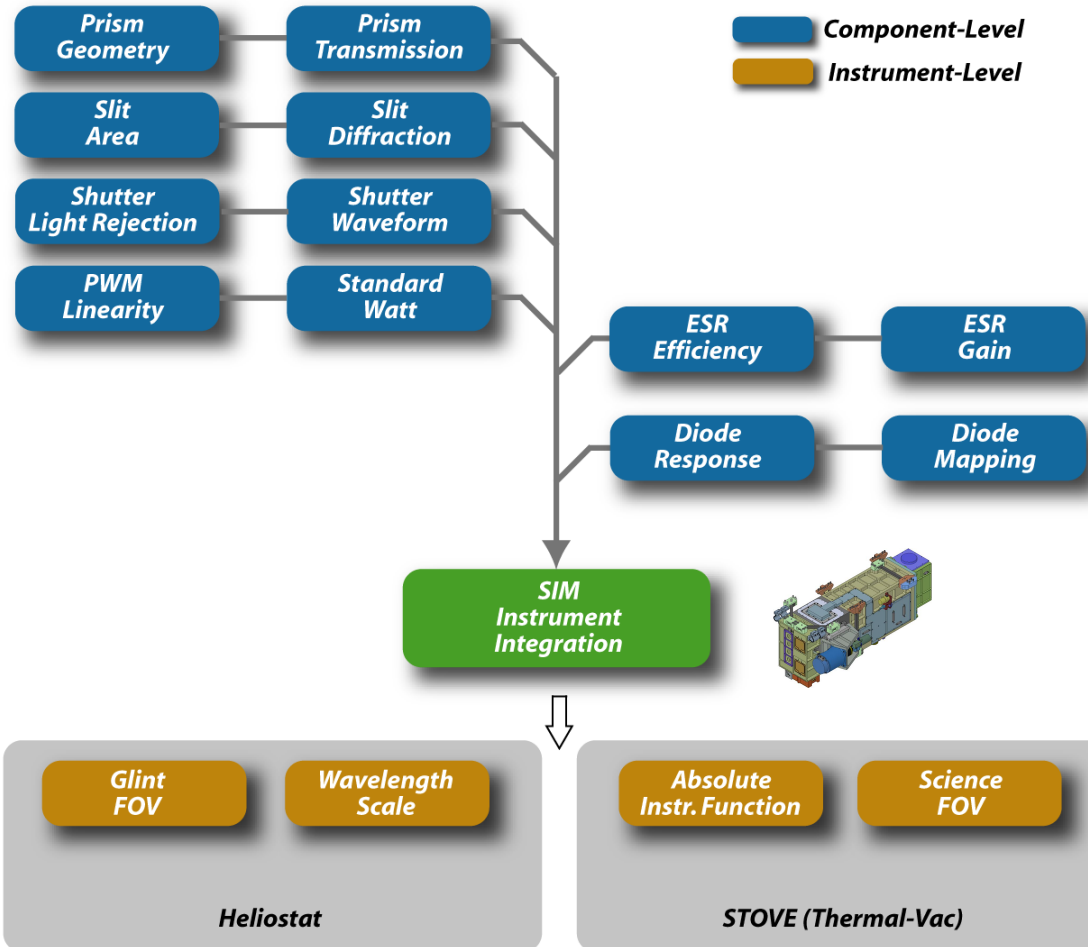
- Slit Area
- Slit Diffraction
- Standard Watt
- Pulse-Width Modulation Linearity
- Shutter Waveform
- Servo Gain

### Unique SIM Calibrations:

- Prism Geometry
- Prism Transmission
- ESR Efficiency
- Photodiode Sensitivity

## Instrument-Level Calibrations:

- Glint Field of View
- Wavelength Scale
- Absolute Instrument Function Area
- Scattered Light
- Science Field of View
- Servo Gain, Nonequivalence, Noise



**Calibration and characterization follows a measurement equation approach at the unit-level for full validation of end-to-end performance at the instrument-level**

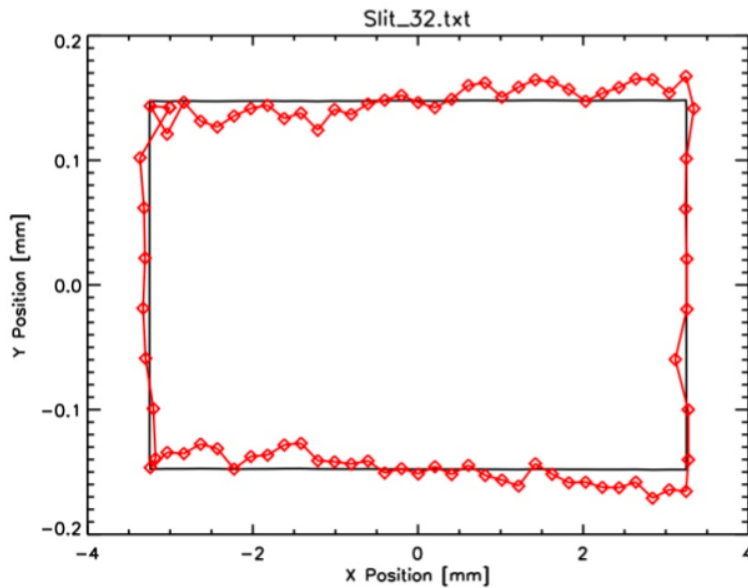


# TSIS SIM Calibration Error Budget

S/C  
Component-Level  
Instrument-Level

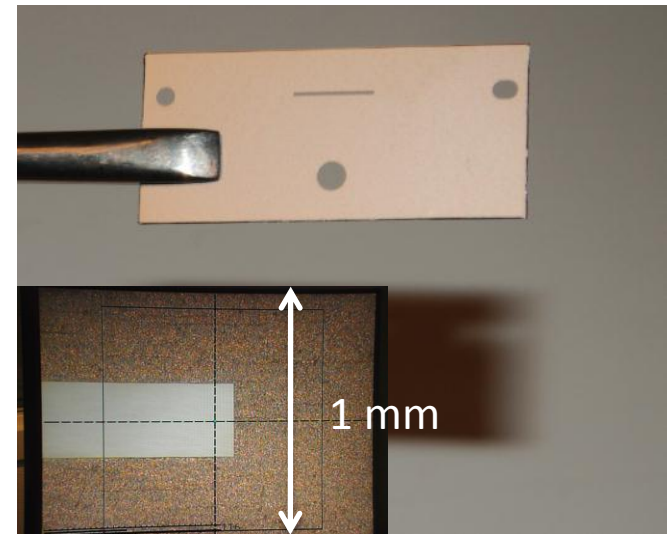
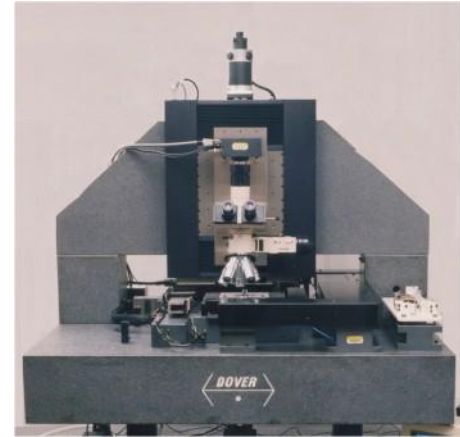
Measurement Correction	Origin	Value (ppm)	1 $\sigma$ (ppm)	Status
Distance to Sun, Earth & S/C	<i>Analysis</i>	33,537	0.1	
Doppler Velocity	<i>Analysis</i>	43	1	
Pointing	<i>Analysis</i>	0	100	
Shutter Waveform	<i>Component</i>	100	10	
Slit Area	<i>Component</i>	1,000,000	500	
Diffraction	<i>Component</i>	5,000-62,000	500	
Prism Transmittance	<i>Component</i>	230,000-450,000	1,000	
ESR Efficiency	<i>Component</i>	1,000,000	1,500	
Standard Volt + DAC	<i>Component</i>	1,000,000	50	
Pulse Width Linearity	<i>Component</i>	0	50	
Standard Ohm + Leads	<i>Component</i>	1,000,000	50	
Instrument Function Area	<i>Instrument</i>	1,000,000	1,000	
Wavelength ( $\Delta\lambda/\lambda = 150$ ppm)	<i>Instrument</i>	1,000,000	1,000	
Non-Equivalence, $Z_H/Z_R - 1$	<i>Instrument</i>	2,000	100	
Servo Gain	<i>Instrument</i>	2,000	100	
Dark Signal	<i>Instrument</i>	0	100	
Scattered Light	<i>Instrument</i>	0	200	
Noise	<i>Instrument</i>	-	100	
<b>Combined Rel. Std. Uncertainty</b>			<b>2418</b>	

# Entrance Slit Area (NIST)



## Non-contact video microscopy technique

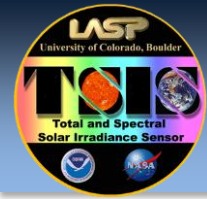
(J. Fowler & M. Litorja, *Metrologia*, **40** (2003) S9-S12)



Slit	Rectangle Fit			Polygon Fit		
	Area (mm <sup>2</sup> )	Std. Unc (k=1)	Rel. Unc (ppm)	Area (mm <sup>2</sup> )	Std. Unc (k=1)	Rel. Unc (ppm)
<b>A</b>	1.92756	1.8 (10) <sup>-4</sup>	95	1.92756	1.8 (10) <sup>-4</sup>	95
<b>B</b>	1.94450	3.1 (10) <sup>-4</sup>	154	1.94447	2.0 (10) <sup>-4</sup>	102
<b>C</b>	1.94416	3.3 (10) <sup>-4</sup>	162	1.94414	2.4 (10) <sup>-4</sup>	125

**Calibration Budget Allocation 500 ppm - still evaluating corner issues**

**Component-Level**



# Wavelength Dependence: NIST SIRCUS

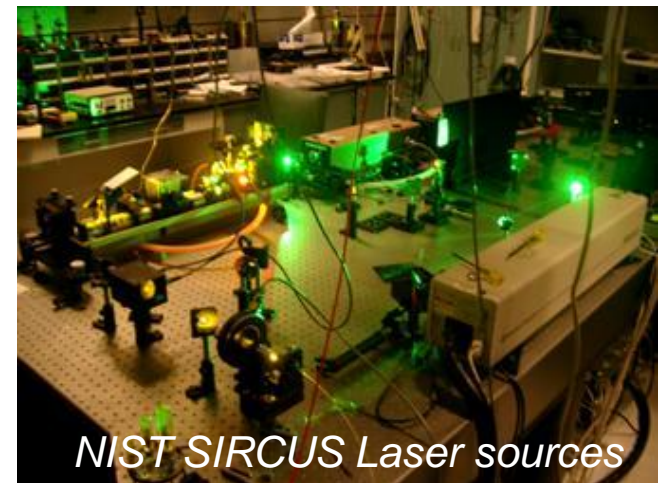
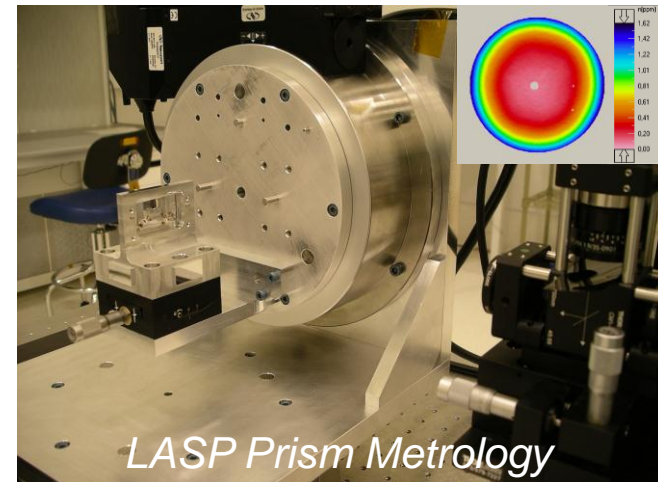
- We need to carefully measure the wavelength dependence of several items in the uncertainty budget:

- Prism Transmission
- ESR Efficiency
- Photodiode Sensitivity
- Wavelength Scales
- Absolute Instrument Function Area

$$\mathcal{E}_\lambda(\lambda_s) = \frac{\mathcal{P}_{\text{ESR}}(\lambda_s)}{A_{\text{slit}} \cdot \int \alpha_\lambda \cdot T_\lambda \cdot \phi_\lambda \cdot S(\lambda, \lambda_s) d\lambda}$$

- This wavelength dependence is always a smoothly-varying function, nonetheless calibrations must be performed at many wavelengths
- We utilize NIST travelling SIRCUS\* (Spectral Irradiance and Radiance Responsivity Calibrations using Uniform Sources) to generate laser light across the SIM spectrum of 210-2400 nm

\*S. Brown, et al. *Applied Optics*, 45, (2006), 8218-8237

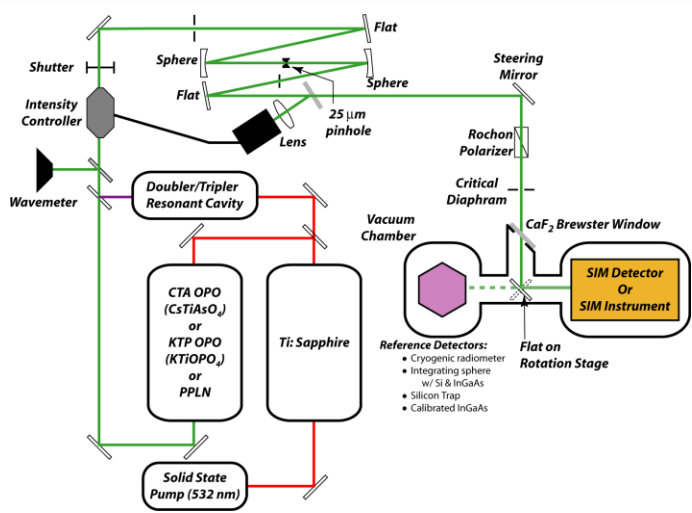


Component-Level  
 Instrument-Level

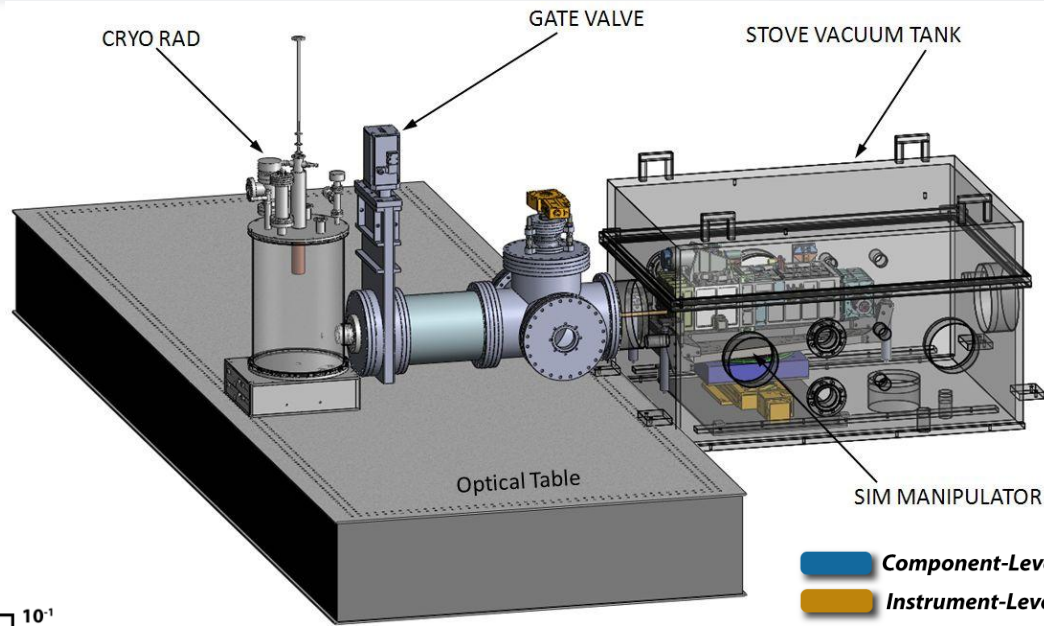
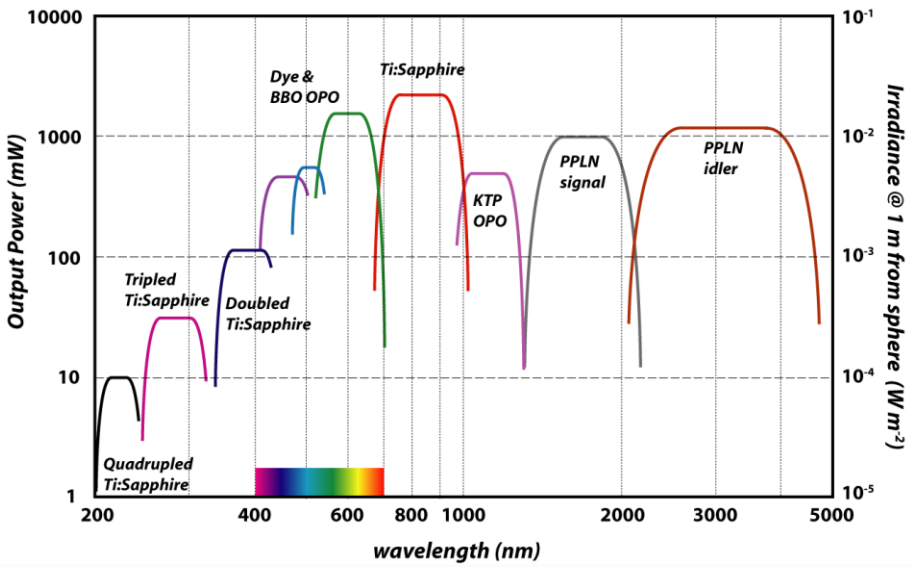




# SIM Radiometric Facility (SIMRF) (LASP & NIST)



## SIRCUS laser spectral coverage



**This comprehensive calibration facility provides irradiance calibration of the SIM ESR & photodiodes and the full SIM over the operation wavelength range of the SIM instrument.**

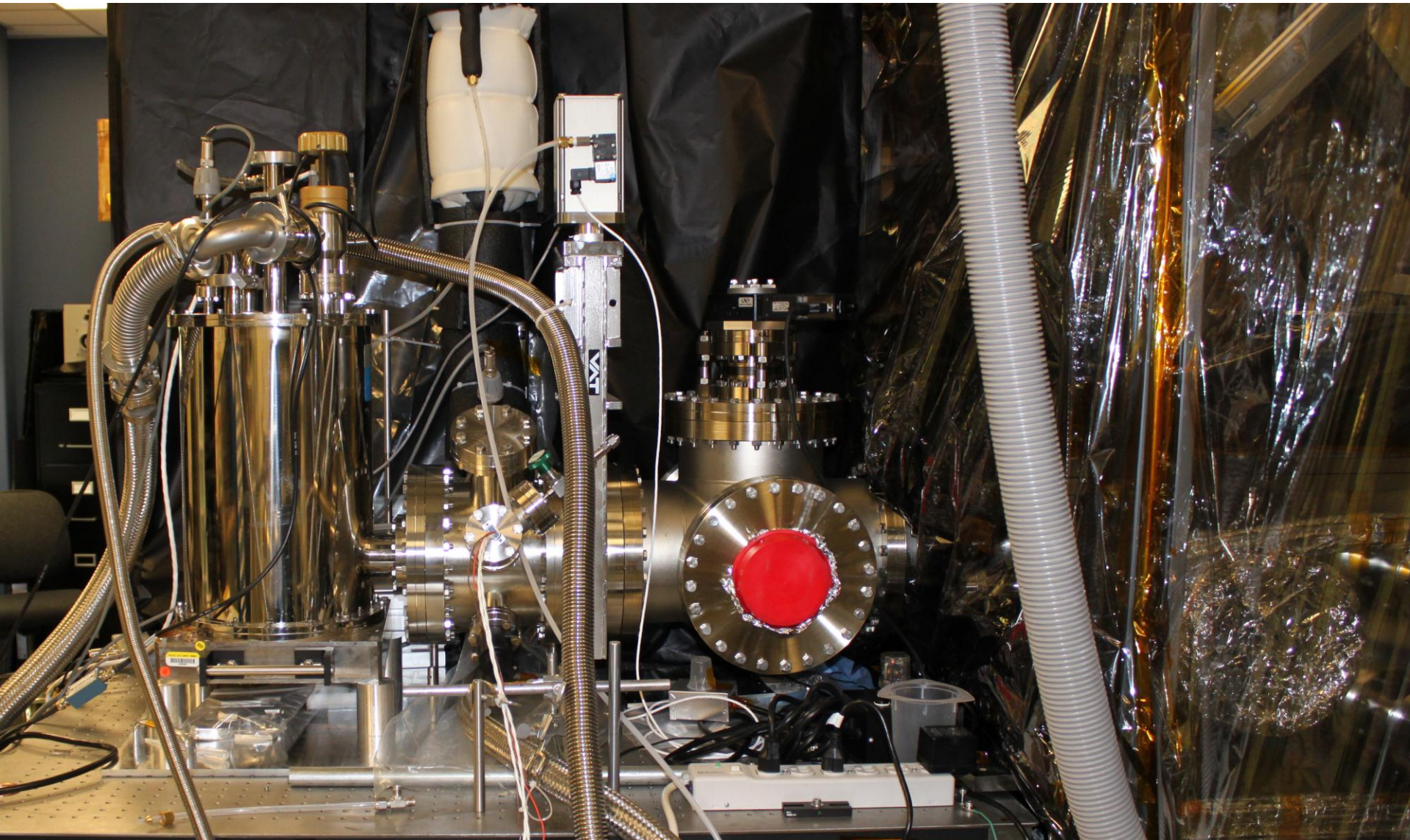
### Main components of SIMRF:

- **STOVE vacuum tank**
- **Manipulator for SIM and ESR subassembly**
- **Steering mirror vacuum housing**
- **L-1 Cryogenic radiometer (NIST)**
- **SIRCUS lasers (NIST)**





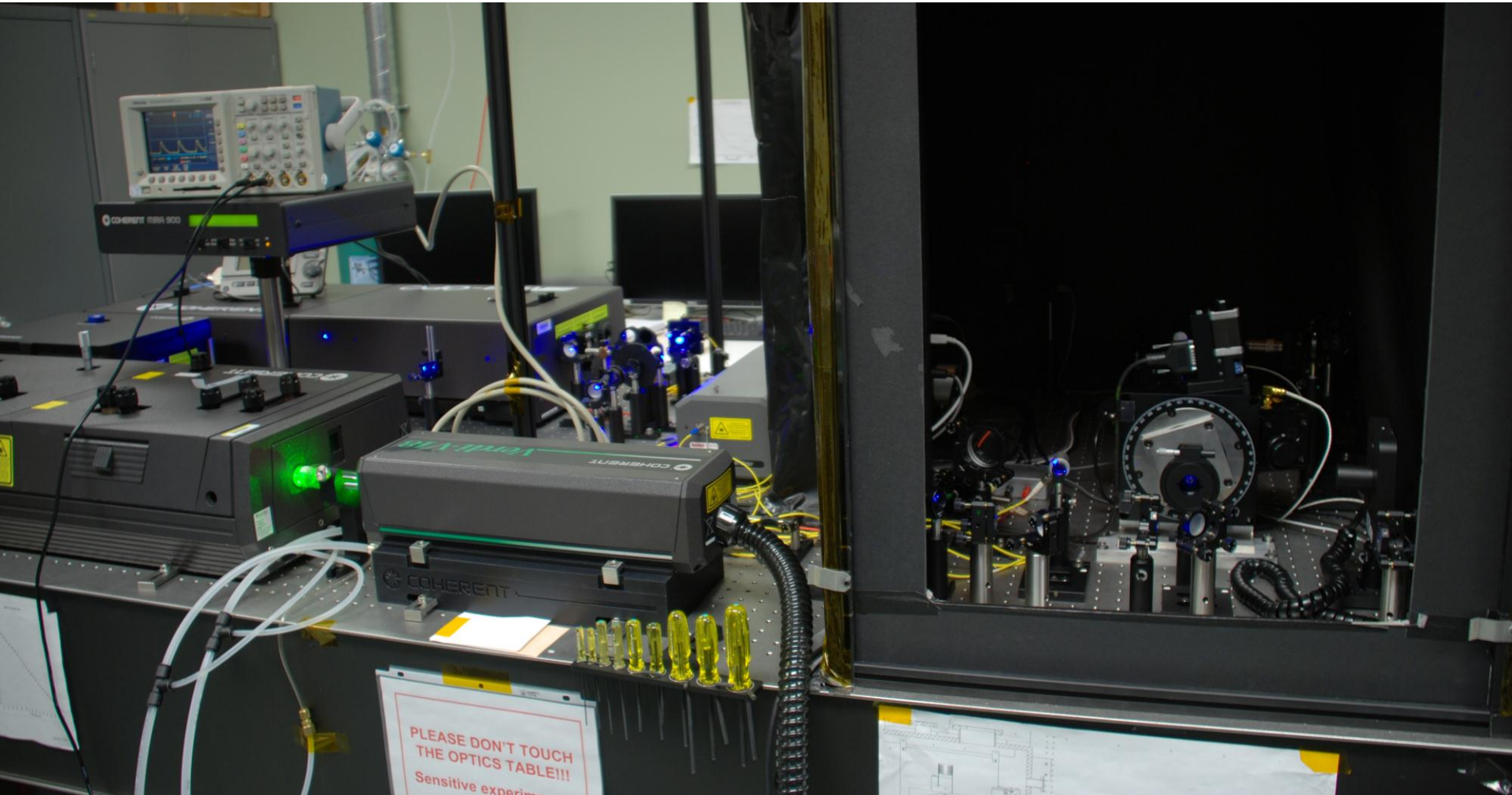
# NIST L1 Cryogenic Radiometer





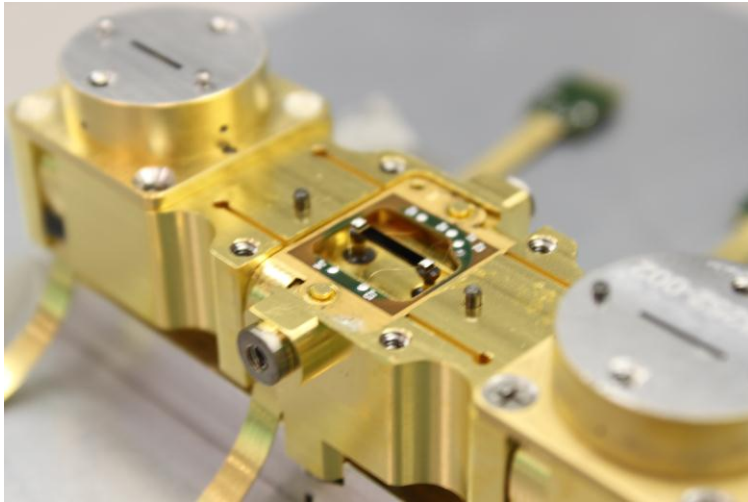


# NIST Travelling SIRCUS

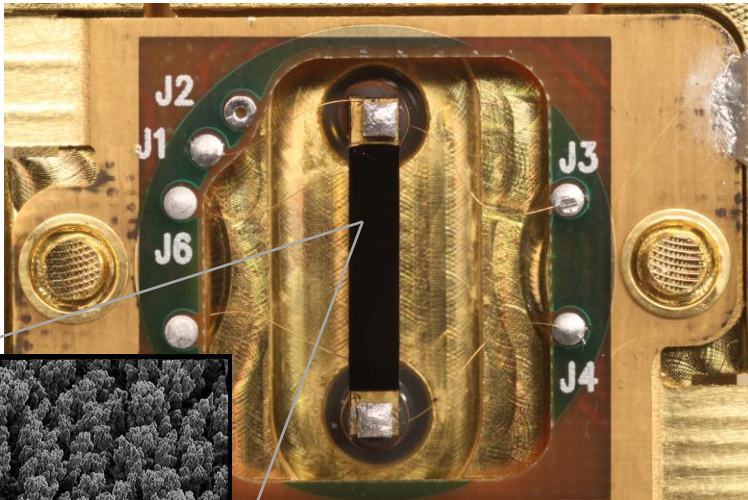




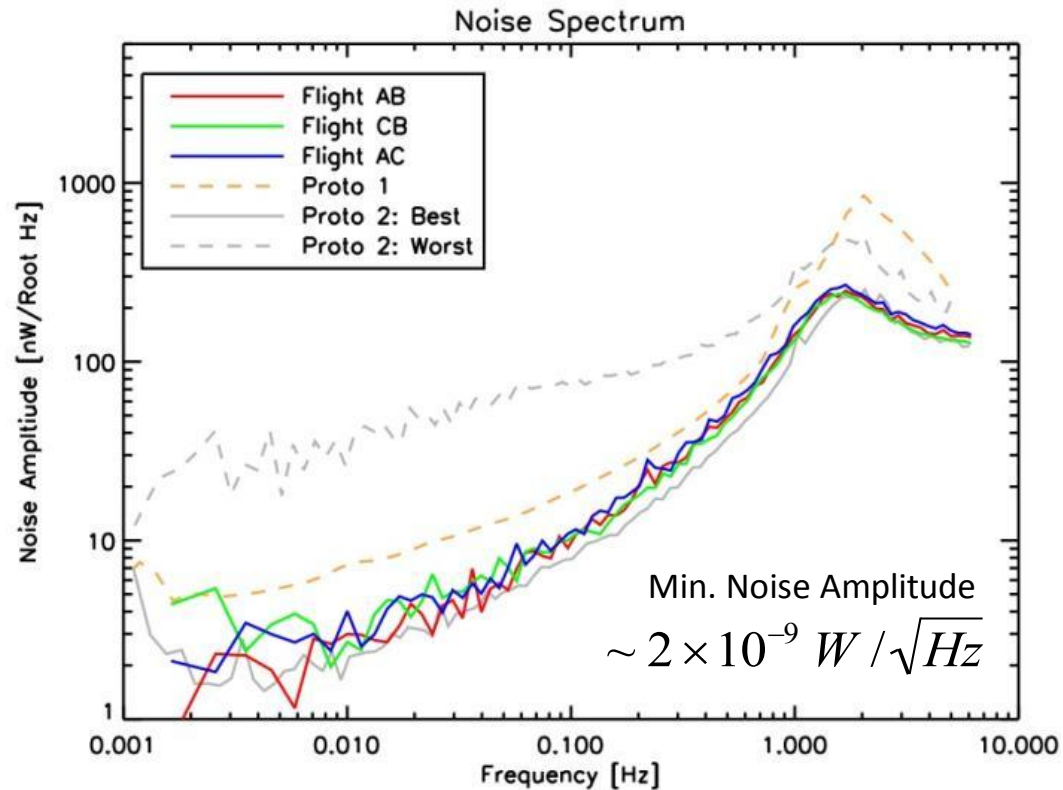
# TSIS SIM ESR



**The ESR provides for a NIST traceable, space-qualified absolute calibration transfer detector**



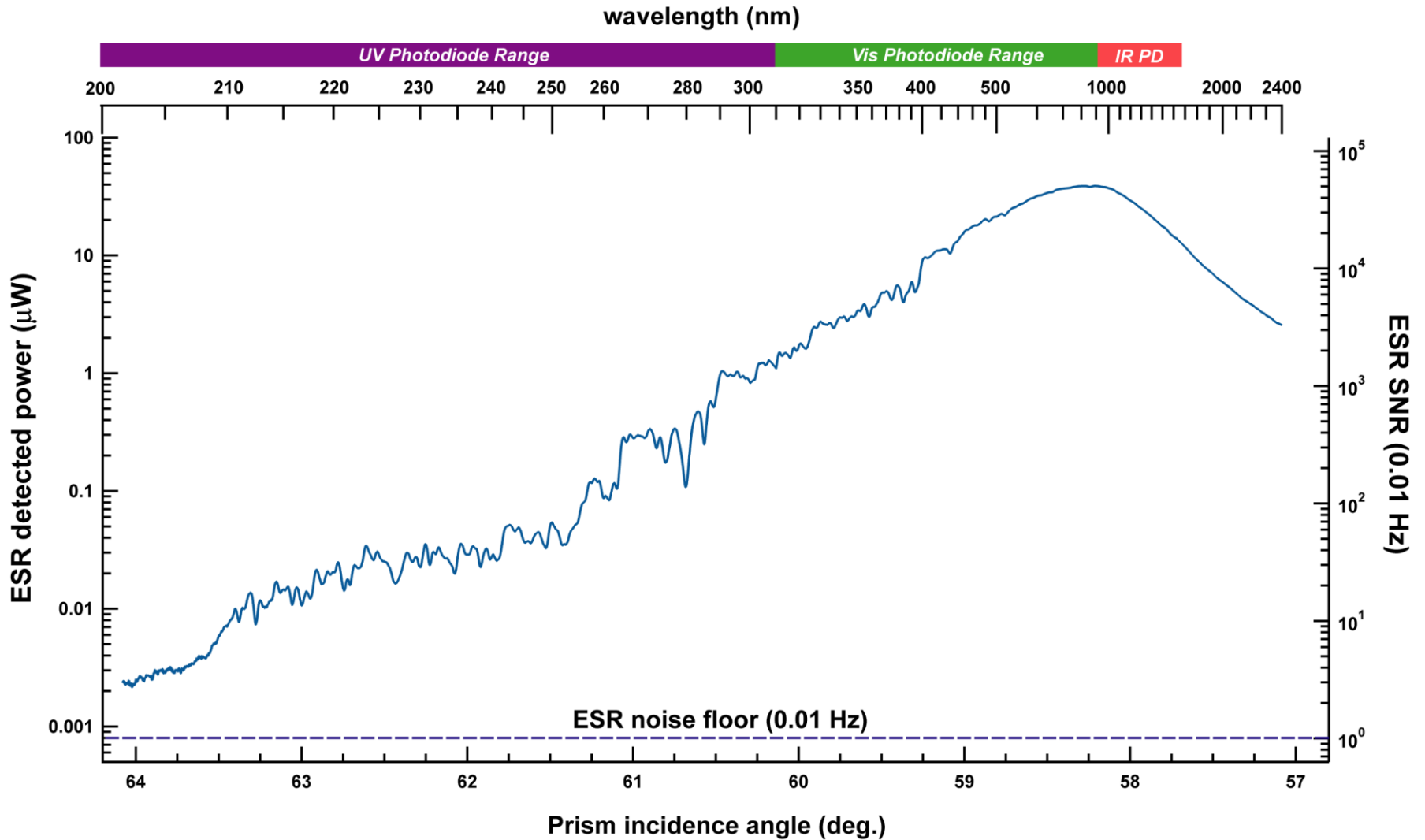
*NiP Black on CVD diamond substrate*



**Component-Level**



# ESR Performance Meets Requirements for Solar Spectral Power

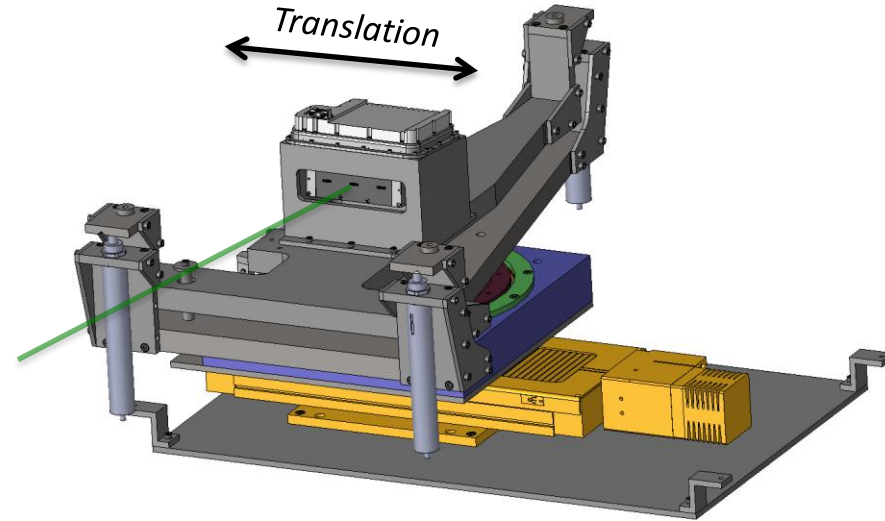
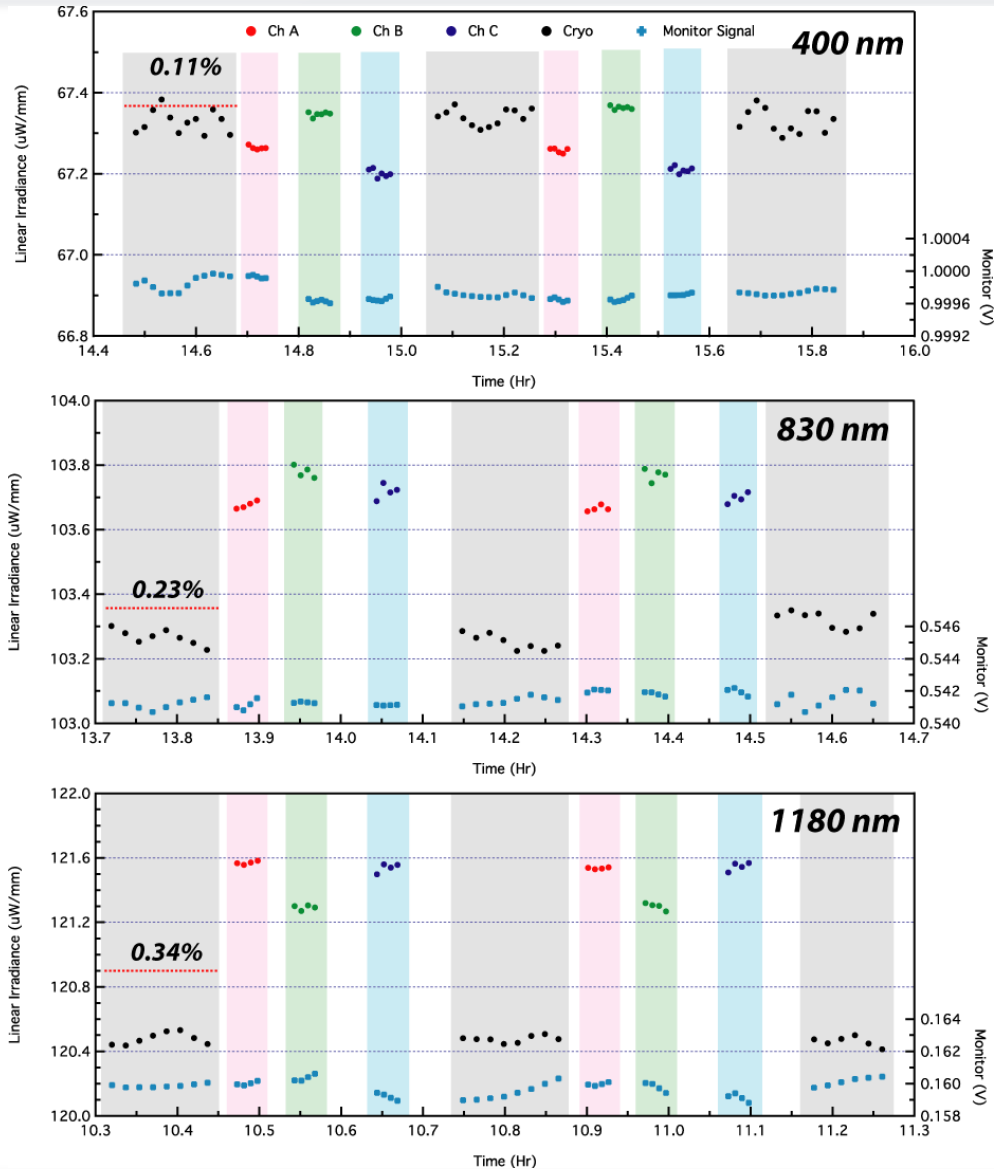


**Instrument-Level**

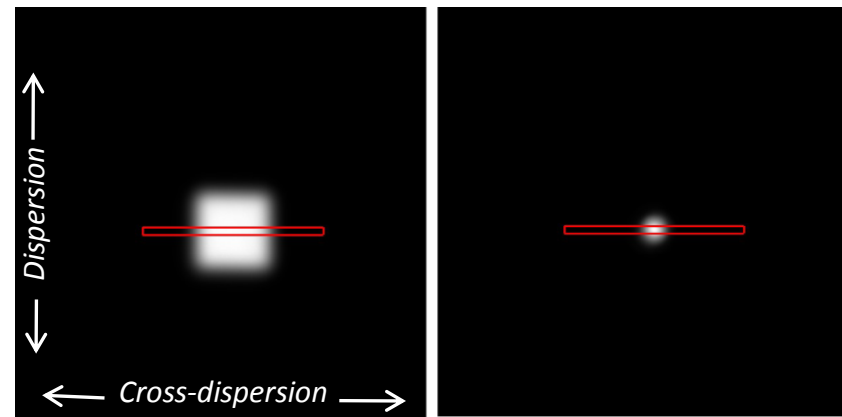




# ESR – Cryo Intercomparison



**SIMRF Laser Scan Image**



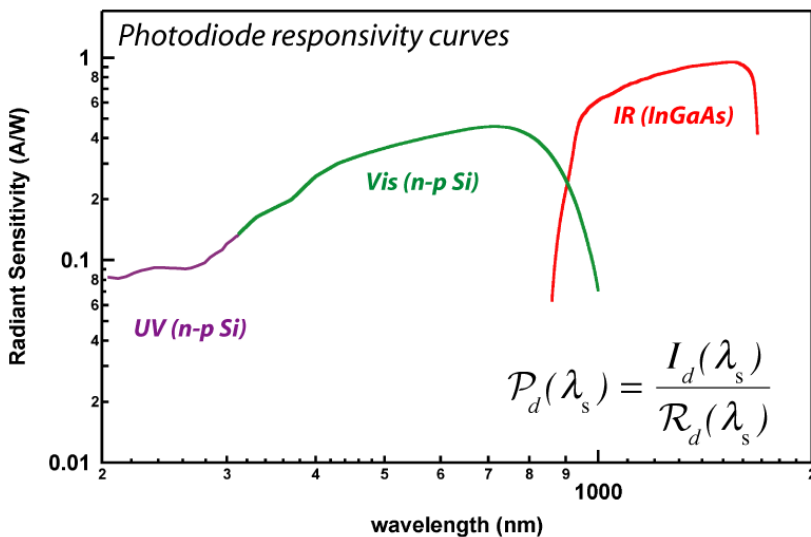
**Component-Level**





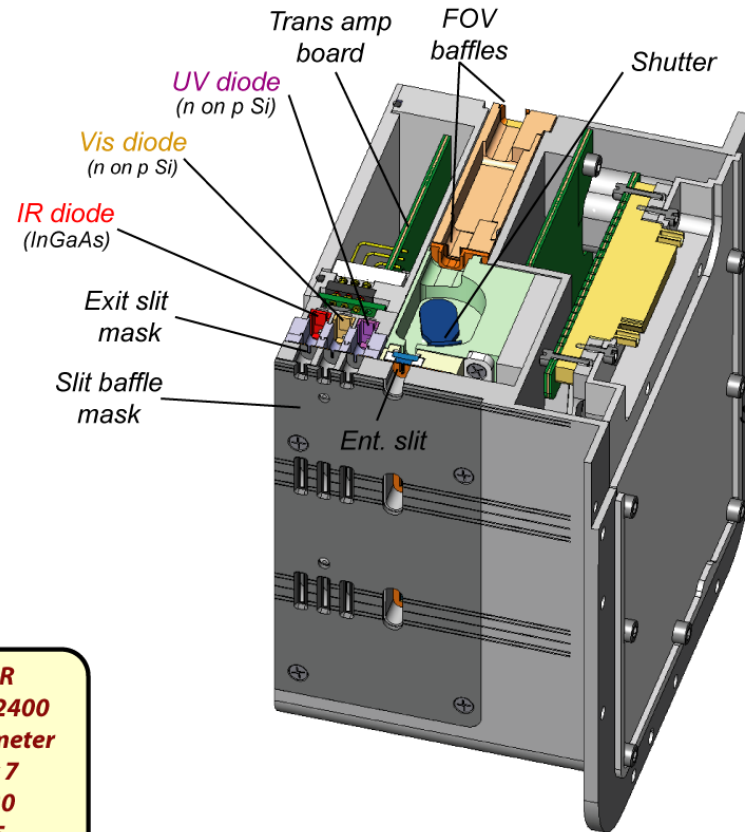
# SIM Photodiode Detectors

**SIM incorporates 3 photodiode detectors to cover the UV through near IR. These fast detectors are calibrated for radiant responsivity on-orbit by periodic scans with the absolute ESR detector**



Detector	UV	Vis	IR	ESR
$\lambda$ range (nm)	200-315	310-950	900-1655	200-2400
Material	n-p Si	n-p Si	InGaAs	bolometer
Active area (mm)	2 x 10	2 x 10	2 x 8	1.5 x 7
Exit slit width (mm)	0.34	0.30	0.30	0.30
FP position (mm)*	10	15	20	45

\*relative to entrance slit

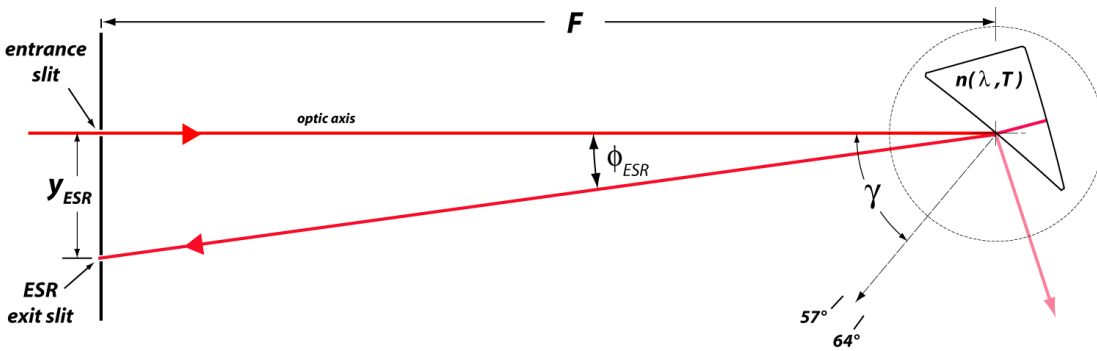


**Shutter / Photodiode Assembly**

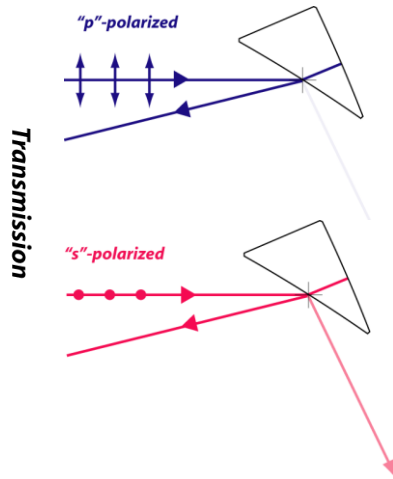
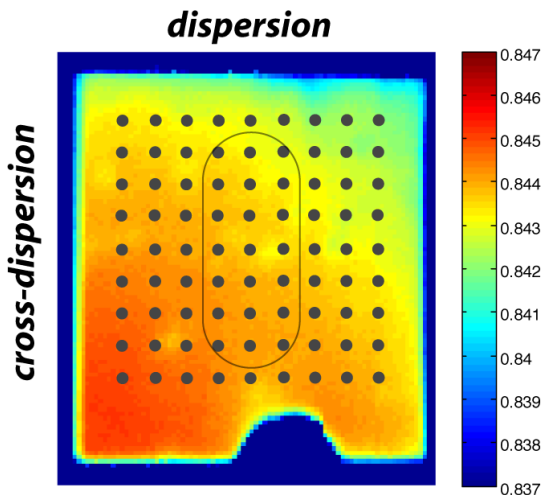


# Full Spatial & Spectral Transmission Mapping

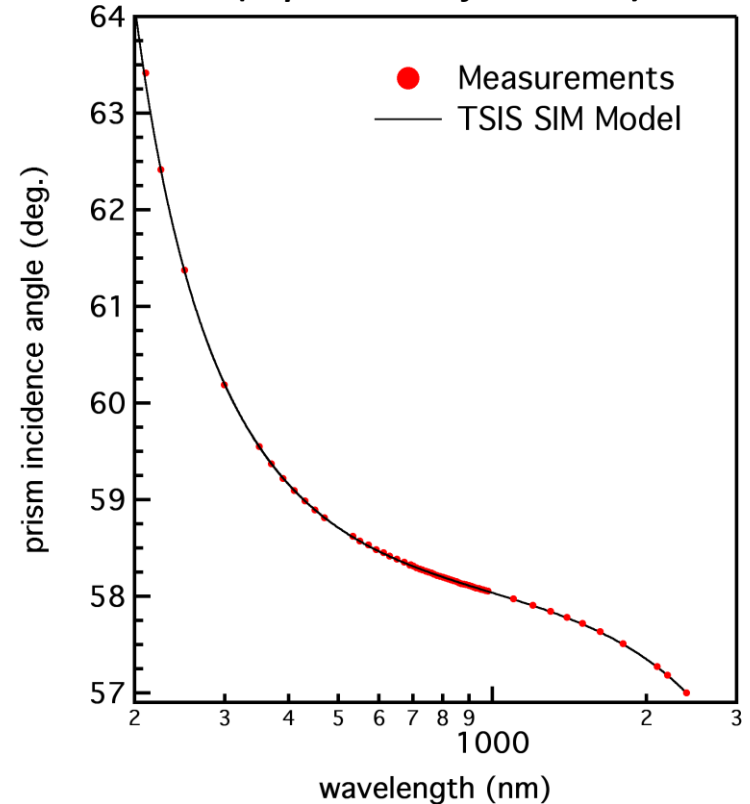
Prism measurement geometry is for ESR optical path  
 Stabilized SIRCUS lasers cover 210 – 2400 nm range



Transmission measured over 10 x 10 grid for both  
*s* and *p*-polarizations



Refraction vs. wavelength  
 (Suprasil 3001 fused silica)

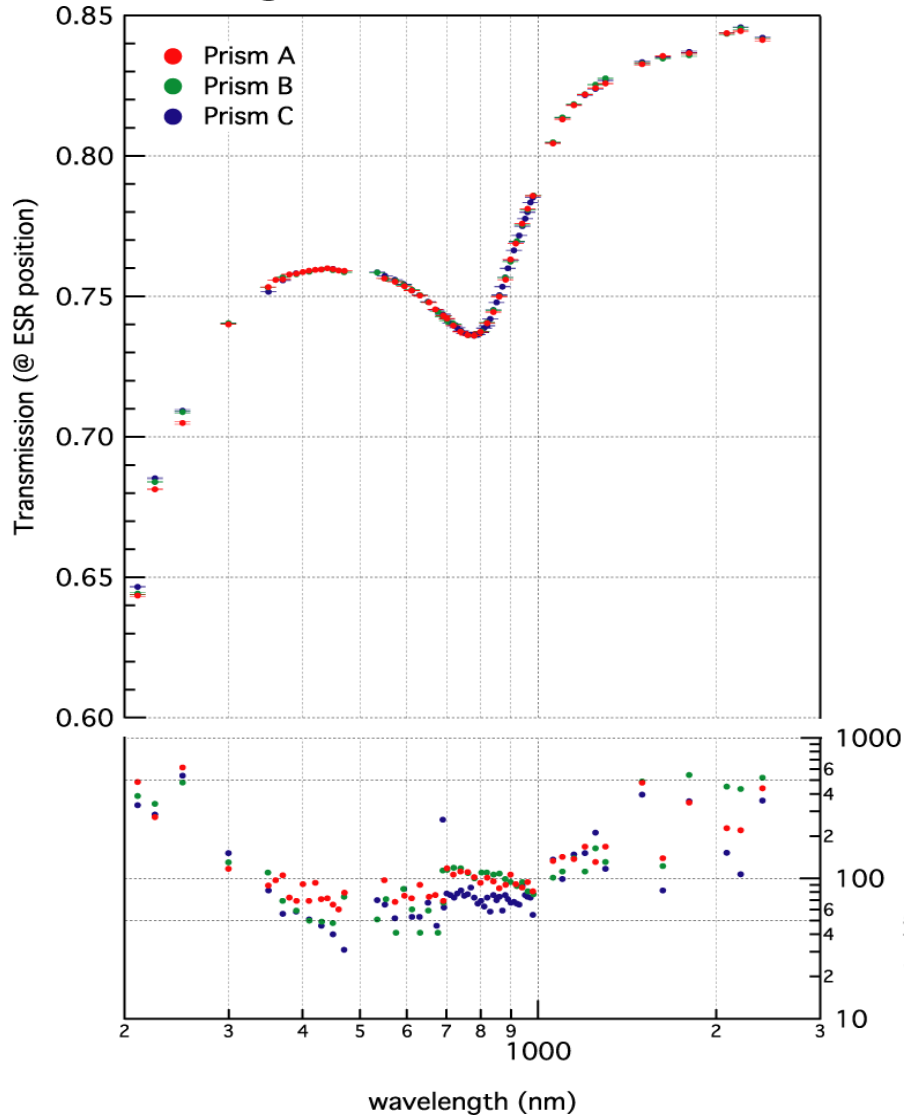


Component-Level

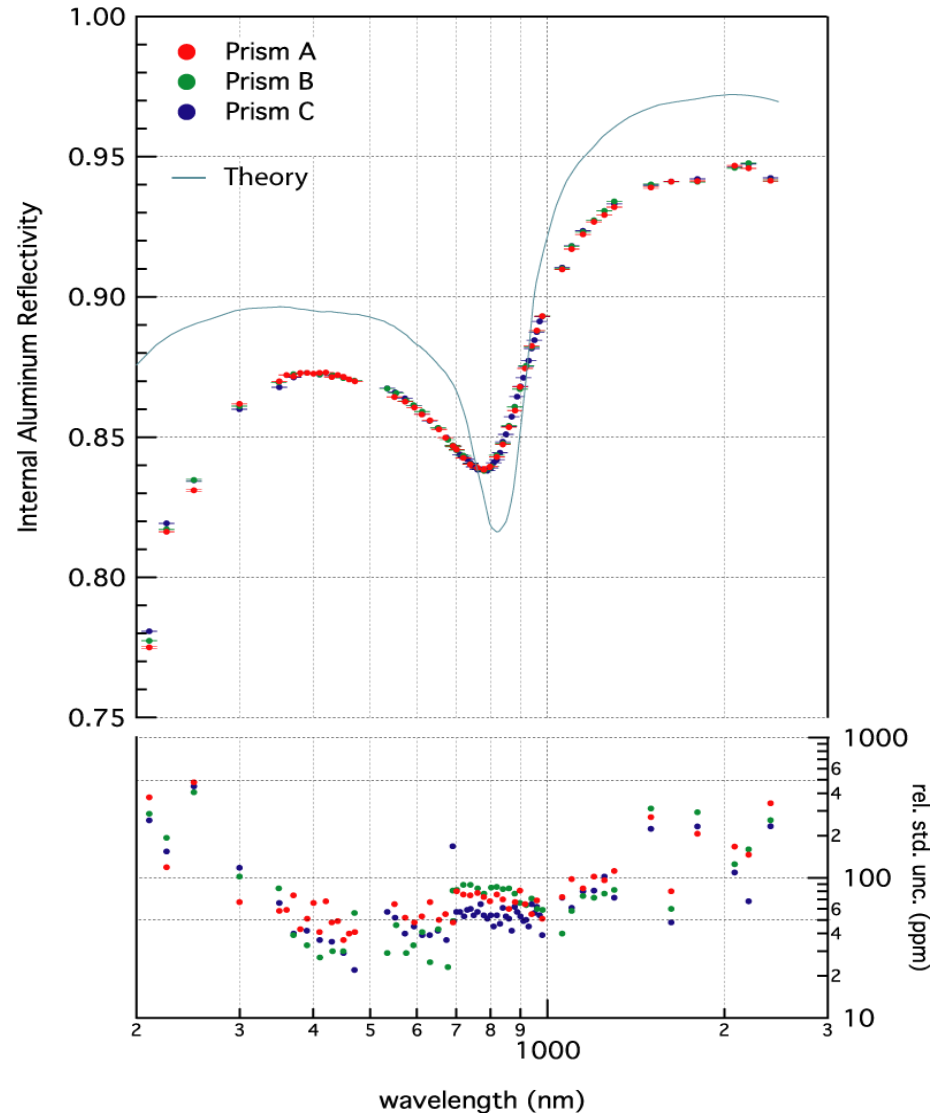


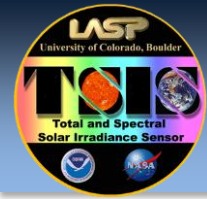
# Prism Transmission

**Average Prism Transmission (@ ESR)**

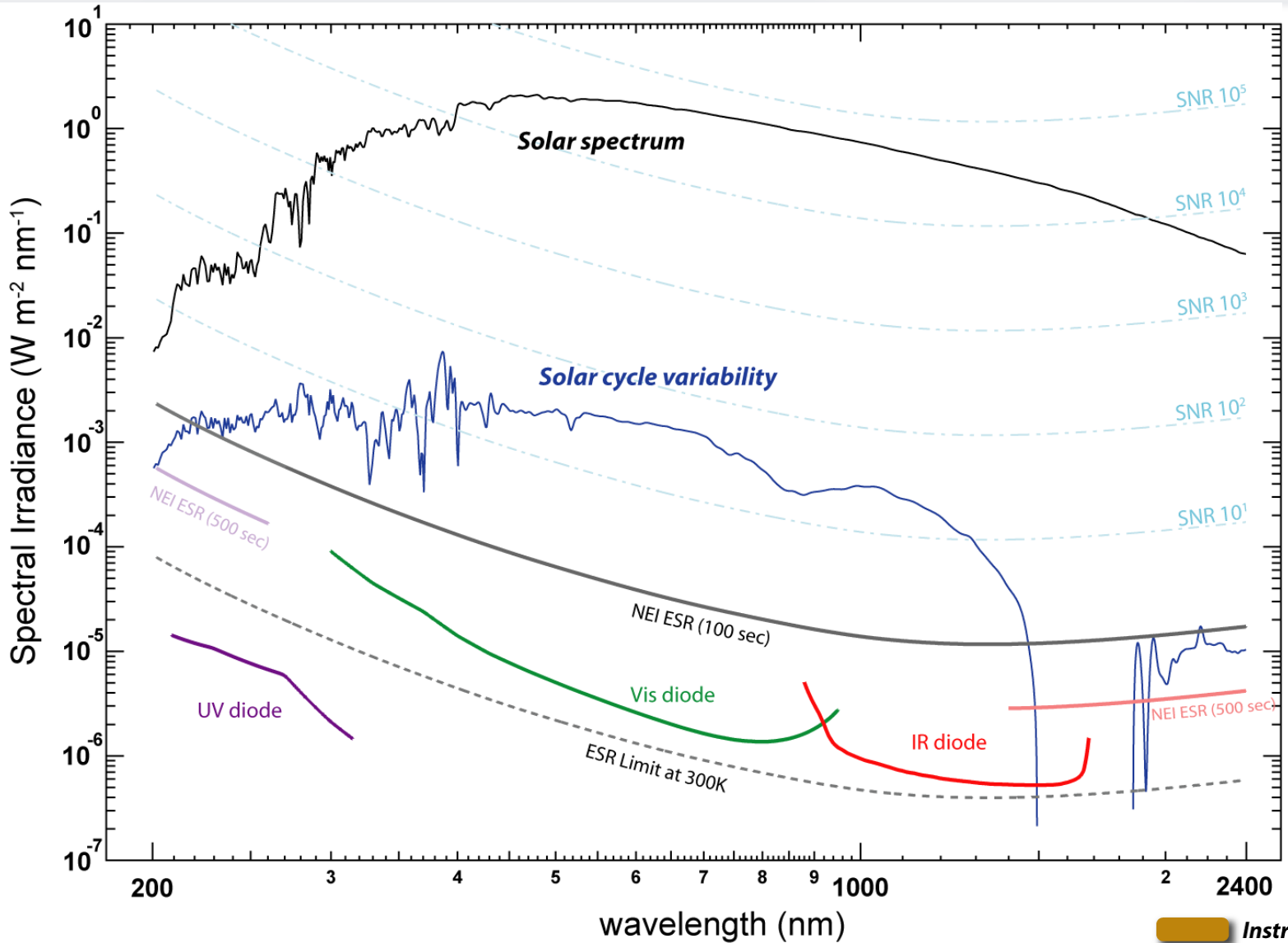


**Internal Al Reflectivity (inc. bulk loss)**





# Spectral Variability and SIM Measurement Capabilities



Instrument-Level



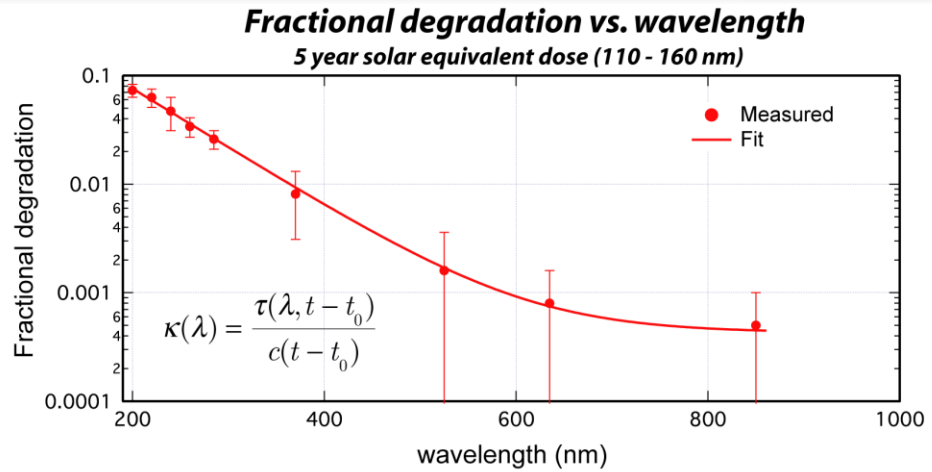
# Exposure Degradation over Mission Life

On-orbit interchannel transmission comparisons track wavelength and exposure time dependent transmission loss

$$I(\lambda, t - t_0) = I(\lambda, t_0) e^{-\tau(\lambda, t - t_0)}$$

$$\tau(\lambda, t) = \kappa(\lambda) \cdot c(t)$$

$\kappa(\lambda)$  evaluated by periodic ESR measurements between separate channels



The degradation correction determined using the Channel A to Channel B ratio data measured twice per month

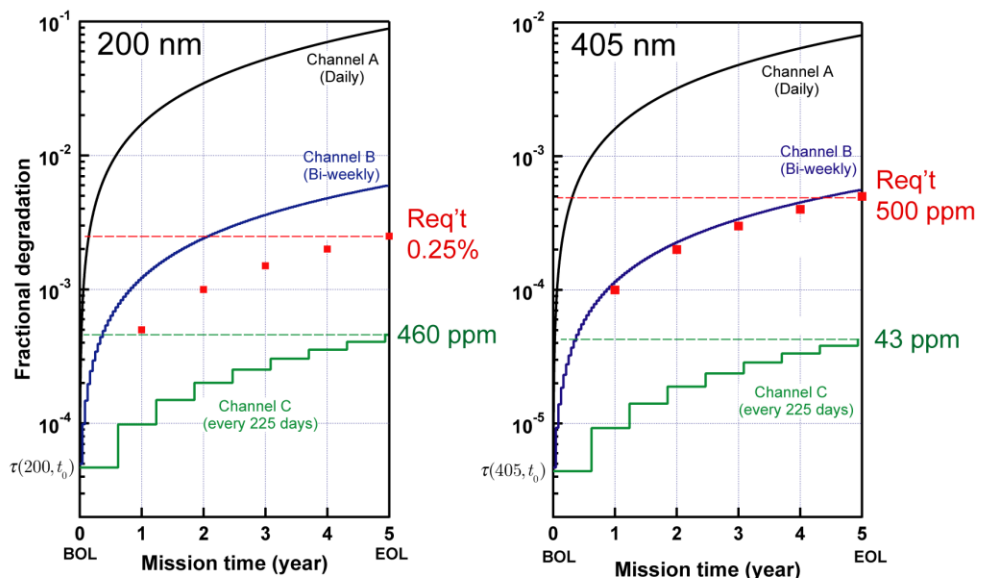
The Channel A to Channel C comparison (~1 per year) verifies the degradation correction

Channel C is to be used infrequently enough so that it can be considered "pristine" (less than 0.01%/year of degradation)

We must be able to measure trends < 0.01% (0.05%) / year in this Channel A to Channel C ratio data for Vis/IR (UV)

This puts a limit on the spectra to spectra repeatability

Fractional degradation vs. exposure time

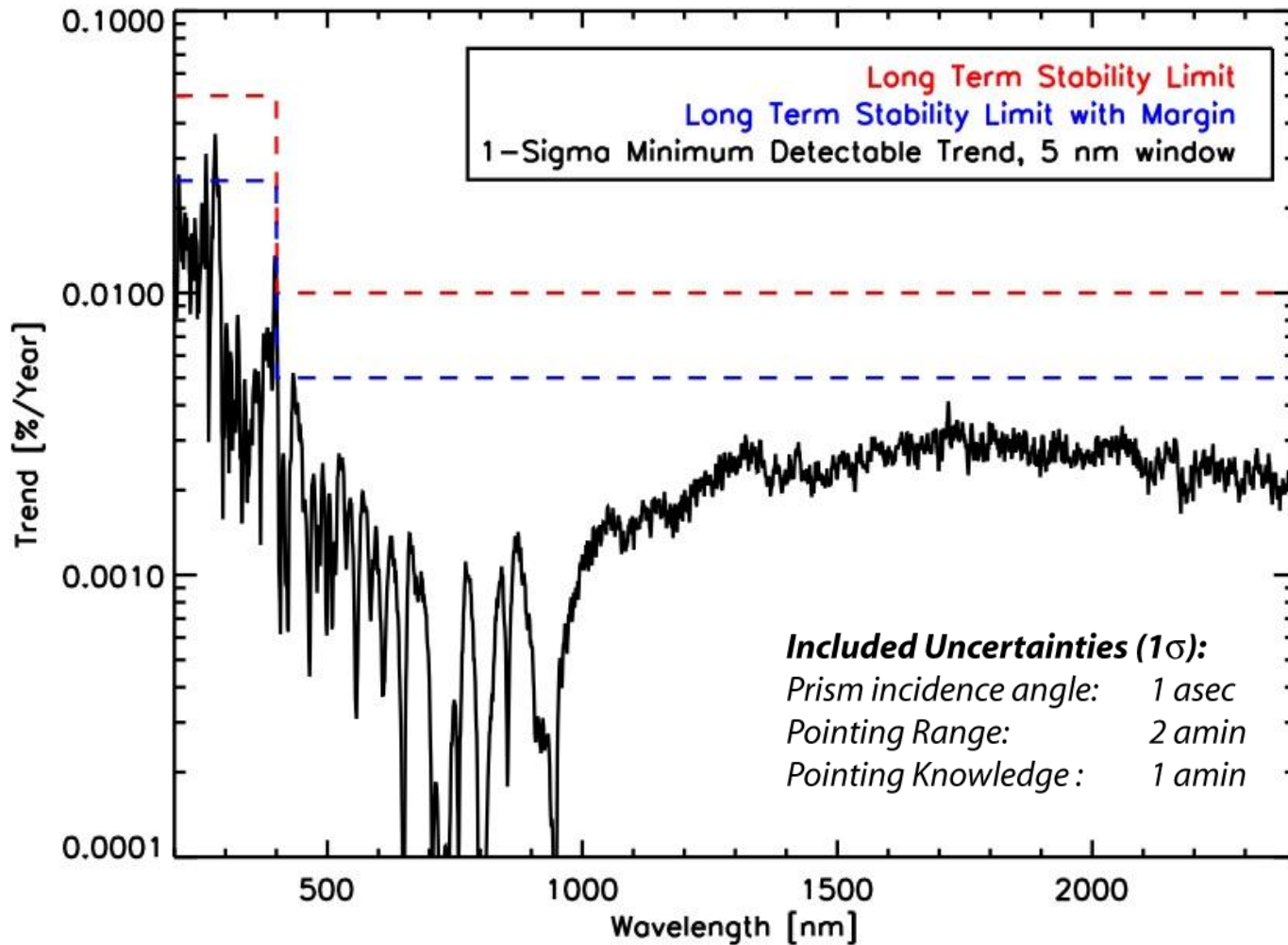






# Long-term Stability Limits on SIM Noise/Repeatability

## Channel A to Channel C Degradation Uncertainty





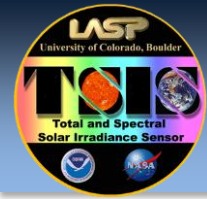
# Summary

***Spectral Solar Irradiance (SSI) is critical to understanding solar variability and its impact on Earth climate***

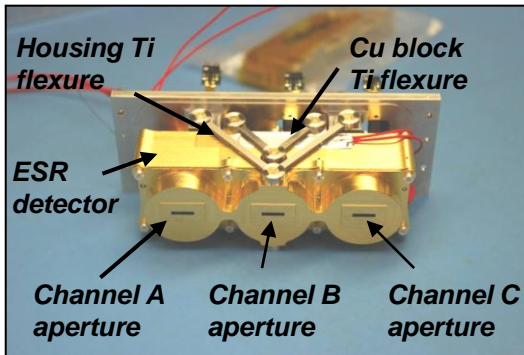
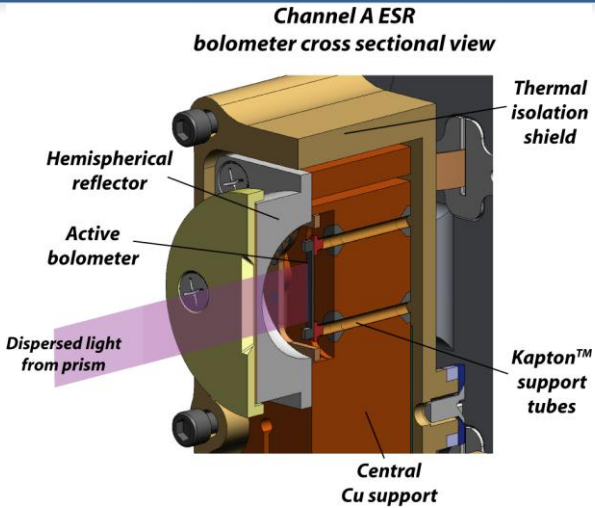
- ***TSIS SIM meets the JPSS measurement requirements for SSI variability, including:***
  - ***High absolute irradiance accuracy ( $\leq 0.25\%$  over full spectrum)***
  - ***High measurement precision ( $< 0.01\%$  relative)***
  - ***On-orbit capability to self-correct long-term drifts and sensitivity changes ( $< 0.05\%$  per year)***
    - ***Channel-to-channel calibrations***
    - ***Direct measurements of optical components***
    - ***Detector-to-detector calibrations***
- ***TSIS SIM significant improvements over SORCE SIM include:***
  - ***Long-term relative stability***
    - ***Improved absolute ESR detector and duty-cycling 3 independent channels provides on-orbit calibration maintenance***
  - ***Measurement accuracy***
    - ***NIST calibration facilities (SIRCUS/POWR) provide SI-traceable pre-launch calibration***



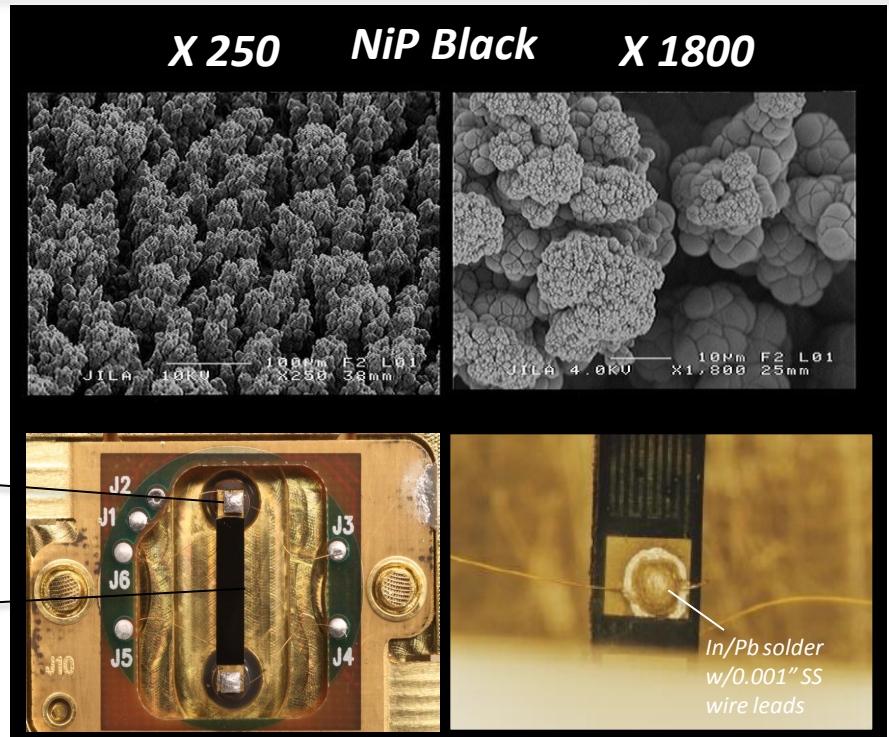
# Backup Slides



# ESR Detector Provides Absolute Power Measurement

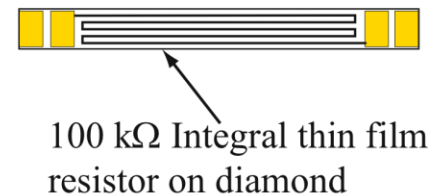
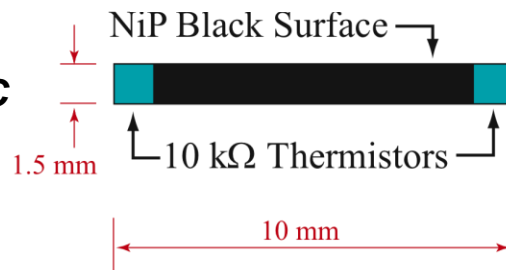


- **Bolometer thermistors part of a precision AC bridge excited at 50 Hz**
- **Incoming solar beam chopped by shutter at 0.01 Hz and signal detected at shutter fundamental**



**Illuminated Face**

**Back Face**



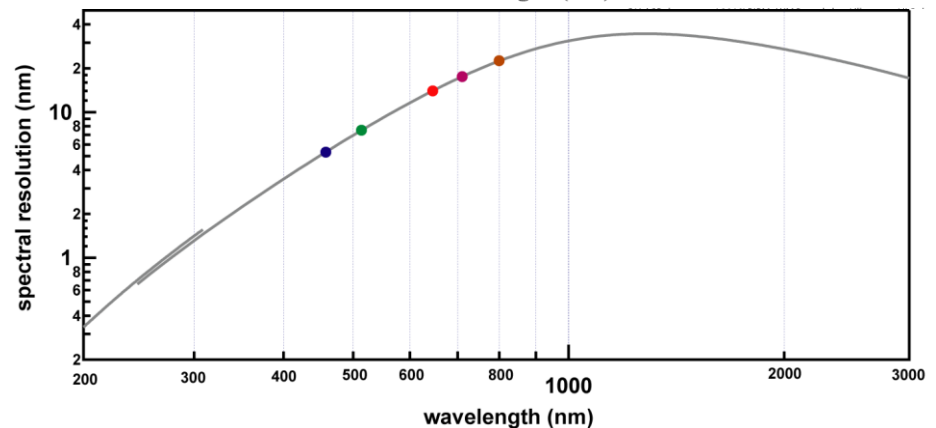
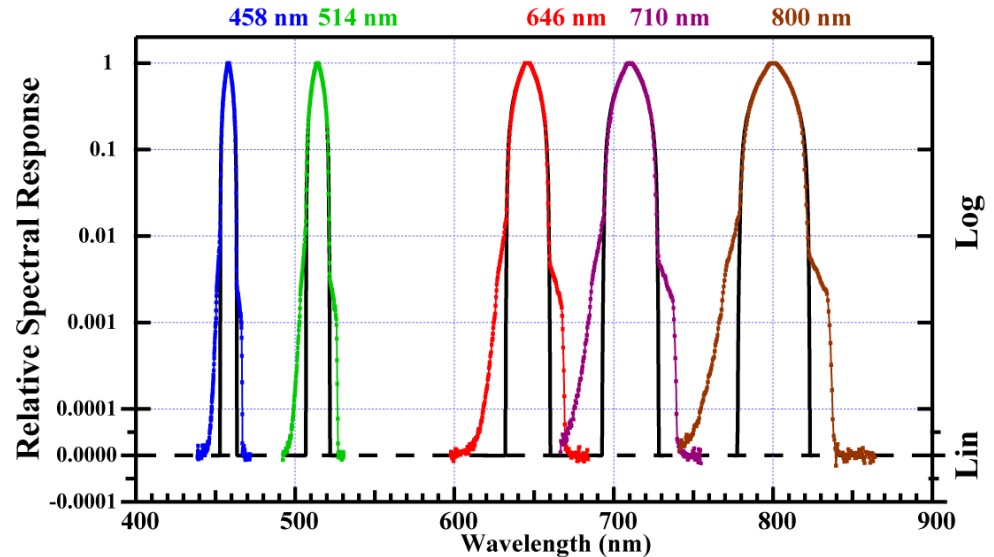
**Component-Level**





# Absolute Instrument Profile

- **Illuminate instrument with a single wavelength and a known irradiance and measure signal vs. prism angle** - Irradiance is measured using the NIST Cryogenic Radiometer
  - The resulting profile provides use in four calibrations:
    - **Instrument Function Area:**
      - Integrated area of profile
    - **Absolute Sensitivity:**
      - Given by the absolute height of the profile
    - **Wavelength Calibration:**
      - Center location of profile provides wavelength calibration
    - **Scattered Light:**
      - The background level of the profile
  - Repeat at different wavelengths using SIRCUS to provide coverage
  - This calibration is performed for both the ESR and diodes



**Instrument-Level**