

LASER-DSR

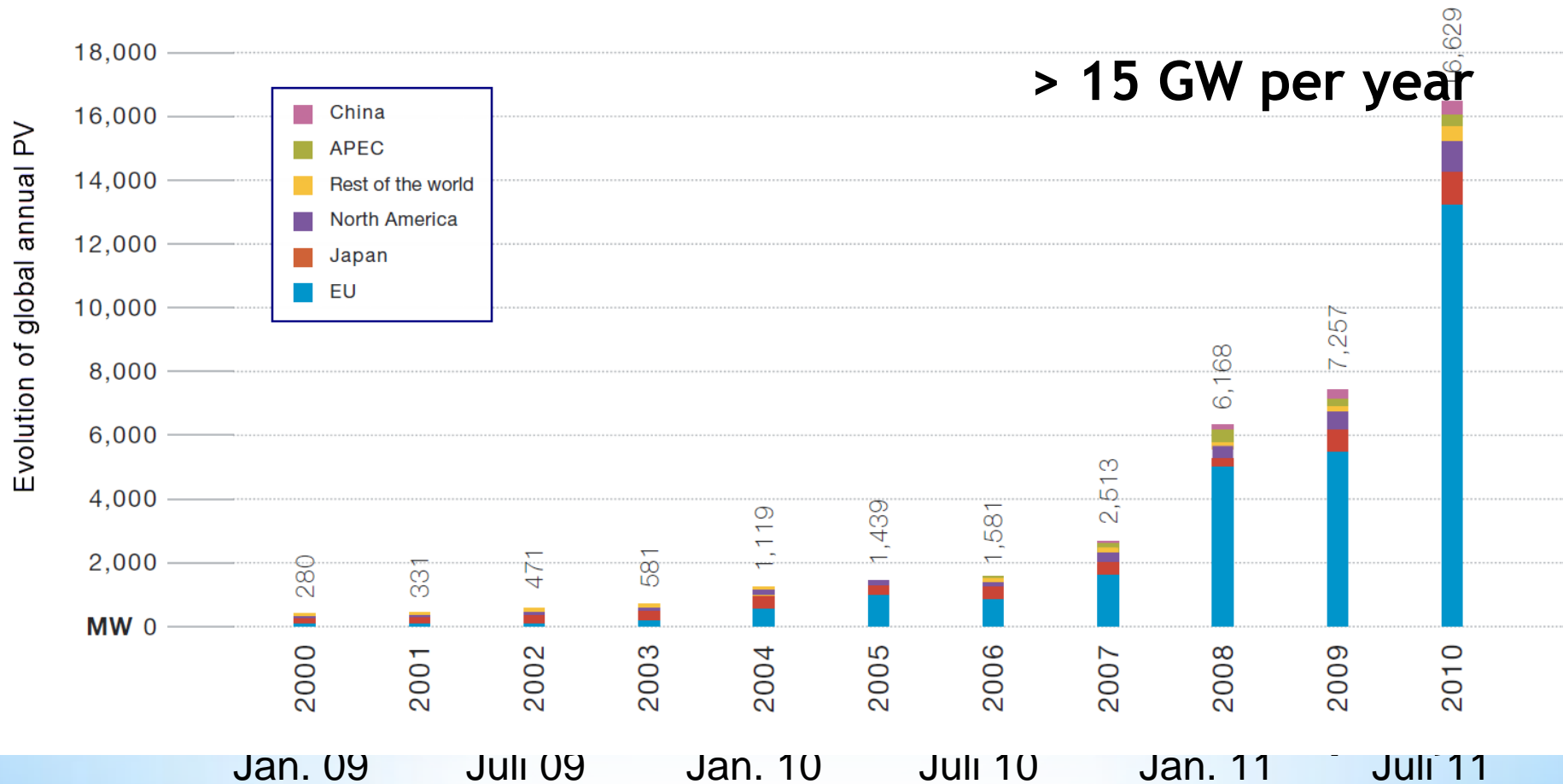
Enabling solar cell and detector calibrations with fs pulses

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Working Group 4.14 - „Solar Cells“

Overview

- Demand for high accuracy solar cell calibrations
- Solutions and their problems
- The new PTB setup: Advantages and disadvantages
- How we solve the disadvantages
- Measurements results
- Outlook

Demand for high accuracy solar cell calibrations

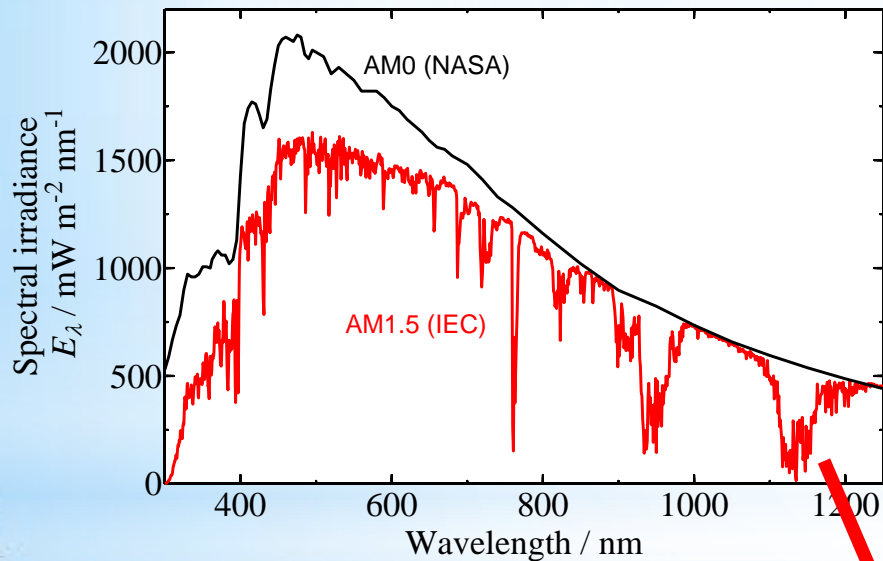


Calibration methods for solar cells

Integral		Spectral	
Outdoor, with solar radiation		Indoor	
Space ($\geq 40\text{km}$) AM0	On the earth AMn	Solar simulator	DSR-Method
Very expensive	$n \geq 1$, direct, total, global radiation	Good reproducibility, independent of place and time	
	Appropriate and stable weather conditions needed	large cells and modules	absolute method without spectral mismatch; cell size $\leq 15 \times 15 \text{ cm}^2$
	Spectral Mismatch correction where required		

Metrological background

Standard solar spectrum
on earth (AM1.5) or in space (AM0)



Spectral responsivity
of different types of solar cells

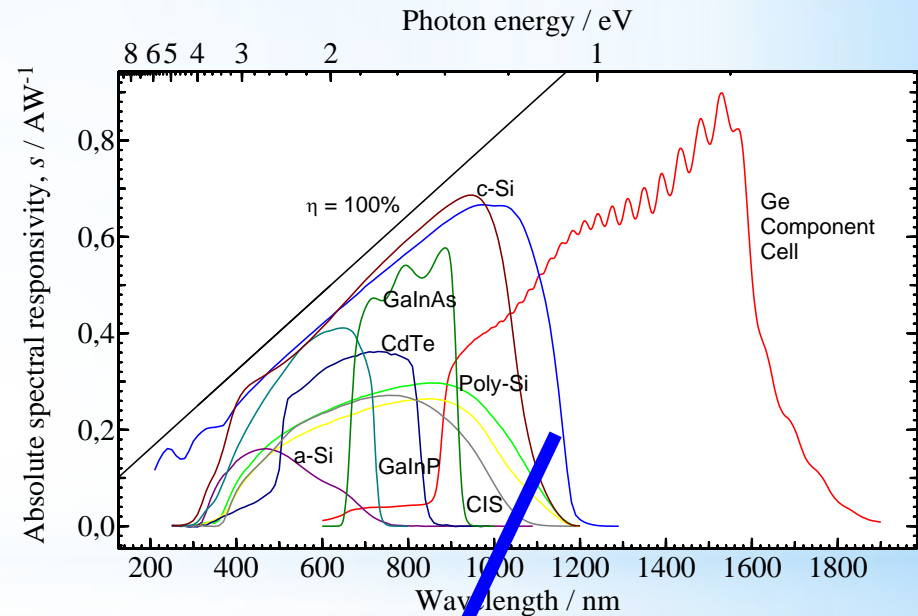
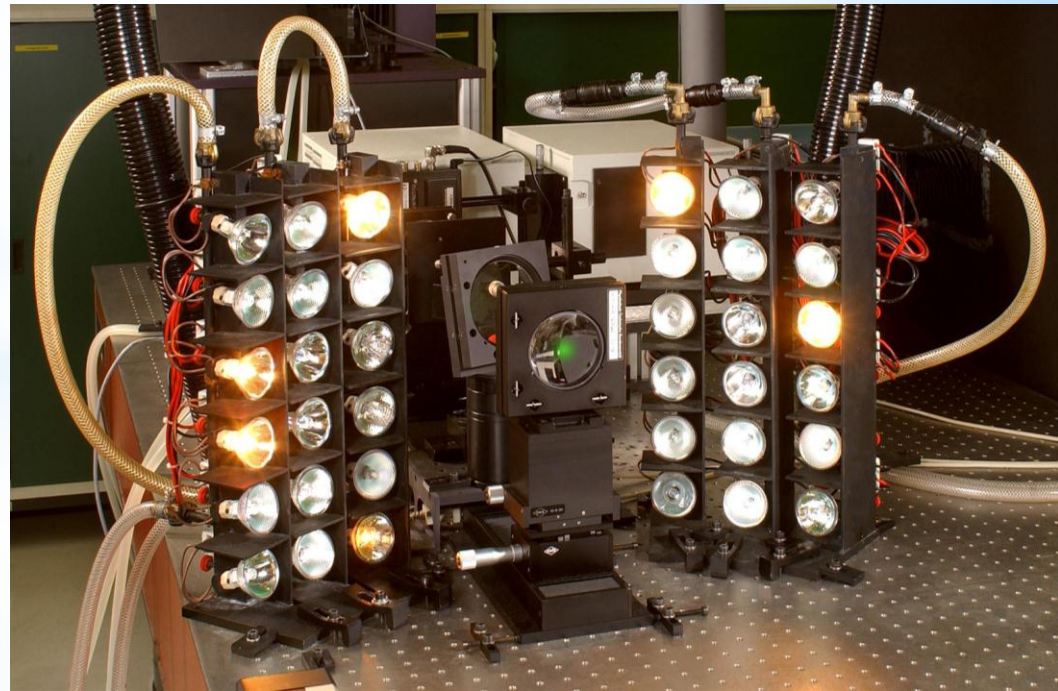


Photo current:
$$I = \int E_{\lambda, \text{Norm}}(\lambda) \cdot s(\lambda) d\lambda$$

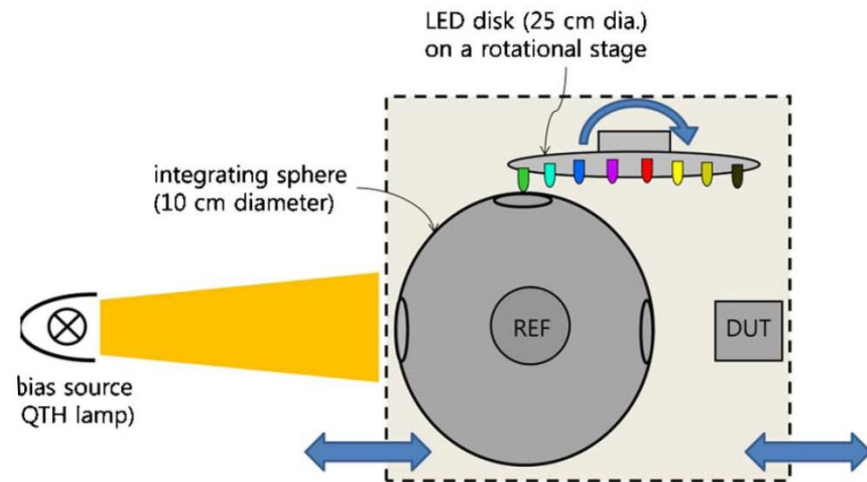
How to measure the spectral responsivity

- Xenon or quartz halogen lamp based system (DSR, SCF)
(or laser-driven xenon lamps or supercontinuum systems)
 - + Easy to use
 - Low Power (100 μ W) with subsequent problems \Rightarrow
 - Uniformity
 - Bandwidth
 - Signal-To-Noise especially at high bias levels
 - Rel. & abs. measurement
 - Size of the solar cells is limited



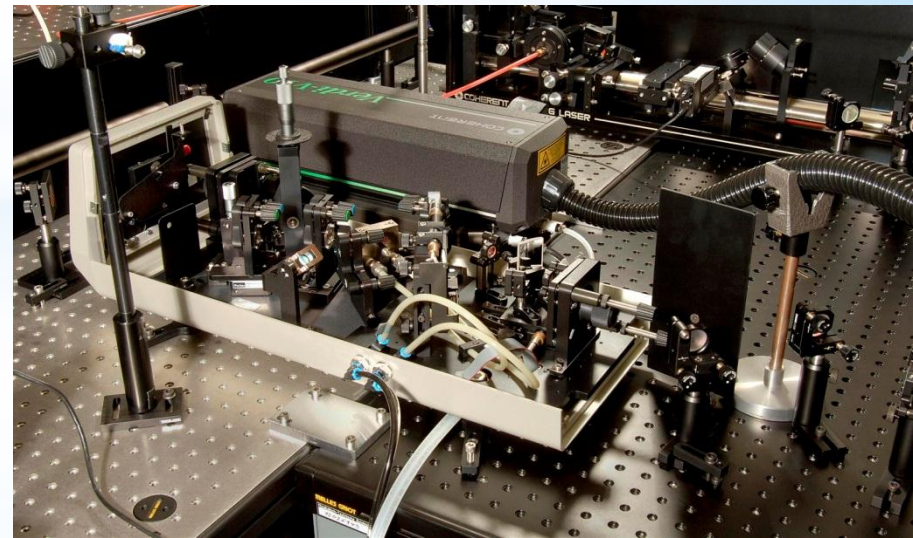
How to measure the spectral responsivity

- LED based systems (KRISS)
 - + High Power level (depending on the setup)
 - + Low priced, as no monochromator is needed
 - Wavelength stability (temperature dependent),
 - High Bandwidth
 - No wavelength tuneability



How to measure the spectral responsivity

- cw-laser based systems (TULIP, SIRCUS)
 - + Very low bandwidth
 - + High power levels
 - Automation is difficult
 - Interference effects
 - Gaps in the spectral range



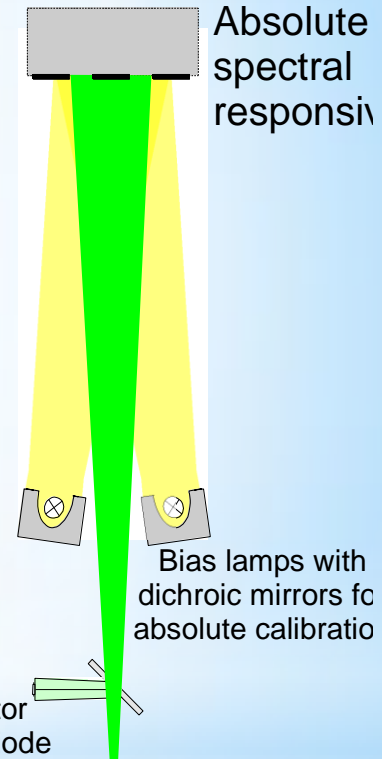
See:
Session 4a, INV 7: Armin Sperling:
“Tuneable lasers for photometry and radiometry”
Poster-Session B: DBR_PO_017, Michaela Schuster:
“Correction algorithm for interference affected measurement data”

How to measure the spectral responsivity

- Quasi cw laser based systems with modelocked Ti-Sapphire (Quasi-cw-TULIP, LASER-DSR)
 - + Very high power (up to > 3500 mW)
 - + Fully tunable
 - + Wavelength range from 190 nm to 4000 nm
 - + Complete automation possible from 210 nm to 4000 nm
 - Unknown behavior of short fs pulses to semiconductor detectors, especially in the UV and blue wavelength region
- Every 12 nm a 120 fs pulse hits the surface: Duty cycle 1/100.000
- ⇒ Integrating spheres are used for pulse stretching
 - ⇒ fluorescence effects + high power losses

The new LASER-DSR setup

- Laser substitutes lamps
 - Power increases up to a factor 1000
 - + Good uniformity
 - + Low bandwidth possible
 - + No interpolation between relative and absolute measurement



Pulsed with 80 MHz

CW signal

monitor photodiode

Modelocked
Ti:Sapphire
laser

Optical parametric
Oscillator (OPO),
SHG, THG, FHG

Pulse-to-CW
converter

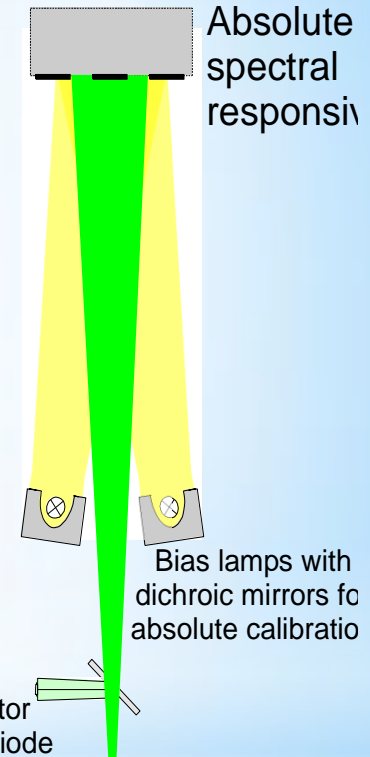
Bandpass
limitation
(monochromator)

680 nm - 1080 nm

190 nm - 4000 nm

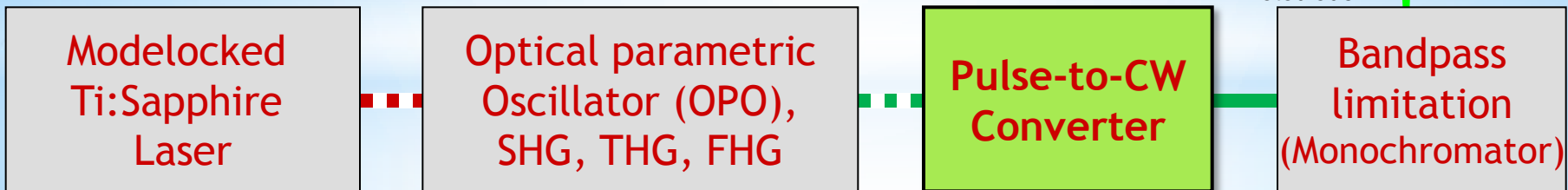
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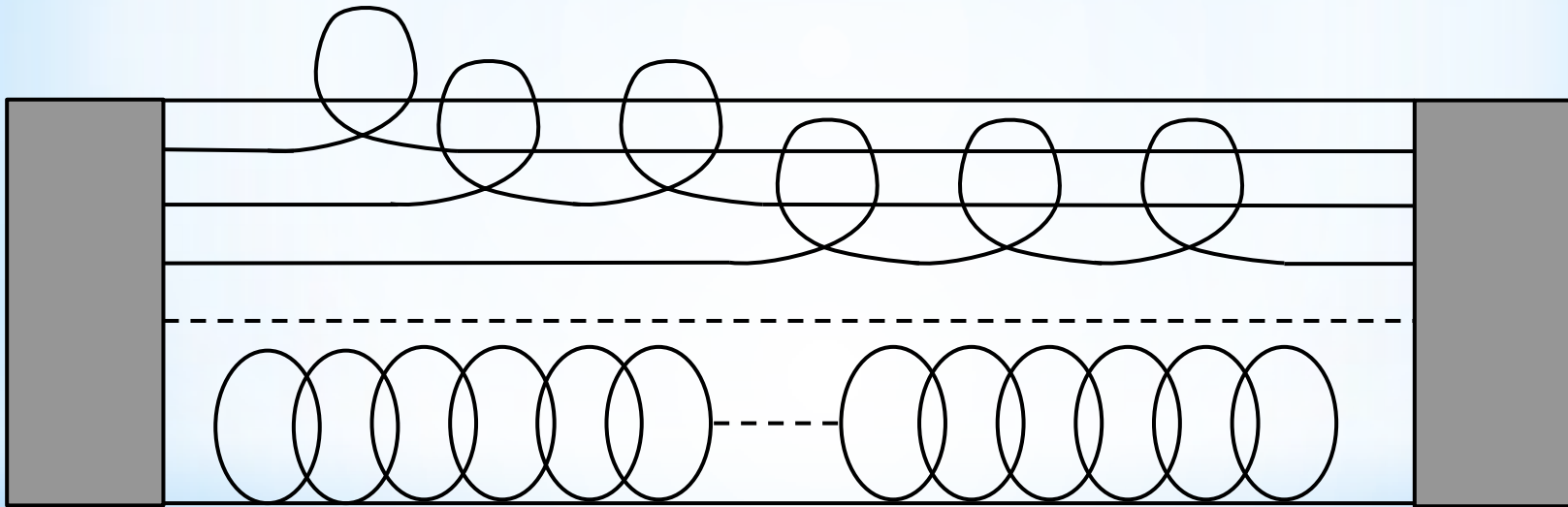
CW signal



680 nm - 1080 nm

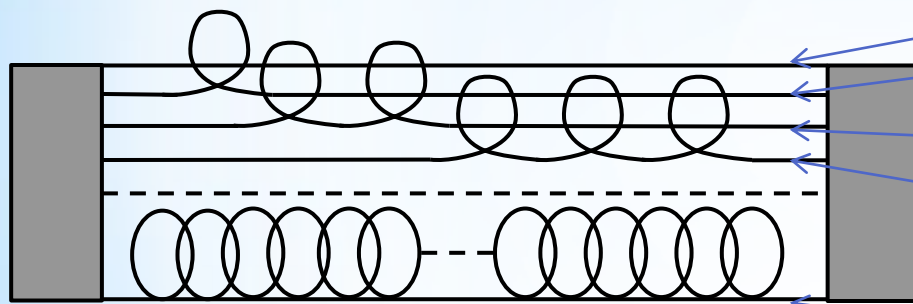
190 nm - 4000 nm

Pulse to cw converter



Pulse to cw converter

Fiber bundle with 100 multimode fibers, each with an individual length



$$L_0 = l_0$$

$$t_0 = t_0$$

$$L_1 = l_0 + 2.5 \text{ cm}$$

$$t_1 = t_0 + 0.125 \text{ ns}$$

$$L_2 = l_0 + 5.0 \text{ cm}$$

$$t_2 = t_0 + 0.250 \text{ ns}$$

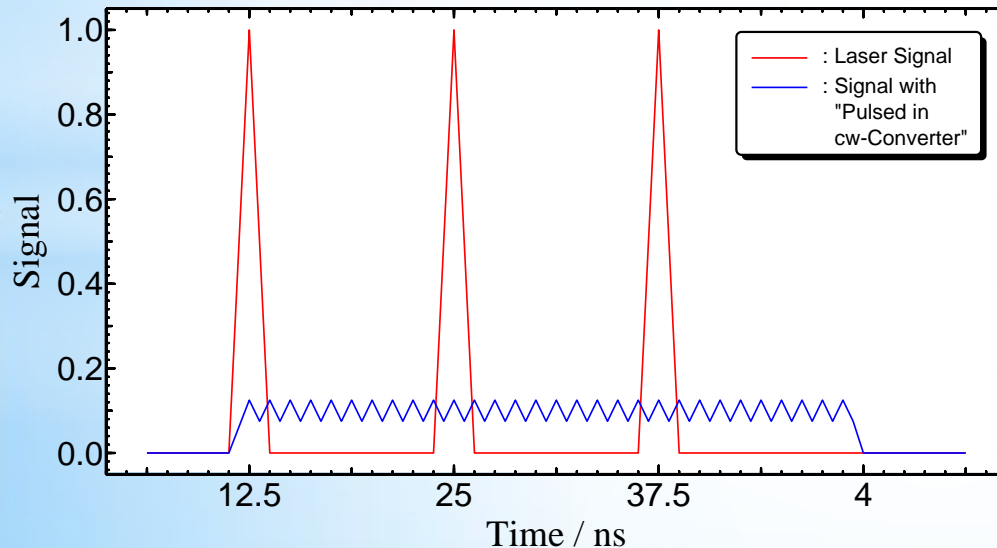
$$L_3 = l_0 + 7.5 \text{ cm}$$

$$t_3 = t_0 + 0.375 \text{ ns}$$

...

...

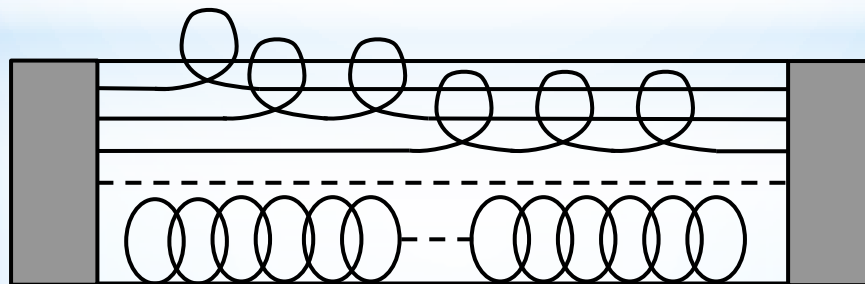
$$L_{99} = l_0 + 247.5 \text{ cm} \quad t_{99} = t_0 + 12.375 \text{ ns}$$



And after 12.5 ns the next pulse from the laser appears.

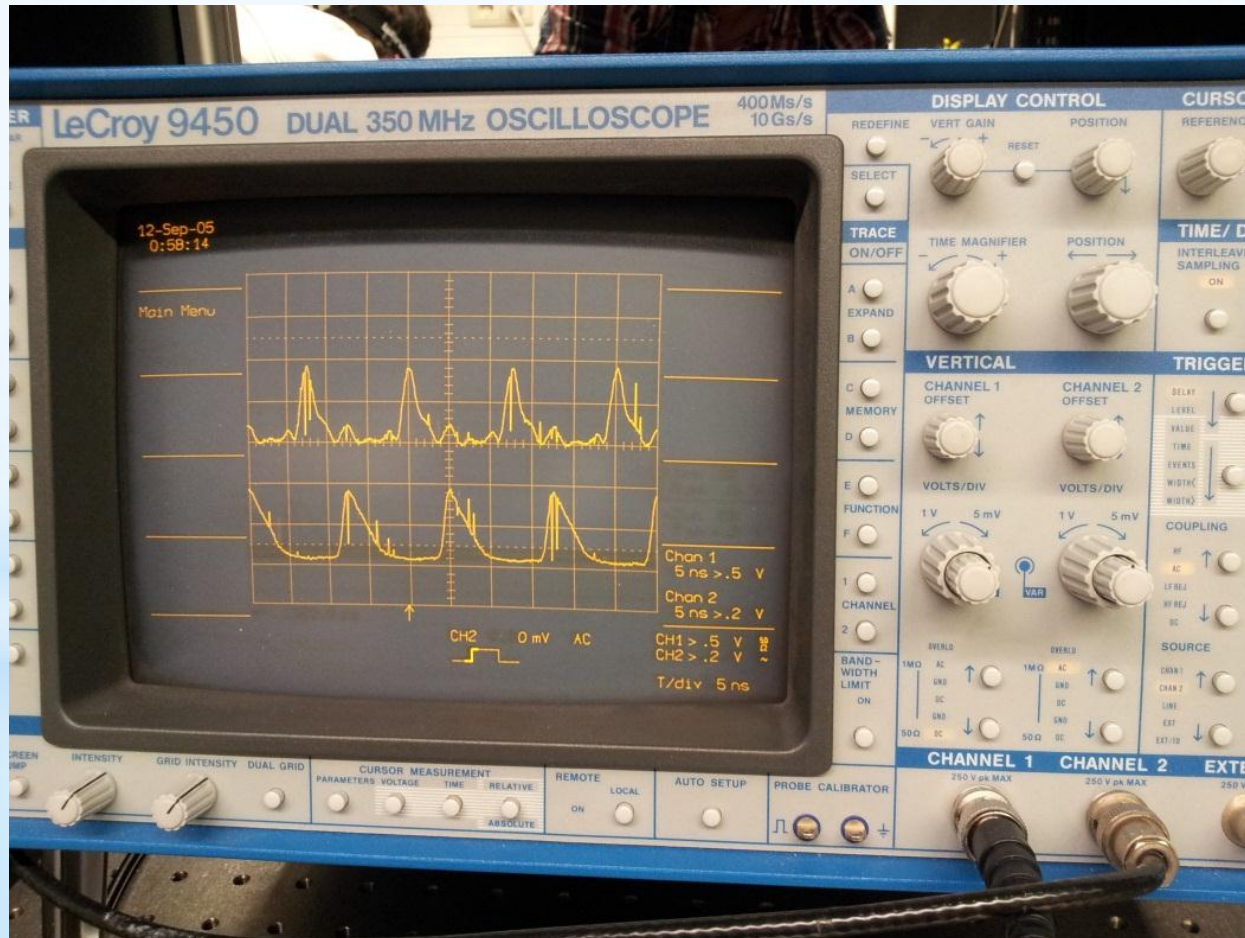
Pulse to cw converter

- Fiber bundle with different length of the single fibers
- The laser is pulsed with 80 MHz
 - ⇒ every 12.5 ns a pulse appears
 - ⇒ During this time the light moves 3 m in air and about 2 m in the fiber
 - ⇒ The distance between the shortest and longest fiber must be 2m
- We need in any case a fiber to couple into the monochromator
- To improve efficiency the ends of the fibers are fused



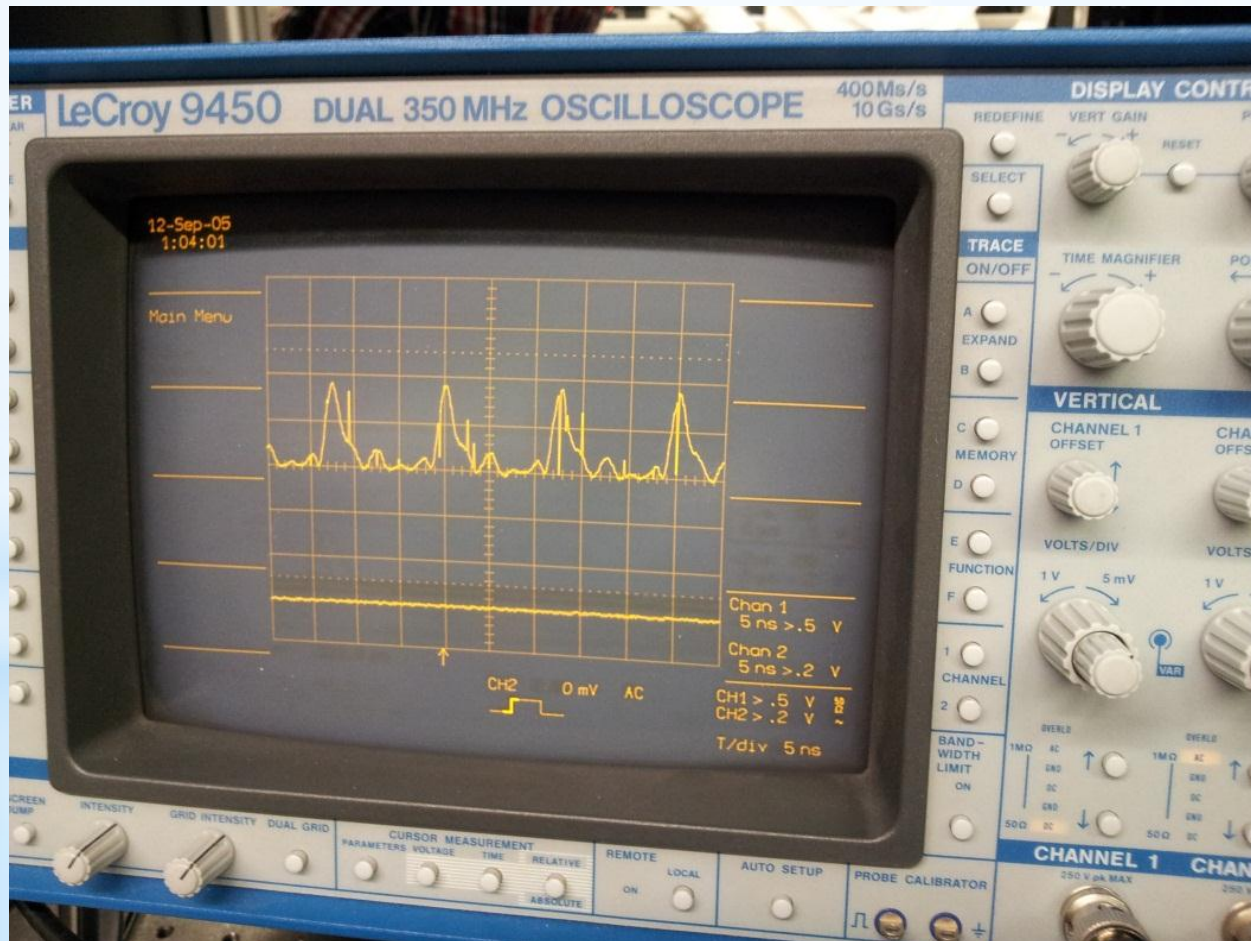
Pulse to cw converter

- Experimental prove: Signal behind a standard fiber



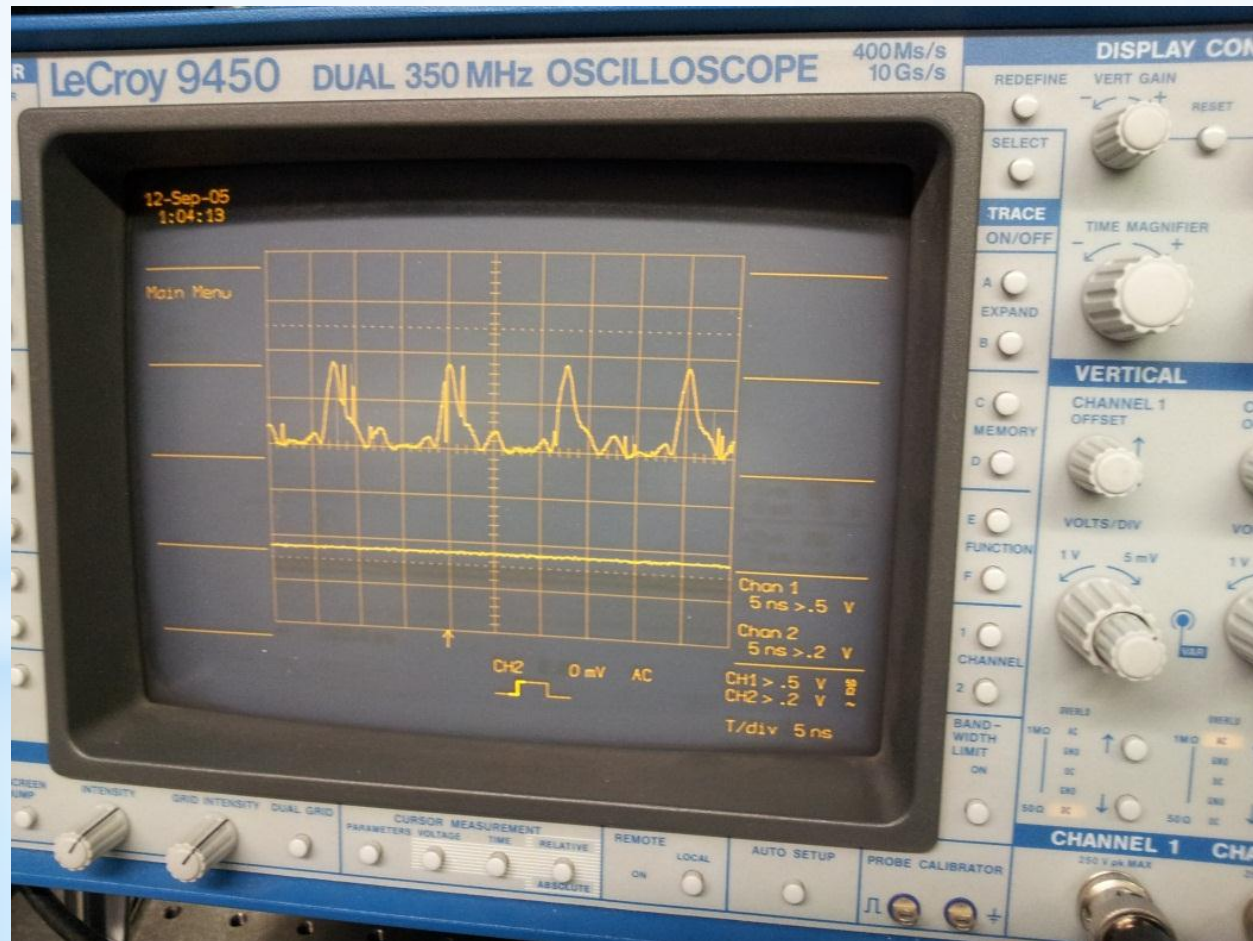
Pulse to cw converter

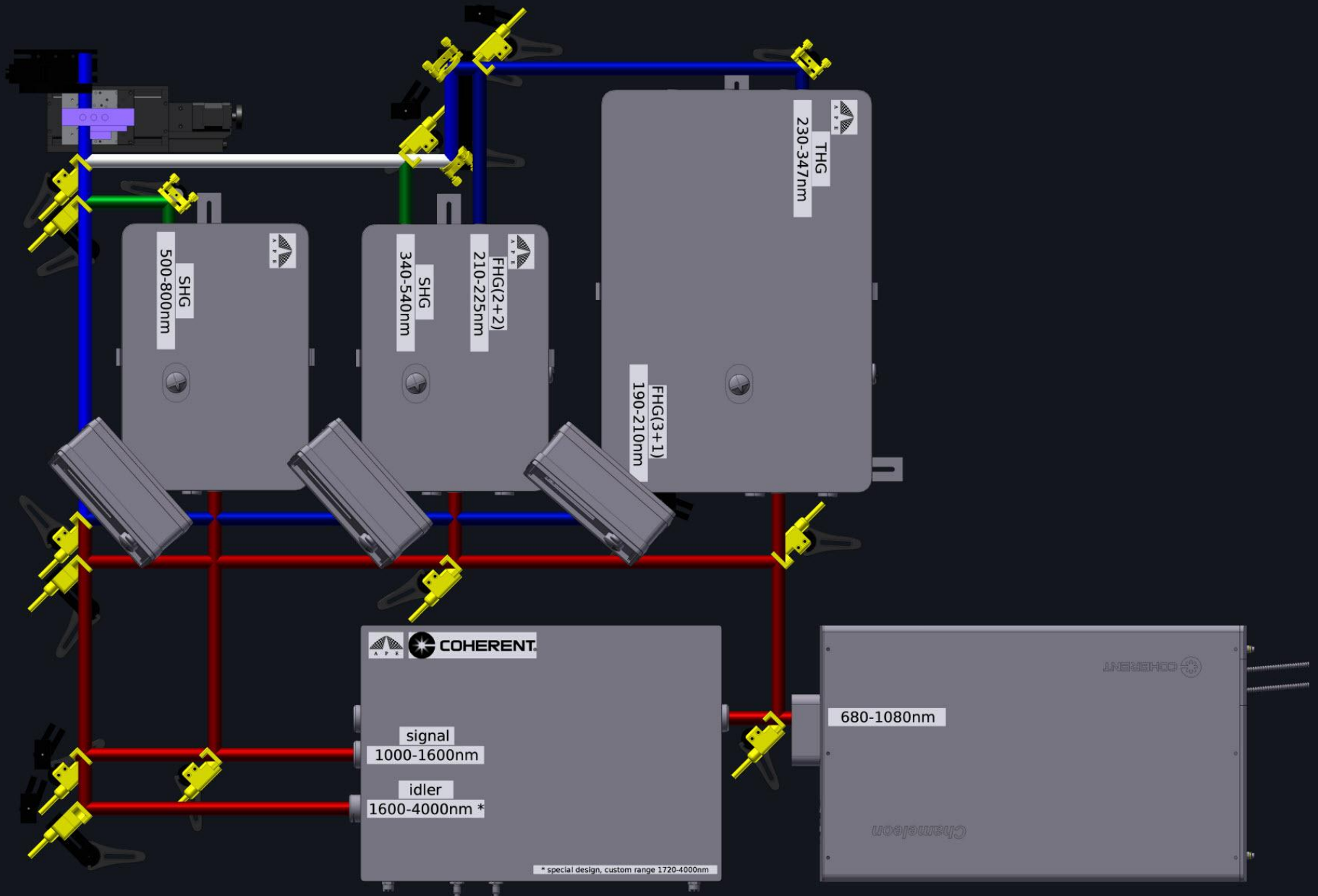
- Experimental prove: Zero line without signal

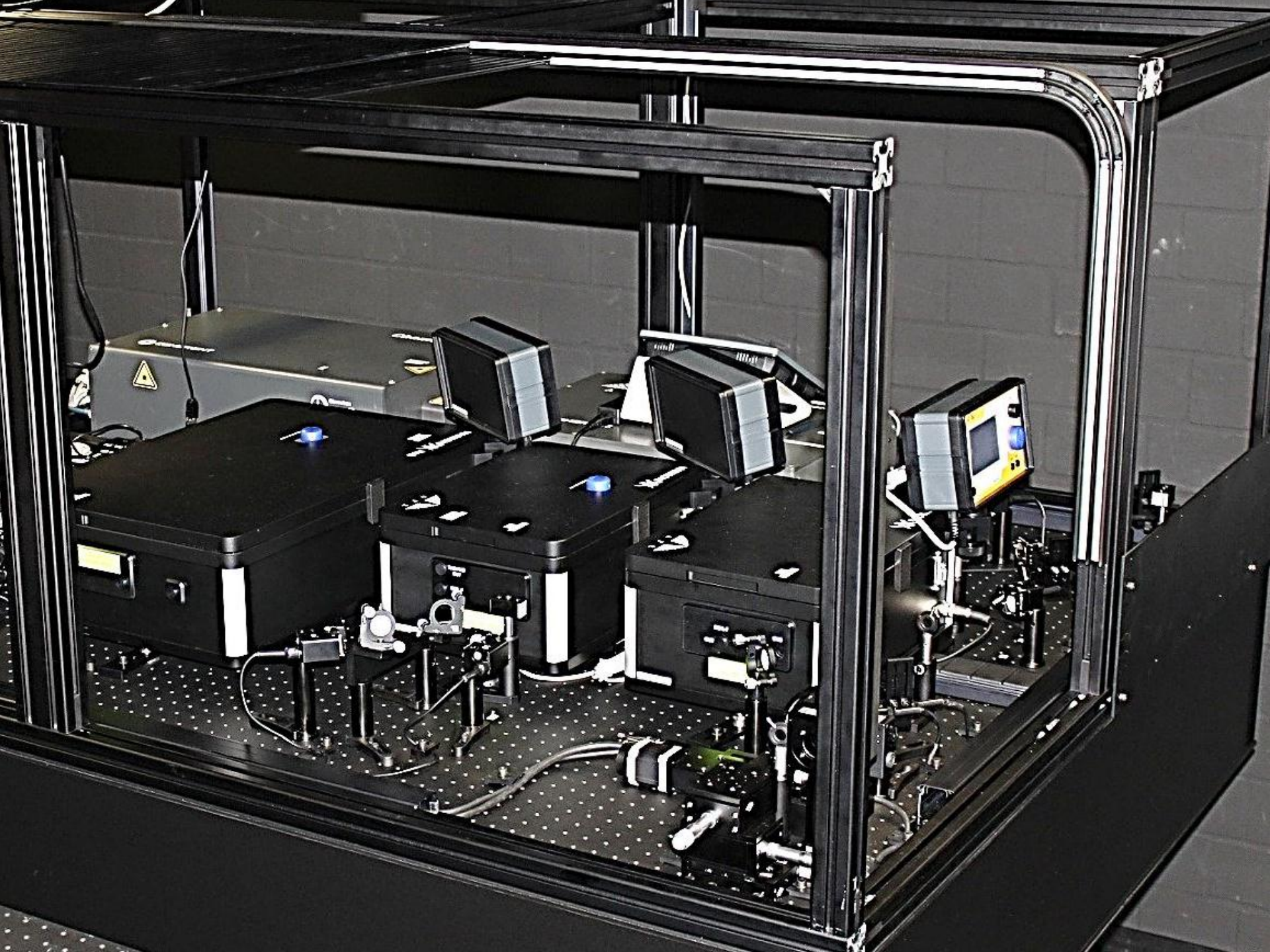


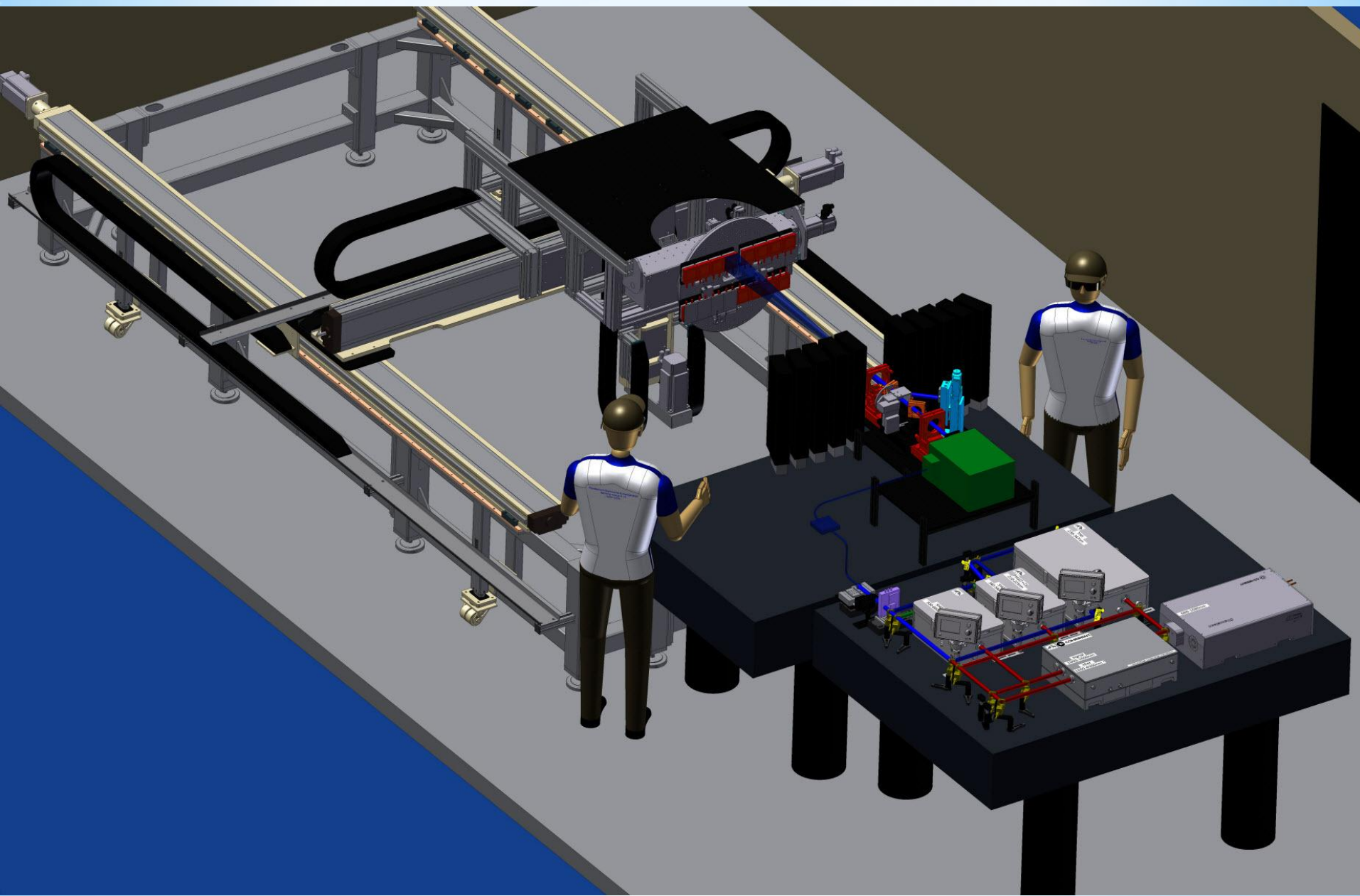
Pulse to cw converter

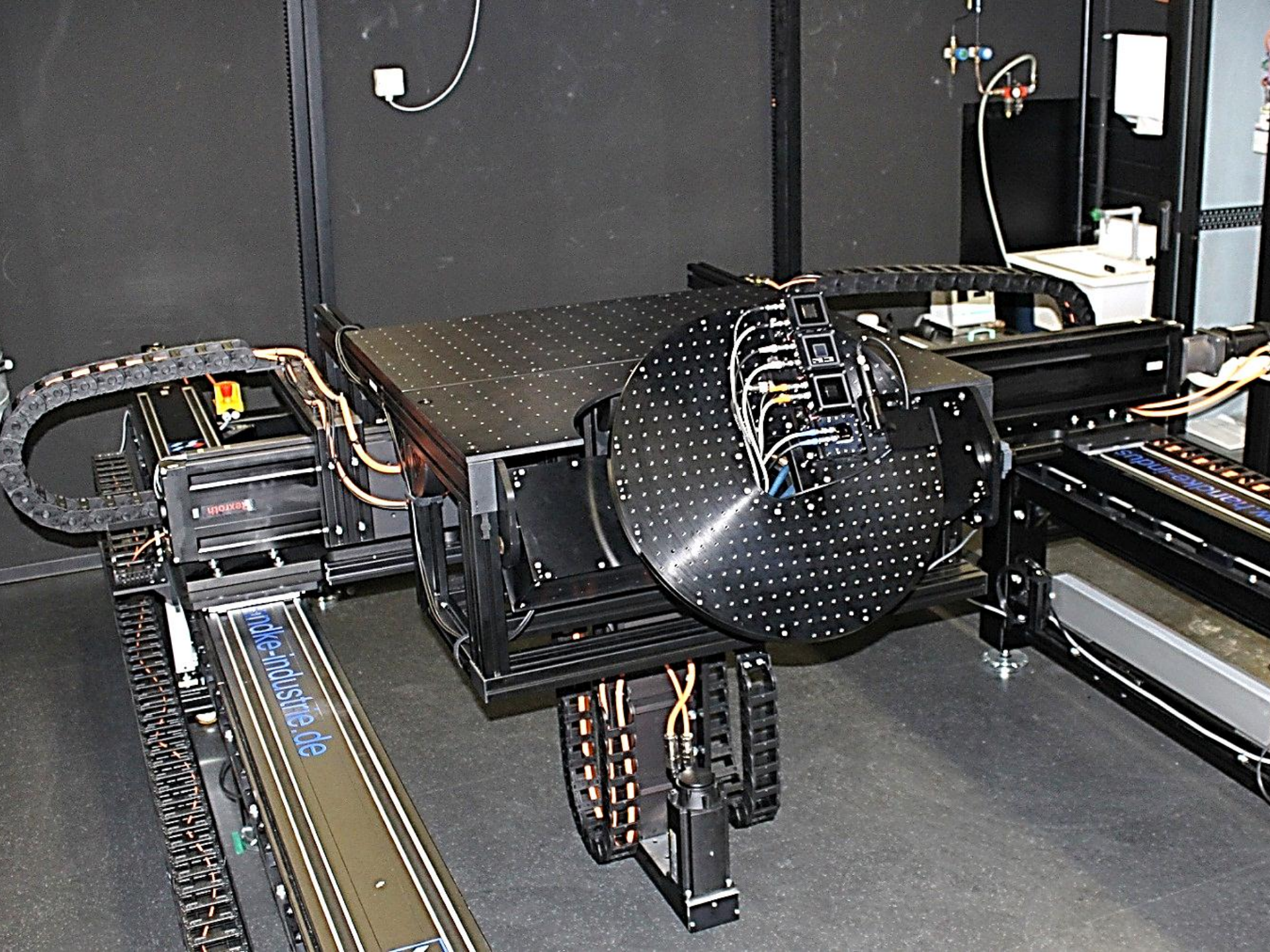
- Experimental prove: Const. Signal behind converter

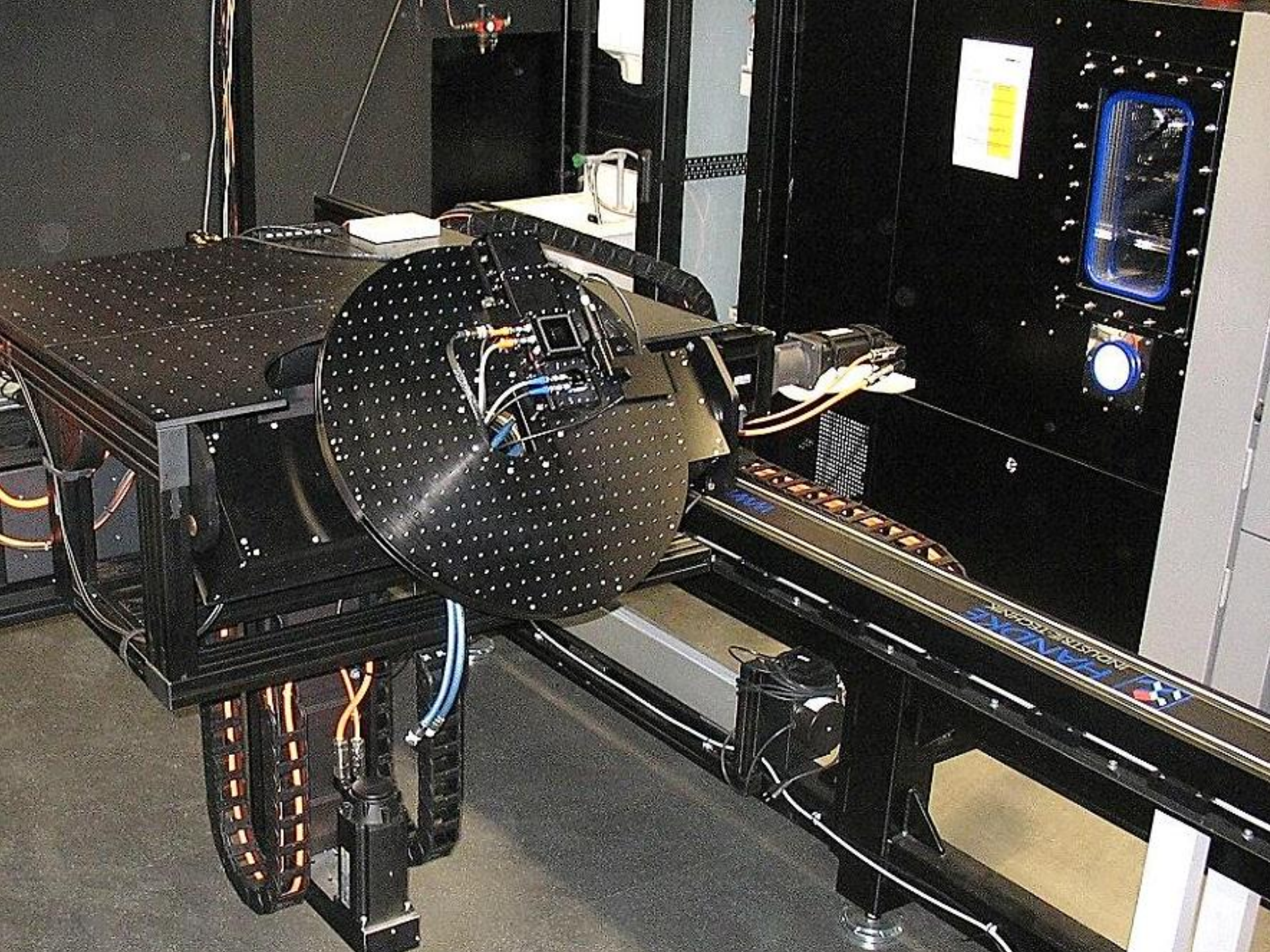




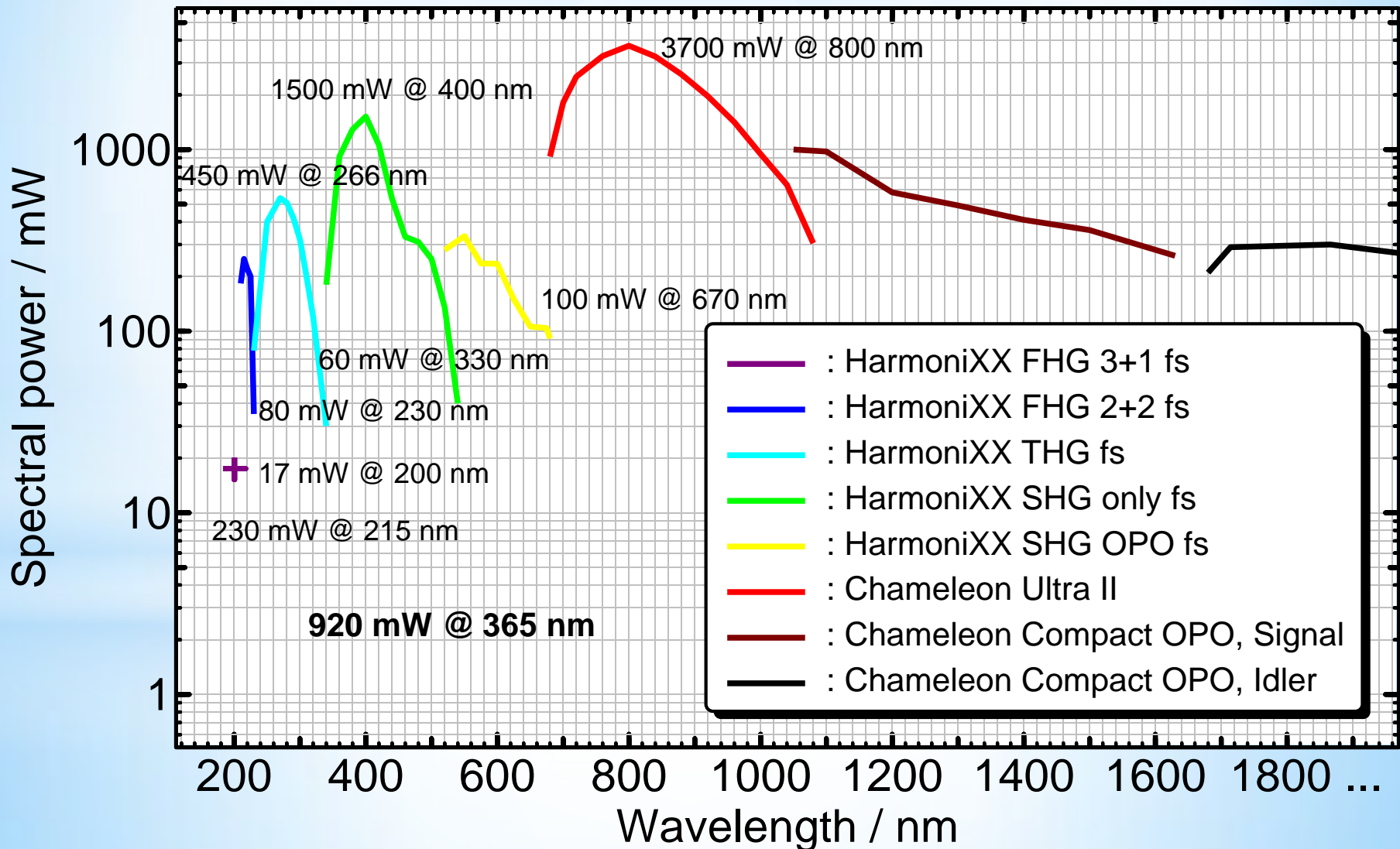








Measurement results



Conclusion

- PTB develops the next-generation of the DSR facility: LASER-DSR
- It has up to 1000 times more optical power than the old facility
- It is a multipurpose spectral comparison facility.

$$S = S(\lambda, E, f_{\text{Chopper}}, T, x, y, z, \varphi, \theta)$$

- We expect a reduction of uncertainty from 1.6% to 0.6% for large solar cells
- A fiber bundle with different lengths converts the fs pulses to a const signal
- The Pulse-To-CW converter can be used for all light sources with high repetition rates:

- ✓ Pulsed Laser
- ✓ Supercontinuum sources
- ✓ Synchrotron radiation



Photometry and Radiometry:
Biology, Medicine
Chemical Analyses